# HARMONIC MAPS AND REPRESENTATIONS OF NON-UNIFORM LATTICES OF $\operatorname{PU}(m, 1)$ 

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#### Abstract

We study representations of lattices of $\mathrm{PU}(m, 1)$ into $\mathrm{PU}(n, 1)$. We show that if a representation $\rho$ is reductive and if $m \geq 2$, then there exists a finite energy harmonic $\rho$-equivariant map from $\mathbb{H}_{\mathbb{C}}^{m}$ to $\mathbb{H}_{\mathbb{C}}^{n}$. This allows us to give a differential geometric proof of rigidity results obtained by M . Burger and A . Iozzi. We also define a new invariant associated to representations into $\mathrm{PU}(n, 1)$ of non-uniform lattices of $\mathrm{PU}(1,1)$, and more generally of fundamental groups of orientable surfaces of finite topological type and negative Euler characteristic. We prove that this invariant is bounded by a constant depending only on the Euler characteristic of the surface and we give a complete characterization of representations with maximal invariant, thus generalizing the results of D . Toledo for uniform lattices.


## 0. Introduction

Lattices in semi-simple Lie groups with no compact factor (say, defined over $\mathbb{R}$ and with trivial center) enjoy several rigidity properties. For example, with the exception of lattices in groups locally isomorphic to $\operatorname{PSL}(2, \mathbb{R})$, they all satisfy Mostow strong rigidity, which roughly means the following. Take two such Lie groups $G$ and $H$, an irreducible lattice $\Gamma$ in $G$, and a representation (that is, a homomorphism of groups) of $\Gamma$ into $H$. Assume that the representation is faithful and that the image of $\Gamma$ is also a lattice in $H$. Then the representation extends to a homomorphism from the ambient Lie group $G$ to $H$ (see (Mo73). Another type of rigidity, known as Margulis superrigidity, provides the same kind of conclusion but with much weaker assumptions: the only hypothesis is that the image of $\Gamma$ should be Zariski-dense in $H$. Superrigidity holds for lattices in Lie groups of rank at least 2 ( $($ Ma91]) and for lattices of quaternionic or octonionic hyperbolic spaces (that is, lattices in the rank one Lie groups $\operatorname{Sp}(m, 1), m \geq 2$, and $\mathrm{F}_{4}^{-20}$ ) (see [Co92] and [GS92]). On the contrary, for lattices of real and complex hyperbolic spaces, namely, lattices in the other rank one Lie groups $\mathrm{PO}(m, 1)$ and $\mathrm{PU}(m, 1)$, superrigidity is known to fail in general.

In this paper, we will focus on lattices in $\mathrm{PU}(m, 1)$, the group of orientation-preserving isometries (or equivalently, of biholomorphisms) of complex hyperbolic $m$-space $\mathbb{H}_{\mathbb{C}}^{m}=$ $\mathrm{PU}(m, 1) / \mathrm{U}(m)$. They are of particular interest because they lie somewhere in between the very flexible lattices of $\mathrm{PO}(m, 1)$ and those, superrigid, of the higher rank Lie groups.

In GM87, W. M. Goldman and J. J. Millson studied representation spaces of uniform torsion-free lattices $\Gamma<\mathrm{SU}(m, 1)$ (which can be considered as lattices in $\mathrm{PU}(m, 1)$ ) into $\operatorname{PU}(n, 1)$, for $n>m \geq 2$. They proved that there are no non-trivial deformations of the standard representation of such a lattice. This means that all nearby representations are $\mathbb{C}$-Fuchsian, namely, they are discrete, faithful, and they stabilize a totally geodesic copy of $\mathbb{H}_{\mathbb{C}}^{m}$ in $\mathbb{H}_{\mathbb{C}}^{n}$. The case $m=1$ was previously treated by Goldman in Go85. Note that the corresponding statement for lattices in $\mathrm{PO}(m, 1)$ is false (cf. for example [JM87]).

They also conjectured that a much stronger rigidity should hold. The volume of a representation $\rho$ of a torsion-free uniform lattice $\Gamma<\mathrm{PU}(m, 1)$ into $\mathrm{PU}(n, 1)$ is defined by pulling-back

[^0]the Kähler form of $\mathbb{H}_{\mathbb{C}}^{n}$ on $\mathbb{H}_{\mathbb{C}}^{m}$ via the representation, taking its $m$-th exterior power to obtain a de Rham cohomology class in $H_{D R}^{2 m}\left(\Gamma \backslash \mathbb{H}_{\mathbb{C}}^{m}\right)$ and evaluating it on the fundamental class of the compact quotient $\Gamma \backslash \mathbb{H}_{\mathbb{C}}^{m}$. Observe that if $\Gamma<\mathrm{SU}(m, 1)$ and if $\rho: \Gamma \longrightarrow \mathrm{PU}(n, 1), n>m$, is the standard representation, then $\operatorname{vol}(\rho)=\operatorname{Vol}\left(\Gamma \backslash \mathbb{H}_{\mathbb{C}}^{m}\right)$. Their conjecture then reads: any representation $\rho$ such that $\operatorname{vol}(\rho)=\operatorname{Vol}\left(\Gamma \backslash \mathbb{H}_{\mathbb{C}}^{m}\right)$ must be $\mathbb{C}$-Fuchsian. This was proved by K. Corlette in Co88 for $m \geq 2$ and by D. Toledo in To89 for $m=1$. Remark that the volume assumption is needed precisely because lattices in $\mathrm{PU}(m, 1)$ are not superrigid.

Recently, M. Burger and A. Iozzi proved in [BI01] (see also [Io02]) that the conjecture is also true for non-uniform lattices of $\mathrm{PU}(m, 1), m \geq 2$, if one suitably modifies the definition of the "volume" of the representation (indeed, with the former one, any representation of a non-uniform lattice has zero volume). We will explain precisely how this invariant is computed in section 3.1 but here we sketch its definition. Again, the Kähler form $\omega_{n}$ of $\mathbb{H}_{\mathbb{C}}^{n}$ is pulledback to the quotient $\Gamma \backslash \mathbb{H}_{\mathbb{C}}^{m}$ via the representation. It turns out that this gives a well-defined $L^{2}$-cohomology class in $H_{(2)}^{2}\left(\Gamma \backslash \mathbb{H}_{\mathbb{C}}^{m}\right)$. Now, integrating a $L^{2}$-representative $\rho^{\star} \omega_{n}$ against the Kähler form $\omega_{m}$ of $\Gamma \backslash \mathbb{H}_{\mathbb{C}}^{m}$, we get the Burger-Iozzi invariant (slightly modified from [BI01]):

$$
\tau(\rho):=\frac{1}{2 m} \int_{\Gamma \backslash \mathbb{H}_{\mathbb{C}}^{m}}\left\langle\rho^{\star} \omega_{n}, \omega_{m}\right\rangle d V_{m}
$$

In complex dimension 1 and for uniform lattices, this invariant coincides with the invariant defined in [To89]. We can now state the main theorem of [BI01]:

Theorem A Let $\Gamma$ be a torsion-free lattice in $\mathrm{PU}(m, 1), m \geq 2$, and let $\rho: \Gamma \longrightarrow \mathrm{PU}(n, 1)$ be a representation. Then $|\tau(\rho)| \leq \operatorname{Vol}\left(\Gamma \backslash \mathbb{H}_{\mathbb{C}}^{m}\right)$ and equality holds if and only if there exists a totally geodesic isometric $\rho$-equivariant embedding of $\mathbb{H}_{\mathbb{C}}^{m}$ into $\mathbb{H}_{\mathbb{C}}^{n}$ (in particular, $\rho(\Gamma)$ seen as a subgroup of $\mathrm{PU}(m, 1)$ is a lattice $)$.

Burger and Iozzi's proof heavily relies on the theory of bounded cohomology developed by Burger and N. Monod in BM02]. As a corollary, they obtain the result of Goldman and Millson for a general lattice:

Corollary $\mathbf{A}^{\prime}$ Let $\Gamma$ be a torsion-free lattice in $\mathrm{SU}(m, 1), m \geq 2$, and let $n>m$. Then there are no non-trivial deformations of the standard representation of $\Gamma$ into $\mathrm{PU}(n, 1)$.

The aim of this paper is to use harmonic map techniques to give a new and more (differential) geometric proof of Theorem A and to extend this result to the case of complex dimension 1, that is, of non-uniform lattices of $\mathrm{PU}(1,1)$.

The over-all harmonic map strategy for proving rigidity results about representations of lattices in a Lie Group $G$ to another Lie group $H$ goes as follows. First, one has to know that there exists a harmonic map between the corresponding symmetric spaces, equivariant w.r.t. the representation. Then, one must prove, generally by using a Bochner-type formula, that there are additional constraints on the harmonic map, which force it to be pluriharmonic, holomorphic, totally geodesic, or isometric, depending on the situation.

For a uniform $\Gamma$ and when the target symmetric space is non-positively curved (which will be assumed from now on), the existence results for harmonic maps go back to J. Eells and J. H. Sampson in ES64 and have been extended by several authors, in particular by Corlette in Co88. The second step was pionneered by Y.-T. Siu in Si80 where he proved a strenghtened version of Mostow strong rigidity theorem in the case of Hermitian locally symmetric spaces. This has later on been applied in different directions by many authors. We should mention the proof of the above conjecture of Goldman and Millson by Corlette in [Co88]
and the geometric proof of Margulis superrigidity theorem in the Archimedean setting worked out by N. Mok, Y.-T. Siu and S.-K. Yeung in MSY93].

When the lattice is not uniform, the only general existence theorem for harmonic maps is due to Corlette in Co92], see Theorem 1.1 below. The main issue is that to apply this theorem, one needs to prove that there exists an equivariant map of finite energy (see section 1 for the definition). If this is the case, the harmonic map also has finite energy and the second step generally goes as if the lattice was uniform, but is technically more involved. The energy finiteness condition is very important, and in general difficult to prove. In some particular cases it is possible to obtain a harmonic map by other means (see for example [JZ97] and section 4 of this paper) but then its energy is infinite and the analysis that follows becomes much harder. These are the reasons why, for example, "geometric superrigidity" for nonuniform lattices is not yet proved.

Our paper is organized as follows. The first three sections are devoted to the proof of Theorem A. In section 1 we give the necessary definitions and we prove that Corlette's general theorem applies in our setting, so that we obtain our main existence theorem (cf. Theorem 1.2):

Theorem B Let $\Gamma$ be a torsion-free lattice in $\mathrm{PU}(m, 1), m \geq 2$, and $\rho: \Gamma \longrightarrow \mathrm{PU}(n, 1)$ be a representation such that $\rho(\Gamma)$ has no fixed point on the boundary at infinity of $\mathbb{H}_{\mathbb{C}}^{n}$. Then there exists a finite energy harmonic $\rho$-equivariant map from $\mathbb{H}_{\mathbb{C}}^{m}$ to $\mathbb{H}_{\mathbb{C}}^{n}$.

In section 2 we prove that the harmonic map previously obtained is pluriharmonic and even holomorphic or antiholomorphic if its rank is high enough (at least 3 at some point). Section 3 is devoted to the precise definition of the Burger-Iozzi invariant and to the proof of Theorem A.

In section 4, we study the case of lattices of $\mathrm{PU}(1,1)$, that is, of fundamental groups of Riemann surfaces with a finite volume hyperbolic metric. The analogue of Theorem A for uniform lattices was proved by Toledo in [To89]. In [GP03] (see also [GP00]), N. Gusevskii and J. R. Parker claim that if one restricts to type-preserving representations, then the original definition of the Toledo invariant can be used to generalize Toledo's result to non-uniform lattices. However, it seems to us that this claim is not entirely exact (see for example the remark following Proposition 4.5).

There are mainly two reasons why the 1-dimensional case is different from the higher dimensional one. First of all, Toledo and/or Burger-Iozzi invariants are not defined for non-uniform lattices. Secondly, there are representations for which no equivariant map of finite energy exists. It should also be noted that Corollary A' fails in this case by a result of Gusevskii and Parker (cf. GP00).

As we shall see, it is in fact more natural to work in the general setting of fundamental groups of orientable surfaces of finite topological type, that is surfaces obtained by removing finitely many points from closed orientable surfaces. Using cohomology with compact support, we define at the beginning of section 4 a new invariant associated to representations of these fundamental groups into $\operatorname{PU}(n, 1)$, which we again call $\tau$. We obtain (see Theorem 4.3):

Theorem C Let $\Gamma$ be the fundamental group of a p-times punctured closed orientable surface $M$ of negative Euler characterisic $\chi(M)$, and let $\rho: \Gamma \longrightarrow \mathrm{PU}(n, 1)$ be a representation. Then $|\tau(\rho)| \leq-2 \pi \chi(M)$ and equality holds if and only if $\rho(\Gamma)$ stabilizes a complex geodesic $L$ in $\mathbb{H}_{\mathbb{C}}^{n}, \rho$ is faithful and discrete, and $M$ is diffeomorphic to the quotient $\rho(\Gamma) \backslash L$.

The proof relies on the fact that though there may be no equivariant map of finite energy, there exists an equivariant harmonic map whose energy density can be controlled. This control allows us to extend the proofs given in the finite energy case to this setting.
Remark. In an earlier version of this paper, Theorem C was proven in a weaker form, and only for what we call tame representations (see Definition 4.2). M. Burger and A. Iozzi then informed us that their methods should allow them to get rid of this tameness assumption. Later, they communicated us the text [BI03], where they define a "bounded Toledo number" and prove Theorem C.

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## 1. EXistence of finite energy equivariant harmonic maps

In this section, we assume that $m \geq 2$.
Let $\Gamma$ be a torsion-free lattice in $\mathrm{PU}(m, 1)$, the group of biholomorphisms of complex hyperbolic $m$-space $\mathbb{H}_{\mathbb{C}}^{m}$ and let $\rho: \Gamma \longrightarrow \mathrm{PU}(n, 1)$ be a representation into the group of biholomorphisms of complex hyperbolic $n$-space $\mathbb{H}_{\mathbb{C}}^{n}$.

We call $M$ the quotient manifold $\Gamma \backslash \mathbb{H}_{\mathbb{C}}^{m}$. The representation $\rho$ determines a flat bundle $M \times{ }_{\rho} \mathbb{H}_{\mathbb{C}}^{n}$ over $M$ with fibers isomorphic to $\mathbb{H}_{\mathbb{C}}^{n}$. Since $\mathbb{H}_{\mathbb{C}}^{n}$ is contractible, this bundle has global sections. This is equivalent to the existence of maps (belonging to the same homotopy class) from $\mathbb{H}_{\mathbb{C}}^{m}$ to $\mathbb{H}_{\mathbb{C}}^{n}$, equivariant w.r.t. the representation $\rho$. Let $f$ be such a map (or section).

We can consider the differential $\mathrm{d} f$ of $f$ as a $f^{\star} T \mathbb{H}_{\mathbb{C}}^{n}$-valued 1-form on $\mathbb{H}_{\mathbb{C}}^{m}$. There is a natural pointwise scalar product on such forms coming from the Riemannian metrics $g_{m}$ and $g_{n}$ (of constant holomorphic sectional curvature -1) on $\mathbb{H}_{\mathbb{C}}^{m}$ and $\mathbb{H}_{\mathbb{C}}^{m}$ : if $\left(e_{i}\right)_{1 \leq i \leq 2 m}$ is a $g_{m^{-}}$ orthonormal basis of $T_{x} \mathbb{H}_{\mathbb{C}}^{m}$, then $\|\mathrm{d} f\|_{x}^{2}:=\sum_{i} g_{n}\left(\mathrm{~d} f\left(e_{i}\right), \mathrm{d} f\left(e_{i}\right)\right)$. Since $f$ is $\rho$-equivariant and the action of $\Gamma$ on $\mathbb{H}_{\mathbb{C}}^{n}$ via $\rho$ is isometric, $\|\mathrm{d} f\|$ is a well-defined function on $M$. We say that $f$ has finite energy if the energy density $e(f):=\frac{1}{2}\|\mathrm{~d} f\|^{2}$ of $f$ is integrable on $M$ :

$$
E(f)=\frac{1}{2} \int_{M}\|\mathrm{~d} f\|^{2} d V_{m}<+\infty
$$

where $d V_{m}$ is the volume density of the metric $g_{m}$. When there is no risk of confusion, we will write $e$ instead of $e(f)$ for the energy density of $f$.

There is also a natural connection $\nabla$ on $f^{\star} T \mathbb{H}_{\mathbb{C}}^{n}$-valued 1-forms on $\mathbb{H}_{\mathbb{C}}^{m}$ coming from the Levi-Civita connections $\nabla^{m}$ and $\nabla^{n}$ of $\mathbb{H}_{\mathbb{C}}^{m}$ and $\mathbb{H}_{\mathbb{C}}^{n}$. If $\nabla^{f^{\star} T \mathbb{H}_{\mathbb{C}}^{n}}$ denotes the connection induced by $\nabla^{n}$ on the bundle $f^{\star} T \mathbb{H}_{\mathbb{C}}^{n} \longrightarrow \mathbb{H}_{\mathbb{C}}^{m}$, then $\nabla \mathrm{d} f(X, Y)=\nabla_{X}^{f^{\star} T \mathbb{H}_{\mathbb{C}}^{n}} \mathrm{~d} f(Y)-\mathrm{d} f\left(\nabla_{X}^{m} Y\right)$. Since $\nabla^{m}$ and $\nabla^{n}$ are torsion-free, $\nabla \mathrm{d} f$ is a symmetric 2 -tensor taking values in $f^{\star} T \mathbb{H}_{\mathbb{C}}^{n}$.

A map $f: \mathbb{H}_{\mathbb{C}}^{m} \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$ is said to be harmonic if $\operatorname{tr}_{g_{m}} \nabla \mathrm{~d} f=0$.
The following theorem of Corlette ( Co92]) implies that if there exists a finite energy $\rho$ equivariant map from the universal cover $\mathbb{H}_{\mathbb{C}}^{m}$ of $M$ to $\mathbb{H}_{\mathbb{C}}^{n}$, and under a very mild assumption on $\rho$, then there exists a harmonic $\rho$-equivariant map of finite energy from $\mathbb{H}_{\mathbb{C}}^{m}$ to $\mathbb{H}_{\mathbb{C}}^{n}$ :

Theorem 1.1. Let $X$ be a complete Riemannian manifold and $Y$ a complete simply-connected manifold with non-positive sectional curvature. Let $\rho: \pi_{1}(X) \longrightarrow \operatorname{Isom}(Y)$ be a representation such that the induced action of $\pi_{1}(X)$ on the sphere at infinity of $Y$ has no fixed point ( $\rho$ is then called reductive). If there exists a $\rho$-equivariant map of finite energy from the universal
cover $\widetilde{X}$ of $X$ to $Y$, then there exists a harmonic $\rho$-equivariant map of finite energy from $\widetilde{X}$ to $Y$.

Theorem B will therefore follow from the
Theorem 1.2. Let $\Gamma$ be a torsion-free lattice in $\operatorname{PU}(m, 1), m \geq 2$, and let $\rho$ be a representation of $\Gamma$ into $\mathrm{PU}(n, 1)$. Then there exists a finite energy $\rho$-equivariant map $\mathbb{H}_{\mathbb{C}}^{m} \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$.

Proof. Of course this is trivially true if the manifold is compact, that is, if $\Gamma$ is a uniform lattice. To prove the theorem in the non-uniform case, we recall some known facts about the structure at infinity of the finite volume complex hyperbolic manifold $M=\Gamma \backslash \mathbb{H}_{\mathbb{C}}^{m}$, cf. for example [Go99], or [Bi97] and HS96.

We will work with the Siegel model of complex hyperbolic space:

$$
\mathbb{H}_{\mathbb{C}}^{m}=\left\{(z, w) \in \mathbb{C}^{m-1} \times \mathbb{C} \mid 2 \operatorname{Re}(w)>\langle\langle z, z\rangle\rangle\right\},
$$

where $\langle\langle.,\rangle$.$\rangle is the standard Hermitian product on \mathbb{C}^{m-1}$. We will call $h$ the function given by $h(z, w)=2 \operatorname{Re}(w)-\langle\langle z, z\rangle\rangle$. The boundary at infinity of $\mathbb{H}_{\mathbb{C}}^{m}$ is the set $\{h=0\} \cup\{\infty\}$ and the horospheres in $\mathbb{H}_{\mathbb{C}}^{m}$ centered at $\infty$ are the level sets of $h$. The complex hyperbolic metric (of constant holomorphic sectional curvature -1) in the Siegel model of $\mathbb{H}_{\mathbb{C}}^{m}$ is given by

$$
g_{m}=\frac{4}{h(z, w)^{2}}[(d w-\langle\langle d z, z\rangle\rangle)(d \bar{w}-\langle\langle z, d z\rangle\rangle)+h(z, w)\langle\langle d z, d z\rangle\rangle] .
$$

The stabilizer $P$ of $\infty$ in $\mathrm{PU}(m, 1)$ is the semi-direct product $\mathcal{N}^{2 m-1} \rtimes\left(\mathrm{U}(m-1) \times\left\{\phi_{s}\right\}_{s \in \mathbb{R}}\right)$ where $\mathcal{N}^{2 m-1}$ is the $(2 m-1)$-dimensional Heisenberg group, $\mathrm{U}(m-1)$ is the unitary group and $\left\{\phi_{s}\right\}_{s \in \mathbb{R}}$ is the one-parameter group corresponding to the horocyclic flow associated to $\infty$. The group $\mathcal{N}^{2 m-1}$ is a central extension of $\mathbb{C}^{m-1}$ and can be seen as $\mathbb{C}^{m-1} \times \mathbb{R}$ with product given by $\left(\xi_{1}, \nu_{1}\right)\left(\xi_{2}, \nu_{2}\right)=\left(\xi_{1}+\xi_{2}, \nu_{1}+\nu_{2}+2 \operatorname{Im}\left\langle\left\langle\xi_{1}, \xi_{2}\right\rangle\right\rangle\right)$. This is a two-step nilpotent group which acts simply transitively and isometrically on horospheres. Its center $\mathcal{Z}$ is the group of "vertical translations": $\{(0, \nu), \nu \in \mathbb{R}\}$.

If we set $u+i v=2 \bar{w}-\langle\langle z, z\rangle\rangle$, we obtain the so-called horospherical coordinates $(z, v, u) \in$ $\mathbb{C}^{m-1} \times \mathbb{R} \times \mathbb{R}_{+}^{\star}$, in which the action of $P$ on $\mathbb{H}_{\mathbb{C}}^{m}$ is given by:

$$
(\xi, \nu) A \phi_{s} \cdot(z, v, u)=\left(A e^{-s} z+\xi, e^{-2 s} v+\nu+2 \operatorname{Im}\left\langle\left\langle\xi, A e^{-s} z\right\rangle\right\rangle, e^{-2 s} u\right)
$$

and the metric $g_{m}$ takes the form

$$
g_{m}=\frac{d u^{2}}{u^{2}}+\frac{1}{u^{2}}(-d v+2 \operatorname{Im}\langle\langle z, d z\rangle\rangle)^{2}+\frac{4}{u}\langle\langle d z, d z\rangle\rangle .
$$

Replacing $u$ by $t=\log u$, the metric tensor decomposes as:

$$
g_{m}=d t^{2}+e^{-2 t}(-d v+2 \operatorname{Im}\langle\langle z, d z\rangle\rangle)^{2}+4 e^{-t}\langle\langle d z, d z\rangle\rangle .
$$

The coordinates $(z, v, t) \in \mathbb{C}^{m-1} \times \mathbb{R} \times \mathbb{R}$ will also be called horospherical coordinates.
A complex hyperbolic manifold $M$ of finite volume is the union of a compact part and a finite number of disjoint cusps. Each cusp $C$ of $M$ is diffeomorphic to the product $N \times[0,+\infty)$, where $N$ is a compact quotient of some horosphere $H S$ in $\mathbb{H}_{\mathbb{C}}^{m}$. We can assume that $H S$ is centered at $\infty$. The fundamental group $\Gamma_{C}$ of $C$, hence of $N$, can be identified with the stabilizer in $\Gamma$ of the horosphere $H S$ : it is therefore equal to $\Gamma \cap\left(\mathcal{N}^{2 m-1} \rtimes \mathrm{U}(m-1)\right)$.

If we call $\beta$ the 1 -form $-d v+2 \operatorname{Im}\langle\langle z, d z\rangle\rangle$ on $\mathbb{H}_{\mathbb{C}}^{m}$, it is easily checked that $\mathrm{d}^{c} t:=J \mathrm{~d} t=e^{-t} \beta$. Therefore, since $\mathcal{N}^{2 m-1} \rtimes \mathrm{U}(m-1)$ preserves the horospheres, $t, d t^{2}$, and $\beta$ are invariant by $\Gamma_{C}$. The decomposition of $g_{m}$ hence goes down to the cusp $C$ and we have:

$$
g_{m}=d t^{2}+e^{-2 t} \beta^{2}+e^{-t} g,
$$

where $g$ is the image of $4\langle\langle d z, d z\rangle\rangle$.
Remark. The Kähler form $\omega_{m}$, which we normalize so that $\omega_{m}(X, J X) \geq 0$, is of course exact on $\mathbb{H}_{\mathbb{C}}^{m}$. More precisely, $\omega_{m}=-\mathrm{dd}^{c} t=-\mathrm{d}\left(e^{-t} \beta\right)$. The invariance of $t$ and $\beta$ implies that this relation still holds in the cusps of $M$.

For lattices in $\operatorname{Sp}(m, 1), m \geq 2$, or in $\mathrm{F}_{4}^{-20}$, Corlette proves in Co92 a simple lemma that allows him to deduce the existence of finite energy equivariant maps. Here, the same idea will only provide the result for $m \geq 3$ :

Lemma 1.3. Assume $m \geq 3$. Then there exists a finite energy retraction of $M=\Gamma \backslash \mathbb{H}_{\mathbb{C}}^{m}$ onto a compact subset of $M$.

Proof. It is enough to construct the retraction on a cusp $C=N \times[0,+\infty)$ of $M$ : we define $r: N \times[0,+\infty) \longrightarrow N \times\{0\}$ obviously by $r(x, t)=(x, 0)$.

If $\left(\frac{\partial}{\partial t}, \frac{\partial}{\partial v}, e_{3}, \ldots, e_{2 m}\right)$ is an orthonormal basis of $T_{(x, 0)} C$ compatible with the splitting of $g_{m}$, then $\left(\frac{\partial}{\partial t}, e^{t} \frac{\partial}{\partial v}, e^{t / 2} e_{3}, \ldots, e^{t / 2} e_{2 m}\right)$ is such a basis of $T_{(x, t)} C$. Hence, $\|\mathrm{d} r\|_{(x, t)}^{2}=e^{2 t}+(2 m-2) e^{t}$. If we call $d V_{N}$ the volume element of $N \times\{0\}$, the volume element of $N \times\{t\}$ is given by $e^{\frac{1}{2}(-2 t-(2 m-2) t)} d V_{N}=e^{-m t} d V_{N}$. Hence, the energy of $r$ on $C$ is

$$
\frac{1}{2} \int_{C}\|\mathrm{~d} r\|^{2} d V_{m}=\frac{1}{2} \int_{0}^{+\infty} \int_{N}\left(2(m-1) e^{t}+e^{2 t}\right) e^{-m t} d V_{N} d t
$$

This is clearly finite if $m \geq 3$.
This retraction lifts to a map $\tilde{r}: \mathbb{H}_{\mathbb{C}}^{m} \longrightarrow \mathbb{H}_{\mathbb{C}}^{m}$, invariant by $\Gamma$. Therefore, if $f: \mathbb{H}_{\mathbb{C}}^{m} \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$ is any $\rho$-equivariant map, so is $f \circ \tilde{r}$, and its energy is finite. The theorem is therefore proved if $m \geq 3$.

In the case $m=2$, the energy density of the retraction $r$ grows like $e^{2 t}$ when $t$ goes to infinity whereas the volume element grows like $e^{-2 t}$ : the energy of $r$ is infinite and we need a deeper analysis of the situation in the cusps.

We fix a cusp $C=N \times[0,+\infty)$ of $M$ and we look for a finite energy map from the universal cover $H S \times[0,+\infty)$ of $C$ to $\mathbb{H}_{\mathbb{C}}^{n}$, equivariant w.r.t. the fundamental group $\Gamma_{C}$ of $C$ (equivalently, a section of the restriction of the flat bundle $M \times{ }_{\rho} \mathbb{H}_{\mathbb{C}}^{n}$ to $C \subset M$ ).

As we said, $\Gamma_{C}$ can be seen as a subgroup of $\mathcal{N} \rtimes \mathrm{U}(1)$, where now $\mathcal{N}:=\mathcal{N}^{3}$ is just $\mathbb{C} \times \mathbb{R}$. It follows from L. Auslander's generalization of Bieberbach's theorem (cf. Au60]) that $\Gamma_{\mathcal{N}}:=\Gamma_{C} \cap \mathcal{N}$ is a discrete uniform subgroup of $\mathcal{N}$, of finite index in $\Gamma_{C}$. Therefore (Au60, Lemma 1.3.), $\Gamma_{\mathcal{N}}$ cannot be contained in any proper analytic subgoup of $\mathcal{N}$. From this, it is easy to deduce that there exists $\varepsilon>0$ such that, for all $\gamma=\left(\xi_{\gamma}, \nu_{\gamma}\right) \in \Gamma_{\mathcal{N}}$, $\left|\xi_{\gamma}\right|>\varepsilon$ as soon as $\xi_{\gamma} \neq 0$. In other words, the image of the homomorphism of groups $T: \Gamma_{\mathcal{N}} \longrightarrow \mathbb{C}, \gamma \longmapsto \xi_{\gamma}$, is a lattice in $\mathbb{C}$. Let $\gamma_{1}=\left(\xi_{1}, \nu_{1}\right)$ and $\gamma_{2}=\left(\xi_{2}, \nu_{2}\right)$ be two elements of $\Gamma_{\mathcal{N}}$ such that $\xi_{1}$ and $\xi_{2}$ generate the lattice $T\left(\Gamma_{\mathcal{N}}\right)$. A straightforward computation yields $\left[\gamma_{1}, \gamma_{2}\right]:=\gamma_{1} \gamma_{2} \gamma_{1}^{-1} \gamma_{2}^{-1}=\left(0,-2 \operatorname{Im}\left(\xi_{1} \bar{\xi}_{2}\right)\right)$. Since $\xi_{1}$ and $\xi_{2}$ are linearly independent (over $\mathbb{R}$ ), $\operatorname{Im}\left(\xi_{1} \bar{\xi}_{2}\right) \neq 0$ and hence the subgroup $\Gamma_{\mathcal{Z}}:=\Gamma_{C} \cap \mathcal{Z}$ of $\Gamma_{\mathcal{N}}$ is non trivial. It is therefore isomorphic to $\mathbb{Z}$ and we call $\gamma_{0}$ its generator. Note that $\gamma_{0}, \gamma_{1}$ and $\gamma_{2}$ generate $\Gamma_{\mathcal{N}}$.

The construction of the equivariant map will depend on the type of $\rho\left(\gamma_{0}\right)$. Recall that an isometry of $\mathbb{H}_{\mathbb{C}}^{n}$ can be of one of the following (exclusive) 3 types:

- elliptic if it has a fixed point in $\mathbb{H}_{\mathbb{C}}^{n}$;
- parabolic if it has exactly one fixed point on the sphere at infinity of $\mathbb{H}_{\mathbb{C}}^{n}$ and no fixed points in $\mathbb{H}_{\mathbb{C}}^{n}$;
- hyperbolic if it has exactly two fixed points on the sphere at infinity of $\mathbb{H}_{\mathbb{C}}^{n}$ and no fixed points in $\mathbb{H}_{\mathbb{C}}^{n}$. In this case, the isometry acts by translation on the geodesic joining its fixed points at infinity.

Claim 1.4. $\rho\left(\gamma_{0}\right)$ can not be a hyperbolic isometry of $\mathbb{H}_{\mathbb{C}}^{n}$.
Proof. Assume that $\rho\left(\gamma_{0}\right)$ is a hyperbolic isometry of $\mathbb{H}_{\mathbb{C}}^{n}$ and call $A_{0}$ its axis (the geodesic joining its fixed points). Then, since $\gamma_{1}$ and $\gamma_{2}$ commute with $\gamma_{0}$, their images by $\rho$ commute with $\rho\left(\gamma_{0}\right)$, hence they must fix $A_{0}$ and act on it by translations: there exist $\tau_{1}, \tau_{2} \in \mathbb{R}$ such that $\rho\left(\gamma_{1}\right) A_{0}(t)=A_{0}\left(t+\tau_{1}\right)$ and $\rho\left(\gamma_{2}\right) A_{0}(t)=A_{0}\left(t+\tau_{2}\right)$. This implies that $\rho\left(\left[\gamma_{1}, \gamma_{2}\right]\right)$ acts trivialy on $A_{0}$. But $\left[\gamma_{1}, \gamma_{2}\right]=\gamma_{0}{ }^{p}$ for some $p \in \mathbb{Z}^{\star}$ and $\rho\left(\gamma_{0}\right)$ does not act trivially on $A_{0}$. This is a contradiction.

Hence $\rho\left(\gamma_{0}\right)$ is either elliptic or parabolic. In both cases we will start by constructing an equivariant map from the universal cover $H S \simeq \mathcal{N}$ of $N$ and then we shall extend it to the universal cover of the whole cusp.

Case 1: $\rho\left(\gamma_{0}\right)$ is parabolic. The idea is to find an equivariant map from $H S$ to a horosphere in $\mathbb{H}_{\mathbb{C}}^{n}$ centered at the fixed point of $\rho\left(\gamma_{0}\right)$ on the sphere at infinity $\partial \mathbb{H}_{\mathbb{C}}^{n}$ of $\mathbb{H}_{\mathbb{C}}^{n}$ and then to extend it to $H S \times[0,+\infty)$ using the horocyclic flow defined by the fixed point. Roughly speaking, when $t$ goes to infinity, the image of $H S \times\{t\}$ must go to infinity in $\mathbb{H}_{\mathbb{C}}^{n}$ fast enough so that the decay of the metric in $\mathbb{H}_{\mathbb{C}}^{n}$ prevents the energy density of the map from growing too quickly.

Using again the Siegel model for $\mathbb{H}_{\mathbb{C}}^{n}$, we may assume that the fixed point of $\rho\left(\gamma_{0}\right)$ is $\infty$. Since $\gamma_{0}$ is in the center of $\Gamma_{C}$, the whole group $\rho\left(\Gamma_{C}\right)$ must fix $\infty$, and therefore must be contained in its stabilizer in $\mathrm{PU}(n, 1)$. Moreover, $\rho\left(\Gamma_{C}\right)$ must stabilize each horosphere centered at $\infty$. For, if this was not the case, there would be an element $\gamma \in \Gamma_{C}$ such that $\rho(\gamma)$ is hyperbolic. But then, since $\gamma_{0}$ commutes with $\gamma, \rho\left(\gamma_{0}\right)$ would fix the axis of $\rho(\gamma)$. This is impossible since we assumed that $\rho\left(\gamma_{0}\right)$ is parabolic.

We see $\mathbb{H}_{\mathbb{C}}^{n}$ as $\mathbb{C}^{n-1} \times \mathbb{R} \times \mathbb{R}$ with horospherical coordinates $\left(z^{\prime}, v^{\prime}, t^{\prime}=\log u^{\prime}\right)$. The metric $g_{n}$ at a point $\left(z^{\prime}, v^{\prime}, t^{\prime}\right)$ is given by $g_{n}=d t^{\prime 2}+e^{-2 t^{\prime}}\left(-d v^{\prime}+2 \operatorname{Im}\left\langle\left\langle z^{\prime}, d z^{\prime}\right\rangle\right\rangle\right)^{2}+4 e^{-t^{\prime}}\left\langle\left\langle d z^{\prime}, d z^{\prime}\right\rangle\right\rangle$.

Let $H S^{\prime} \subset \mathbb{H}_{\mathbb{C}}^{n}$ be the horosphere $\mathbb{C}^{n-1} \times \mathbb{R} \times\{0\}$. The representation $\rho$ can be seen as a homomorphism from the fundamental group of $N$ to the isometry group of $H S^{\prime}$. Since $H S^{\prime}$ is contractible, there exists a $\rho$-equivariant map $\varphi$ from the universal cover $H S \subset \mathbb{H}_{\mathbb{C}}^{m}$ of $N$ to the horosphere $H S^{\prime} \subset \mathbb{H}_{\mathbb{C}}^{n}$. Now, define a $\rho$-equivariant map $f$ from the universal cover $H S \times[0,+\infty)$ of the cusp $C=N \times[0,+\infty)$ to $\mathbb{H}_{\mathbb{C}}^{n}$ by:

$$
\begin{aligned}
f: H S \times[0,+\infty) & \longrightarrow H S^{\prime} \times[0,+\infty) \subset \mathbb{H}_{\mathbb{C}}^{n} \\
(x, t) & \longmapsto(\varphi(x), 2 t)
\end{aligned}
$$

Using the same notation as in Lemma 1.3, the energy density of $f$ can be estimated as follows:

$$
\begin{aligned}
\|\mathrm{d} f\|_{(x, t)}^{2} & =\left|\mathrm{d} f\left(\frac{\partial}{\partial t}\right)\right|_{(\varphi(x), t)}^{2}+e^{2 t}\left|\mathrm{~d} f\left(\frac{\partial}{\partial v}\right)\right|_{(\varphi(x), t)}^{2}+e^{t} \sum_{k=3}^{4}\left|\mathrm{~d} f\left(e_{k}\right)\right|_{(\varphi(x), t)}^{2} \\
& \leq 4+e^{-2 t}\left(e^{2 t}\left|\mathrm{~d} \varphi\left(\frac{\partial}{\partial v}\right)\right|_{(\varphi(x), 0)}^{2}+e^{t} \sum_{k=3}^{4}\left|\mathrm{~d} \varphi\left(e_{k}\right)\right|_{(\varphi(x), 0)}^{2}\right) \\
& \leq 4+\|\mathrm{d} \varphi\|_{x}^{2}
\end{aligned}
$$

where $\|\mathrm{d} \varphi\|$ denotes the norm of the differential of $\varphi: H S \longrightarrow H S^{\prime}$ computed with the metrics of $H S$ and $H S^{\prime}$ induced from $g_{m}$ and $g_{n}$.

The energy of $f$ in the cusp $C$ is therefore finite since:

$$
E_{C}(f)=\frac{1}{2} \int_{C}\|\mathrm{~d} f\|^{2} d V_{m} \leq \frac{1}{2} \int_{0}^{+\infty} \int_{N}\left(4+\|\mathrm{d} \varphi\|^{2}\right) e^{-2 t} d V_{N} d t<+\infty
$$

Case 2: $\rho\left(\gamma_{0}\right)$ is elliptic. In this case, there is no canonical "direction" in which to send the slices $H S \times\{t\}$ to infinity in $\mathbb{H}_{\mathbb{C}}^{n}$. Once the equivariant map $f$ is constructed on $H S \times\{0\}$, the most natural way to define it on $H S \times\{t\}$ is to set $f_{\mid H S \times\{t\}}=f_{\mid H S \times\{0\}}$. Therefore, the growth of the energy density in the cusp cannot be controled by some decay of the metric in $\mathbb{H}_{\mathbb{C}}^{n}$, and we must control it at the start. We shall achieve this by demanding the equivariant map $H S \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$ to be constant in the "vertical direction" $\mathbb{R}$ of $H S=\mathbb{C} \times \mathbb{R}$.

As mentionned before, $\Gamma_{\mathcal{N}}$ is a finite index subgroup of $\Gamma_{C}$ and we have the tower of coverings:

$$
H S=\mathbb{C} \times \mathbb{R} \xrightarrow{\Gamma_{\mathcal{N}}} \widehat{N} \xrightarrow{\Gamma_{C} / \Gamma_{\mathcal{N}}} N,
$$

where $\widehat{N}=(\mathbb{C} \times \mathbb{R}) / \Gamma_{\mathcal{N}}$ is a circle bundle over the 2-torus $\mathbb{T}=\mathbb{C} / T\left(\Gamma_{\mathcal{N}}\right)$, and $\Gamma_{C} / \Gamma_{\mathcal{N}}$ can be seen as a finite subgroup of $\mathrm{U}(1)$, acting freely on this bundle.

The group $\Gamma_{C} / \Gamma_{\mathcal{N}}$ is generated by a primitive $p$-th root of unity $a$ and its action on $\mathbb{C}$ preserves the lattice $T\left(\Gamma_{\mathcal{N}}\right) \subset \mathbb{C}$. This implies that $a$ is a root of a degree 2 polynomial with integer coefficients and hence the possible values of $a$ are $1,-1, e^{i \frac{2 \pi}{3}}, i$, or $e^{i \frac{\pi}{3}}$. On the other hand, the number of possible lattices is also restricted:

- if $a=1$ or $a=-1, T\left(\Gamma_{\mathcal{N}}\right)$ can be any lattice of $\mathbb{C}$;
- if $a=i, T\left(\Gamma_{\mathcal{N}}\right)$ must be a square lattice, meaning that we can choose the first two generators $\gamma_{1}=\left(\xi_{1}, \nu_{1}\right)$ and $\gamma_{2}=\left(\xi_{2}, \nu_{2}\right)$ of $\Gamma_{\mathcal{N}}$ so that $\xi_{2}=i \xi_{1}$;
- if $a=e^{i \frac{2 \pi}{3}}$ or $a=e^{i \frac{\pi}{3}}, T\left(\Gamma_{\mathcal{N}}\right)$ must be an equilateral triangle lattice, meaning that we can choose $\gamma_{1}$ and $\gamma_{2}$ so that $\xi_{2}=e^{i \frac{\pi}{3}} \xi_{1}$.
We start with the case $a=1$, namely $\Gamma_{\mathcal{N}}=\Gamma_{C}$. We want to define a map $\varphi$ from $\mathbb{C}$ to $\mathbb{H}_{\mathbb{C}}^{n}$ and then to extend $\varphi$ to $\mathbb{C} \times \mathbb{R}$ by $\varphi(z, v)=\varphi(z)$, so that this extended map is equivariant w.r.t. the action of $\Gamma_{C}$. An obvious necessary condition is that $\varphi: \mathbb{C} \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$ must be equivariant w.r.t. the action of $T\left(\Gamma_{C}\right)$ on $\mathbb{C}$. Another necessary condition is that $\varphi$ should send $\mathbb{C}$ to the fixed points set $\mathrm{Fix}_{0}$ of $\rho\left(\gamma_{0}\right)$ in $\mathbb{H}_{\mathbb{C}}^{n}$. Indeed, for any $z \in \mathbb{C}, \gamma_{0}(z, 0)$ belongs to $\{z\} \times \mathbb{R}$ and $\varphi$ maps $\{z\} \times \mathbb{R}$ to the point $\varphi(z)$. These two conditions are also clearly sufficient.

So let $x_{0} \in \mathbb{H}_{\mathbb{C}}^{n}$ be a fixed point of $\rho\left(\gamma_{0}\right)$ and set $\varphi(0)=x_{0}$. Since $\gamma_{i}=\left(\xi_{i}, \nu_{i}\right), i=1$ or 2 , commutes with $\gamma_{0}$, the point $x_{i}=\rho\left(\gamma_{i}\right) x_{0}$ must also be fixed by $\rho\left(\gamma_{0}\right)$. Let $\sigma_{0 i}$ be the geodesic arc in $\mathbb{H}_{\mathbb{C}}^{n}$ joining $x_{0}$ to $x_{i}$. Note that $\mathrm{Fix}_{0}$ is a convex subset of $\mathbb{H}_{\mathbb{C}}^{n}$ and hence $\sigma_{0 i}$ is included in $\operatorname{Fix}_{0}$. Let $\varphi$ map the segment $\left[0, \xi_{i}\right]$ onto $\sigma_{0 i}$. We then map the segment $\left[\xi_{1}, \xi_{1}+\xi_{2}\right]$ to $\rho\left(\gamma_{1}\right) \sigma_{02}$ and the segment $\left[\xi_{2}, \xi_{1}+\xi_{2}\right]$ to $\rho\left(\gamma_{2}\right) \sigma_{01}$. This is well defined since $\rho\left(\gamma_{1}\right)\left(x_{2}\right)=\rho\left(\gamma_{1} \gamma_{2}\right)\left(x_{0}\right)=$ $\rho\left(\gamma_{2} \gamma_{1}\right) \rho\left(\gamma_{1}^{-1} \gamma_{2}^{-1} \gamma_{1} \gamma_{2}\right) x_{0}=\rho\left(\gamma_{2} \gamma_{1}\right) \rho\left(\gamma_{0}^{k}\right) x_{0}=\rho\left(\gamma_{2} \gamma_{1}\right) x_{0}=\rho\left(\gamma_{2}\right)\left(x_{1}\right)$. Moreover, because of the commutation of $\gamma_{1}$ and $\gamma_{2}$ with $\gamma_{0}, \rho\left(\gamma_{1}\right) \sigma_{02}$ and $\rho\left(\gamma_{2}\right) \sigma_{01}$ are included in Fix ${ }_{0}$.

Hence we get an equivariant map $\varphi$ from the boundary of a fundamental domain of $T\left(\Gamma_{C}\right)$ in $\mathbb{C}$ to $\operatorname{Fix}_{0}$ ( $\varphi$ can be made smooth, for example by taking it constant near $0, \xi_{1}$ and $\xi_{2}$ ). We can therefore extend $\varphi$ to a $T\left(\Gamma_{C}\right)$-equivariant map from $\mathbb{C}$ to $\mathrm{Fix}_{0}$.

Define now $f: H S \times[0,+\infty)=\mathbb{C} \times \mathbb{R} \times[0,+\infty) \longrightarrow \operatorname{Fix}_{0} \subset \mathbb{H}_{\mathbb{C}}^{n}$ by $f(z, v, t)=\varphi(z)$. The map $f$ is $\rho$-equivariant and its energy density is:

$$
\|\mathrm{d} f\|_{(x, t)}^{2}=\left|\mathrm{d} f\left(\frac{\partial}{\partial t}\right)\right|_{\varphi(x)}^{2}+e^{2 t}\left|\mathrm{~d} f\left(\frac{\partial}{\partial v}\right)\right|_{\varphi(x)}^{2}+e^{t} \sum_{k=3}^{4}\left|\mathrm{~d} f\left(e_{k}\right)\right|_{\varphi(x)}^{2}=0+0+e^{t}\|\mathrm{~d} \varphi\|_{x}^{2}
$$

where $\|\mathrm{d} \varphi\|$ denotes the norm of the differential of $\varphi: \mathbb{C} \longrightarrow \mathrm{Fix}_{0}$ computed with the metrics of $\mathbb{C} \times\{0\} \subset H S \subset \mathbb{H}_{\mathbb{C}}^{m}$ and $\mathrm{Fix}_{0} \subset \mathbb{H}_{\mathbb{C}}^{n}$ induced from $g_{m}$ and $g_{n}$.

Therefore, $E_{C}(f)=\frac{1}{2} \int_{0}^{+\infty} \int_{N}\|\mathrm{~d} \varphi\|^{2} e^{-t} d V_{N} d t<+\infty$.
Now, consider the cases where $a \neq 1$. We want to proceed as we just did, namely, we want to first construct a map $\varphi$ from $\mathbb{C}$ to $\mathbb{H}_{\mathbb{C}}^{n}$ and then extend it to $H S$ by requiring that
$\varphi(z, v)=\varphi(z)$. The two conditions we mentioned are of course still necessary but we need to be more careful because of the action of $\Gamma_{C} / \Gamma_{\mathcal{N}}$.

Let $\gamma_{3}$ be an element of $\Gamma_{C}$ such that $\gamma_{3} \Gamma_{\mathcal{N}}=a$. Then $\gamma_{3}=\left(\xi_{3}, \nu_{3}, a\right)$ for some $\xi_{3} \in \mathbb{C}$ and $\nu_{3} \in \mathbb{R}$. It is easy to check that $\gamma_{0}, \gamma_{1}, \gamma_{2}$ and $\gamma_{3}$ generate $\Gamma_{C}$.

The first thing to notice is that the point $\zeta=\frac{\xi_{3}}{1-a}$ is fixed by the action of $\gamma_{3}$ on the $\mathbb{C}$-factor. Since we want $\varphi$ to be constant on $\{\zeta\} \times \mathbb{R}, \varphi$ must send $\zeta$ to a fixed point of $\rho\left(\gamma_{3}\right)$. This can be done because of the:

Claim 1.5. Let $\gamma=(\xi, \nu, b) \in \Gamma_{C}$ be such that $b \neq 1$. Then $\rho(\gamma)$ and $\rho\left(\gamma_{0}\right)$ have a common fixed point in $\mathbb{H}_{\mathbb{C}}^{n}$.

Proof. Since $\gamma$ and $\gamma_{0}$ commute, $\rho(\gamma)$ stabilizes the totally geodesic submanifold $\mathrm{Fix}_{0}$ of $\mathbb{H}_{\mathbb{C}}^{n}$. Let $q$ be such that $b^{q}=1$. Computing, we get $\gamma^{q}=(\xi, \nu, b)^{q}=\left(\left(\sum_{k=0}^{q-1} b^{k}\right) \xi, v, b^{q}\right)=(0, v, 1)$ for some $v \in \mathbb{R}$. Hence $\gamma^{q}$ belongs to $\Gamma_{\mathcal{Z}}: \gamma^{q}$ is a power of $\gamma_{0}$. The orbit under the group generated by $\rho(\gamma)$ of any point in $\mathrm{Fix}_{0}$ must therefore be finite and this implies that the action of $\rho(\gamma)$ on Fix $_{0}$ has a fixed point.

With this in mind, it is now possible to complete the proof by constructing $\varphi$ on the boundary $\partial F$ of a fundamental domain $F$ of the action of $\Gamma_{C}$ on the $\mathbb{C}$-factor. Since $\gamma_{3}$ acts on $\mathbb{C}$ by rotation around its fixed point $\zeta$, we can choose a fundamental domain $G$ of the action of $T\left(\Gamma_{\mathcal{N}}\right)$ on $\mathbb{C}$, centered at $\zeta$ and invariant by $\gamma_{3}$. For $F$ we then take a fundamental domain for the action of $\gamma_{3}$ on $G$.

We do it in the case $a=e^{i \frac{\pi}{3}}$, the other cases are handled similarly.
The lattice $T\left(\Gamma_{\mathcal{N}}\right)$ is generated by $\xi_{1}$ and $\xi_{2}=e^{i \frac{\pi}{3}} \xi_{1}$. Let $G$ be the regular hexagon centered at $\zeta$ with one vertex at the point $\zeta+\frac{1}{3}\left(\xi_{1}+\xi_{2}\right)$. G is a fundamental domain for the action of $T\left(\Gamma_{\mathcal{N}}\right)$ and it is invariant by $\gamma_{3}$. Let then $F$ be the quadrilateral whose vertices are $\zeta, \zeta+\frac{1}{2} \xi_{1}$, $\zeta+\frac{1}{3}\left(\xi_{1}+\xi_{2}\right)$ and $\zeta+\frac{1}{2} \xi_{2}$. $F$ is clearly a fundamental domain for the action of $\Gamma_{C}$ on $\mathbb{C}$. See Fig. 1 for a picture.


Fig. 1
Let now $x_{0} \in \mathbb{H}_{\mathbb{C}}^{n}$ be a fixed point of both $\gamma_{0}$ and $\gamma_{3}$ (such a point exists by Claim 1.5). Set $\varphi(\zeta)=x_{0}$.

The point $\zeta+\frac{1}{2} \xi_{1}$ is fixed by $\gamma_{1} \gamma_{3}^{3}$, hence it must be sent by $\varphi$ to a fixed point of $\rho\left(\gamma_{1} \gamma_{3}^{3}\right)$. It follows from Claim 1.5 that $\rho\left(\gamma_{1} \gamma_{3}^{3}\right)$ and $\rho\left(\gamma_{0}\right)$ have a common fixed point, say $x_{1}$. Let $\varphi$ send the vertex $\zeta+\frac{1}{2} \xi_{1}$ to $x_{1}$ and the edge $\left[\zeta, \zeta+\frac{1}{2} \xi_{1}\right]$ of $F$ to the geodesic arc $\sigma_{01}$ joining $x_{0}$
 it to a fixed point $x_{2}$ of $\rho\left(\gamma_{2} \gamma_{3}^{4}\right)$ in Fix ${ }_{0}$. We map the edge $\left[\zeta+\frac{1}{2} \xi_{1}, \zeta+\frac{1}{3}\left(\xi_{1}+\xi_{2}\right)\right]$ to the geodesic arc $\sigma_{12}$ joining $x_{1}$ and $x_{2}$ in Fix 0 .

Now the edge $\left[\zeta, \zeta+\frac{1}{2} \xi_{2}\right]$ is the image of $\left[\zeta, \zeta+\frac{1}{2} \xi_{1}\right]$ under $\gamma_{3}$ so we must map it to $\rho\left(\gamma_{3}\right)\left(\sigma_{01}\right)$. In the same way, $\left[\zeta+\frac{1}{2} \xi_{2}, \zeta+\frac{1}{3}\left(\xi_{1}+\xi_{2}\right)\right]$ is the image of $\left[\zeta+\frac{1}{2} \xi_{1}, \zeta+\frac{1}{3}\left(\xi_{1}+\xi_{2}\right)\right]$ by $\gamma_{2} \gamma_{3}^{4}$ and we must therefore map it to $\rho\left(\gamma_{2} \gamma_{3}^{4}\right)\left(\sigma_{12}\right)$. These definitions of $\varphi$ agree at the point $\frac{1}{2} \xi_{2}$. Indeed, a simple computation shows that there exists $q$ such that $\gamma_{2} \gamma_{3}=\gamma_{0}^{q} \gamma_{3} \gamma_{1}$ and therefore, $\rho\left(\gamma_{2} \gamma_{3}^{4}\right) x_{1}=\rho\left(\gamma_{2} \gamma_{3} \gamma_{3}^{3}\right) x_{1}=\rho\left(\gamma_{0}^{q} \gamma_{3} \gamma_{1} \gamma_{3}^{3}\right) x_{1}=\rho\left(\gamma_{0}^{q} \gamma_{3}\right) \rho\left(\gamma_{1} \gamma_{3}^{3}\right) x_{1}=\rho\left(\gamma_{3}\right) \rho\left(\gamma_{0}^{q}\right) x_{1}=\rho\left(\gamma_{3}\right) x_{1}$. Hence $\varphi$ is well defined on $\partial F$. By construction, $\varphi$ is equivariant w.r.t. $\gamma_{0}$ and the facepairings $\gamma_{3}$ and $\gamma_{2} \gamma_{3}^{4}$ which generate the whole group $\Gamma_{C}$.

The construction of $\varphi$ and $f$ then goes on as in the case $a=1$.
In this way we obtain a section $f_{i}$ of the bundle $M \times{ }_{\rho} \mathbb{H}_{\mathbb{C}}^{n}$ on each cusp $C_{i}$ of $M$. This section can be extended to a section $f$ defined on the whole manifold $M$ and since the energy of $f_{i}: C_{i} \longrightarrow M \times_{\rho} \mathbb{H}_{\mathbb{C}}^{n}$ is finite for each $i$, the energy of $f: M \longrightarrow M \times_{\rho} \mathbb{H}_{\mathbb{C}}^{n}$ is finite and we are done.

## 2. Pluriharmonicity and consequences

In this section, we study the properties of finite energy harmonic maps $\mathbb{H}_{\mathbb{C}}^{m} \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$ which are equivariant w.r.t. a representation $\rho$ of a torsion-free lattice $\Gamma<\mathrm{PU}(m, 1)$ into $\mathrm{PU}(n, 1)$.

### 2.1. Pluriharmonicity.

Theorem 2.1. Let $f: \mathbb{H}_{\mathbb{C}}^{m} \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$ be a $\rho$-equivariant harmonic map of finite energy. Then $f$ is pluriharmonic, namely, the $J$-invariant part $(\nabla \mathrm{d} f)^{1,1}$ of $\nabla \mathrm{d} f$ vanishes identically. Moreover the complexified sectional curvature of $\mathbb{H}_{\mathbb{C}}^{n}$ is zero on $\mathrm{d} f\left(T^{1,0} \mathbb{H}_{\mathbb{C}}^{m}\right)$.

We first prove a general Bochner-type formula due to Mok, Siu and Yeung (cf. [MSY93]) in case $\Gamma$ is a uniform lattice. We state it in the case of maps $\mathbb{H}_{\mathbb{C}}^{m} \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$ but it is valid in the more general setting of equivariant maps from an irreducible (rank 1) symmetric space of non-compact type to a negatively curved manifold, as can be seen from the proof. Our exposition follows [Pa95].

Let $R^{m}$ and $R^{n}$ be the curvature tensors of $g_{m}$ and $g_{n}$, and $Q$ be any parallel tensor of curvature type on $\mathbb{H}_{\mathbb{C}}^{m}$. For $h$ a symmetric 2 -tensor with values in a vector bundle over $\mathbb{H}_{\mathbb{C}}^{m}$, define $(\stackrel{\circ}{Q} h)(X, Y)=\operatorname{tr}(W \longmapsto h(Q(W, X) Y, W))$.

Remark that if $f$ is a $\rho$-equivariant map $\mathbb{H}_{\mathbb{C}}^{m} \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$, then, since $Q$ is parallel and $\rho(\Gamma)$ acts by isometries on $\mathbb{H}_{\mathbb{C}}^{n}$, the $\mathbb{R}$-valued functions $\left\langle Q, f^{\star} R^{n}\right\rangle$ and $\langle\stackrel{\circ}{Q} \nabla \mathrm{~d} f, \nabla \mathrm{~d} f\rangle$ on $\mathbb{H}_{\mathbb{C}}^{m}$ are in fact $\rho$-invariant and hence can be considered as functions on $M=\Gamma \backslash \mathbb{H}_{\mathbb{C}}^{m}$ :

Proposition 2.2. Let $f$ be a $\rho$-equivariant harmonic map of finite energy from $\mathbb{H}_{\mathbb{C}}^{m}$ to $\mathbb{H}_{\mathbb{C}}^{n}$ and $Q$ a parallel tensor of curvature type on $\mathbb{H}_{\mathbb{C}}^{m}$. Then,

$$
\int_{M}\left[\langle\stackrel{\circ}{Q} \nabla \mathrm{~d} f, \nabla \mathrm{~d} f\rangle-\frac{1}{2}\left\langle Q, f^{\star} R^{n}\right\rangle\right] d V_{m}=-\frac{1}{4 m} \int_{M}\left\langle Q, R^{m}\right\rangle\|\mathrm{d} f\|^{2} d V_{m}
$$

where if $\Gamma$ is non-uniform, that is if $M$ is non-compact, the left-hand side should read $\lim _{R \rightarrow \infty} \int_{M} \eta_{R}\left[\langle\stackrel{\circ}{Q} \nabla \mathrm{~d} f, \nabla \mathrm{~d} f\rangle-\frac{1}{2}\left\langle Q, f^{\star} R^{n}\right\rangle\right] d V_{m}$, for $\left\{\eta_{R}\right\}$ a well-chosen family of cut-off functions on $M$.
Proof. Let us first assume that $M$ is compact. All the computations will be made in a normal coordinates system.

By definition, $(\stackrel{\circ}{Q} \nabla \mathrm{~d} f)(X, Y)=\sum_{k}(\nabla \mathrm{~d} f)\left(Q\left(e_{k}, X\right) Y, e_{k}\right)$. Since $\mathrm{d} f$ is closed, i.e. $\nabla \mathrm{d} f$ is symmetric, and $Q$ is parallel, we have in fact
$\stackrel{\circ}{Q} \nabla \mathrm{~d} f(X, Y)=\sum_{k}\left(\nabla_{e_{k}} \mathrm{~d} f\right)\left(Q\left(e_{k}, X\right) Y\right)=\sum_{k}\left(\nabla_{e_{k}} \mathrm{~d} f \circ Q\right)\left(e_{k}, X\right) Y=-\nabla^{\star}(\mathrm{d} f \circ Q)(X, Y)$,
where $\nabla^{\star}$ is the formal adjoint of $\nabla$ : if $T$ is a $(p+1)$-tensor, $\left(\nabla^{\star} T\right)\left(X_{1}, \ldots, X_{p}\right):=-\operatorname{tr}(W \longmapsto$ $\left.\left(\nabla_{W} T\right)\left(W, X_{1}, \ldots, X_{p}\right)\right)$.

Integrating this relation over $M$ (we assumed $M$ compact), we get

$$
\int_{M}\langle\stackrel{\circ}{Q} \nabla \mathrm{~d} f, \nabla \mathrm{~d} f\rangle d V_{m}=-\int_{M}\left\langle\nabla^{\star}(\mathrm{d} f \circ Q), \nabla \mathrm{d} f\right\rangle d V_{m}=-\int_{M}\left\langle\mathrm{~d} f \circ Q, \nabla^{2} \mathrm{~d} f\right\rangle d V_{m}
$$

where $\nabla^{2} \mathrm{~d} f$ is the 3 -tensor $\nabla(\nabla \mathrm{d} f)$.
Using that $Q$, and hence $\mathrm{d} f \circ Q$, is skew-symmetric in its first two variables, one checks that

$$
\left\langle\mathrm{d} f \circ Q, \nabla^{2} \mathrm{~d} f\right\rangle=-\frac{1}{2}\left[\left\langle\mathrm{~d} f \circ Q, f^{\star} R^{n}\right\rangle-\left\langle\mathrm{d} f \circ Q, \mathrm{~d} f \circ R^{m}\right\rangle\right]
$$

We have $\left\langle\mathrm{d} f \circ Q, f^{\star} R^{n}\right\rangle=\left\langle Q, f^{\star} R^{n}\right\rangle$, where in the r.h.s. $f^{\star} R^{n}$ and $Q$ are considered as (4,0)-tensors. Moreover, computations show that

$$
\left\langle\mathrm{d} f \circ Q, \mathrm{~d} f \circ R^{m}\right\rangle=\sum_{a, b} \frac{1}{2}\left(\left\langle\iota_{e_{a}} Q, \iota_{e_{b}} R^{m}\right\rangle+\left\langle\iota_{e_{b}} Q, \iota_{e_{a}} R^{m}\right\rangle\right) f^{\star} g_{n}\left(e_{a}, e_{b}\right)
$$

where $\iota$ denotes interior product. Now, since $M$ is locally symmetric, the symmetric 2 -tensor $\theta$ given by

$$
\theta(X, Y)=\frac{1}{2}\left(\left\langle\iota_{X} Q, \iota_{Y} R^{m}\right\rangle+\left\langle\iota_{Y} Q, \iota_{X} R^{m}\right\rangle\right)
$$

is parallel. Thus it must be proportionnal to $g_{m}\left(M\right.$ is locally irreducible): $\theta=\frac{1}{2 m}\left(\operatorname{tr}_{g_{m}} \theta\right) g_{m}$. Now, $\operatorname{tr}_{g_{m}} \theta=\left\langle Q, R^{m}\right\rangle$ and $\left\langle g_{m}, f^{\star} g_{n}\right\rangle=\|\mathrm{d} f\|^{2}$, so that $\left\langle\mathrm{d} f \circ Q, \mathrm{~d} f \circ R^{m}\right\rangle=\frac{1}{2 m}\left\langle Q, R^{m}\right\rangle\|\mathrm{d} f\|^{2}$ and hence

$$
\int_{M}\langle\stackrel{\circ}{Q} \nabla \mathrm{~d} f, \nabla \mathrm{~d} f\rangle d V_{m}=\frac{1}{2} \int_{M}\left[\left\langle Q, f^{\star} R^{n}\right\rangle-\frac{1}{2 m}\left\langle Q, R^{m}\right\rangle\|\mathrm{d} f\|^{2}\right] d V_{m}
$$

This ends the proof in the compact case.
Now assume $M$ is non-compact of finite volume. The only global step in the preceding proof is the initial integration by parts. Thus we only have to show that this can be done in the finite volume case. We mimic the argument given by Corlette in Co92.

As mentionned earlier, $M$ is the union of a compact manifold with boundary $M_{0}$ and of a finite number of pairwise disjoint cusps $C_{i}$, each diffeomorphic to a compact ( $2 m-1$ )-manifold $N_{i}$ times $[0,+\infty)$. For each $i$, let $t_{i}$ be the parameter in the $[0,+\infty)$ factor.

For $R>1$, we define a cut-off function $\eta_{R}$ on $M$ in the following manner. Take a smooth function $\eta$ on $[0,+\infty)$ identically equal to 1 on $[0,1]$ and to 0 on $[2,+\infty)$. Set

$$
\eta_{R}(x)= \begin{cases}1 & \text { if } x \in M_{0} \\ \eta\left(\frac{t_{i}}{R}\right) & \text { if } x \in C_{i}\end{cases}
$$

Since $\eta_{R}$ is a horofunction along each cusp (cf. [Hel84], II.3.8), the absolute value $\left|\Delta \eta_{R}\right|$ of its Laplacian is bounded independently of $R$. Moreover, the norm $\left\|\mathrm{d} \eta_{R}\right\|$ of its differential is bounded by a constant times $\frac{1}{R}$.

Introducing $\eta_{R}$ in the integration by parts, we obtain

$$
\begin{aligned}
\int_{M} \eta_{R}\langle\stackrel{\circ}{Q} \nabla \mathrm{~d} f, \nabla \mathrm{~d} f\rangle d V_{m} & =-\int_{M}\left\langle\nabla^{\star}(\mathrm{d} f \circ Q), \eta_{R} \nabla \mathrm{~d} f\right\rangle d V_{m} \\
& =-\int_{M}\left\langle\mathrm{~d} f \circ Q, \eta_{R} \nabla^{2} \mathrm{~d} f+\mathrm{d} \eta_{R} \otimes \nabla \mathrm{~d} f\right\rangle d V_{m} \\
& =-\int_{M} \eta_{R}\left\langle\mathrm{~d} f \circ Q, \nabla^{2} \mathrm{~d} f\right\rangle d V_{m}-\int_{M}\left\langle\mathrm{~d} f \circ Q, \mathrm{~d} \eta_{R} \otimes \nabla \mathrm{~d} f\right\rangle d V_{m} \\
& =\frac{1}{2} \int_{M} \eta_{R}\left[\left\langle Q, f^{\star} R^{n}\right\rangle-\frac{1}{2 m}\left\langle Q, R^{m}\right\rangle\|\mathrm{d} f\|^{2}\right] d V_{m} \\
& \quad-\int_{M}\left\langle\mathrm{~d} f \circ Q, \mathrm{~d} \eta_{R} \otimes \nabla \mathrm{~d} f\right\rangle d V_{m} .
\end{aligned}
$$

Thus,

$$
\begin{aligned}
\int_{M} \eta_{R}\left[\langle\stackrel{\circ}{Q} \nabla \mathrm{~d} f, \nabla \mathrm{~d} f\rangle-\frac{1}{2}\left\langle Q, f^{\star} R^{n}\right\rangle\right] d V_{m}= & -\frac{1}{4 m} \int_{M} \eta_{R}\left\langle Q, R^{m}\right\rangle\|\mathrm{d} f\|^{2} d V_{m} \\
& -\int_{M}\left\langle\mathrm{~d} f \circ Q, \mathrm{~d} \eta_{R} \otimes \nabla \mathrm{~d} f\right\rangle d V_{m}
\end{aligned}
$$

The tensors $Q$ and $R^{m}$ are parallel and hence $\left\langle Q, R^{m}\right\rangle$ is constant on $M$. Therefore, $\left\langle Q, R^{m}\right\rangle\|\mathrm{d} f\|^{2}$ is integrable and the first term in the r.h.s. goes to $-\frac{1}{4 m} \int_{M}\left\langle Q, R^{m}\right\rangle\|\mathrm{d} f\|^{2} d V_{m}$ as $R$ goes to infinity. On the other hand we have $\left\|\mathrm{d} \eta_{R}\right\| \leq \frac{C}{R}$ for some constant $C$ independent of $R$ and hence

$$
\begin{aligned}
\left(\int_{M}\left\langle\mathrm{~d} f \circ Q, \mathrm{~d} \eta_{R} \otimes \nabla \mathrm{~d} f\right\rangle d V_{m}\right)^{2} & \leq\left(\int_{M}\|\mathrm{~d} f \circ Q\|^{2} d V_{m}\right)\left(\int_{M}\left\|\mathrm{~d} \eta_{R} \otimes \nabla \mathrm{~d} f\right\|^{2} d V_{m}\right) \\
& \leq\left(\int_{M} \frac{1}{2 m}\|Q\|^{2}\|\mathrm{~d} f\|^{2} d V_{m}\right)\left(\int_{M} 2 m\left\|\mathrm{~d} \eta_{R}\right\|^{2}\|\nabla \mathrm{~d} f\|^{2} d V_{m}\right) \\
& \leq \frac{C^{2}}{R^{2}}\left(\int_{M}\|Q\|^{2}\|\mathrm{~d} f\|^{2} d V_{m}\right)\left(\int_{M}\|\nabla \mathrm{~d} f\|^{2} d V_{m}\right) .
\end{aligned}
$$

Since $Q$ is parallel, $\|Q\|$ is constant and $\int_{M}\|Q\|^{2}\|\mathrm{~d} f\|^{2} d V_{m}$ is finite.
The next lemma implies that $\lim _{R \rightarrow \infty} \int_{M}\left\langle\mathrm{~d} f \circ Q, \mathrm{~d} \eta_{R} \otimes \nabla \mathrm{~d} f\right\rangle d V_{m}=0$ and therefore ends the proof of Proposition 2.2.

Lemma 2.3. $\|\nabla \mathrm{d} f\|$ belongs to $L^{2}(M): \int_{M}\|\nabla \mathrm{~d} f\|^{2} d V_{m}<+\infty$.
Proof. Because the energy density $e$ of $f$ is integrable on $M$, and using Green's formula, we see that

$$
\int_{M}(\Delta e) \eta_{R} d V_{m}=\int_{M} e\left(\Delta \eta_{R}\right) d V_{m}
$$

is bounded independently of $R$. Now, since we assumed that $f$ is harmonic, the Bochner-type formula of Eells-Sampson ( $[$ ES64 $)$ reads:

$$
\Delta e=-2\|\nabla \mathrm{~d} f\|^{2}+\operatorname{Scal}\left(f^{\star} R^{n}\right)-\left\langle\mathrm{d} f \circ \operatorname{Ric}^{m}, \mathrm{~d} f\right\rangle
$$

where $\operatorname{Scal}\left(f^{\star} R^{n}\right)$ denotes the scalar contraction of the curvature tensor $f^{\star} R^{n}$ and $\operatorname{Ric}^{m}$ is the Ricci tensor of $g_{m}$ seen as an endomorphism of $T \mathbb{H}_{\mathbb{C}}^{m}$. Since $\mathbb{H}_{\mathbb{C}}^{n}$ is negatively curved and $\operatorname{Ric}^{m}=-\frac{1}{2}(m+1) \mathrm{Id}$, we get $2\|\nabla \mathrm{~d} f\|^{2} \leq-\Delta e+(m+1) e$ and thus $\|\nabla \mathrm{d} f\|^{2}$ is integrable.

Let us call $I$, resp. $I_{\mathbb{C}}$, the (3,1)-tensor of curvature type on $M=\Gamma \backslash \mathbb{H}_{\mathbb{C}}^{m}\left(\right.$ or on $\left.\mathbb{H}_{\mathbb{C}}^{m}\right)$ given by $I(X, Y) Z=g_{m}(X, Z) Y-g_{m}(Y, Z) X$, resp. $I_{\mathbb{C}}(X, Y) Z=\frac{1}{4}(I(X, Y) Z+I(J X, J Y) Z+$
$\left.2 g_{m}(J X, Y) J Z\right)$, for all $X, Y, Z \in T \mathbb{H}_{\mathbb{C}}^{m}$. The curvature tensor $R^{m}$ of $M$ (or of $\mathbb{H}_{\mathbb{C}}^{m}$ ) is just $-I_{\mathbb{C}}$. Both $I$ and $I_{\mathbb{C}}$ are parallel tensors, and in fact they form a basis of the space of parallel tensors of curvature type on $M . I$ and $I_{\mathbb{C}}$ will also denote the corresponding (4,0)-tensors.

We will apply the Bochner-type formula $(\diamond)$ to the parallel tensor of curvature type $Q=$ $I_{\mathbb{C}}-I$.

Lemma 2.4. Let $f: \mathbb{H}_{\mathbb{C}}^{m} \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$ be a harmonic map and let $Q=I_{\mathbb{C}}-I$.
Then $\langle\stackrel{\circ}{Q} \nabla \mathrm{~d} f, \nabla \mathrm{~d} f\rangle=-\frac{3}{2}\left\|(\nabla \mathrm{~d} f)^{(1,1)}\right\|^{2}$, where $(\nabla \mathrm{d} f)^{(1,1)}$ is the $J$-invariant part of $\nabla \mathrm{d} f$ : for all $X, Y \in T \mathbb{H}_{\mathbb{C}}^{m},(\nabla \mathrm{~d} f)^{(1,1)}(X, Y):=\frac{1}{2}[\nabla \mathrm{~d} f(X, Y)+\nabla \mathrm{d} f(J X, J Y)]$.

Proof. A straightforward computation shows that for $h$ a symmetric 2-tensor taking values in $f^{\star} T \mathbb{H}_{\mathbb{C}}^{n}, \stackrel{\circ}{I} h=h-g_{m} \operatorname{tr}_{g_{m}} h$ and $\stackrel{\circ}{I_{\mathbb{C}}} h(X, Y)=\frac{1}{4}\left[h(X, Y)-3 h(J X, J Y)-g_{m}(X, Y) \operatorname{tr}_{g_{m}} h\right]$. Therefore, since $\operatorname{tr}_{g_{m}} \nabla \mathrm{~d} f=0, \stackrel{\circ}{Q} \nabla \mathrm{~d} f=-\frac{3}{2}(\nabla \mathrm{~d} f)^{(1,1)}$. The decomposition of a 2-tensor in $J$-invariant and $J$-skew-invariant parts is orthogonal, hence the result.

Lemma 2.5. $I_{\mathbb{C}}$ is the orthogonal projection of $I$ onto the space of Kähler curvature type tensors, namely, the space of tensors of curvature type $T$ such that $T(X, Y) J Z=J(T(X, Y) Z)$, for all $X, Y, Z \in T \mathbb{H}_{\mathbb{C}}^{m}$.
Proof. Since $I_{\mathbb{C}}$ is clearly of Kähler curvature type, it remains to show that $I_{\mathbb{C}}-I$ is orthogonal to all tensors of Kähler curvature type. Simple computations show that if $T$ is any tensor of curvature type, $\langle I, T\rangle=2 \operatorname{Scal}(T)$, whereas

$$
\left\langle I_{\mathbb{C}}, T\right\rangle=\frac{1}{2} \operatorname{Scal}(T)-\frac{1}{2} \sum_{k, l=1}^{2 m}\left(T\left(e_{k}, J e_{l}, J e_{k}, e_{l}\right)+T\left(e_{k}, J e_{k}, J e_{l}, e_{l}\right)\right)
$$

for $\left\{e_{k}\right\}$ an orthonormal basis of $T M$. It is then easy to check that if $T$ is moreover of Kähler type, this last formula reduces to $\left\langle I_{\mathbb{C}}, T\right\rangle=2 \operatorname{Scal}(T)$, hence the result.

Let us recall what the complexified sectional curvature of a Hermitian manifold $(N, g, J)$ is: if $E$ and $F$ are two vectors of the complexified tangent space $T^{\mathbb{C}} N=T N \otimes_{\mathbb{R}} \mathbb{C}$ of $N$ then the complexified sectional curvature of the 2-plane they span is defined to be $R^{N}(E, F, \bar{E}, \bar{F})$ where $R^{N}$ is the curvature tensor of $g$ extended by $\mathbb{C}$-linearity to $T^{\mathbb{C}} N$. Despite its name, the complexified sectional curvature takes real values.

If $T$ is a tensor of curvature type, we define its complexified scalar curvature $\operatorname{Scal}_{\mathbb{C}}(T)$ as follows: $\operatorname{Scal}_{\mathbb{C}}(T):=\sum_{k, l=1}^{m} T\left(\zeta_{k}, \zeta_{l}, \bar{\zeta}_{k}, \bar{\zeta}_{l}\right)$, for $\left\{\zeta_{k}\right\}$ an orthonormal basis of the (1,0)-part of $T^{\mathbb{C}} \mathbb{H}_{\mathbb{C}}^{m}$.

Using the formulae given in the proof of the previous lemma, one gets
Lemma 2.6. $\left\langle I_{\mathbb{C}}-I, T\right\rangle=-6 \operatorname{Scal}_{\mathbb{C}}(T)$.
We are now ready to prove Theorem 2.1. Recall that $Q=I_{\mathbb{C}}-I$. First, Lemma 2.5 implies that the right-hand side in the Bochner-type formula $(\diamond)$ vanishes. Next, it follows from Lemma 2.4 and Lemma 2.6 that
$\int_{M} \eta_{R}\left[\langle\stackrel{\circ}{Q} \nabla \mathrm{~d} f, \nabla \mathrm{~d} f\rangle-\frac{1}{2}\left\langle Q, f^{\star} R^{n}\right\rangle\right] d V_{m}=-\frac{3}{2} \int_{M} \eta_{R}\left[\left\|(\nabla \mathrm{~d} f)^{(1,1)}\right\|^{2}-2 \operatorname{Scal}_{\mathbb{C}}\left(f^{\star} R^{n}\right)\right] d V_{m}$
for any $R>1$. Thus, formula $(\diamond)$ reads:

$$
\lim _{R \rightarrow \infty} \int_{M} \eta_{R}\left[\left\|(\nabla \mathrm{~d} f)^{(1,1)}\right\|^{2}-2 \operatorname{Scal}_{\mathbb{C}}\left(f^{\star} R^{n}\right)\right] d V_{m}=0
$$

It is known that, since the sectional curvature of $\left(\mathbb{H}_{\mathbb{C}}^{n}, g_{n}\right)$ is pinched between -1 and $-\frac{1}{4}$, its complexified sectional curvature is non-positive (see for example Her91). Therefore, $\operatorname{Scal}_{\mathbb{C}}\left(f^{\star} R^{n}\right)$ being a mean of complexified sectional curvatures of $\mathbb{H}_{\mathbb{C}}^{n}$, it is non-positive. Thus $R \longmapsto \int_{M} \eta_{R}\left[\left\|(\nabla \mathrm{~d} f)^{(1,1)}\right\|^{2}-2 \operatorname{Scal}_{\mathbb{C}}\left(f^{\star} R^{n}\right)\right] d V_{m}$ is a non-negative non-decreasing function whose limit as $R$ goes to infinity is zero. It follows at once that $(\nabla \mathrm{d} f)^{(1,1)}$ vanishes identically, that is, $f$ is pluriharmonic. Finally, we also have $\operatorname{Scal}_{\mathbb{C}}\left(f^{\star} R^{n}\right)=0$ everywhere and this implies that $R^{n}\left(f_{\star} \zeta_{k}, f_{\star} \zeta_{l}, \overline{f_{\star} \zeta_{k}}, \overline{f_{\star} \zeta_{l}}\right)=0$ for all $k, l=1, \ldots, n$. Theorem 2.1 is proved.

### 2.2. Holomorphicity of "high" rank harmonic maps.

Let $f: \mathbb{H}_{\mathbb{C}}^{m} \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$ be a finite energy harmonic map, equivariant w.r.t a representation of a torsion-free lattice $\Gamma<\operatorname{PU}(m, 1)$ into $\mathrm{PU}(n, 1)$. In this section, we exploit the full strength of Theorem 2.1 to prove a result that was first obtained by J. A. Carlson and D. Toledo in CT89] in the case $\Gamma$ is cocompact and $\rho(\Gamma)$ is discrete in $\mathrm{PU}(n, 1)$ or in a more general target Lie group $G$. Their proof relies on a careful study of maximal abelian subalgebras of the complexification of the Lie algebra of $G$. In our setting, the simple form of the curvature tensor of $\left(\mathbb{H}_{\mathbb{C}}^{n}, g_{n}\right)$ allows a more elementary proof that we give for completeness.

Proposition 2.7. Let $f: \mathbb{H}_{\mathbb{C}}^{m} \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$ be a finite energy harmonic map equivariant w.r.t. a representation $\rho$ of the torsion-free lattice $\Gamma<\operatorname{PU}(m, 1)$ in $\mathrm{PU}(n, 1)$. If the real rank of $f$ is at least 3 at some point, then $f$ is holomorphic or anti-holomorphic.

Before proving this proposition, we introduce some notations that will be needed in the proof and later on in the paper.

For $l=m, n$, let $T^{\mathbb{C}} \mathbb{H}_{\mathbb{C}}^{l}=T \mathbb{H}_{\mathbb{C}}^{l} \otimes_{\mathbb{R}} \mathbb{C}$ be the complexification of $T \mathbb{H}_{\mathbb{C}}^{l}$ and $T^{\mathbb{C}} \mathbb{H}_{\mathbb{C}}^{l}=T^{1,0} \mathbb{H}_{\mathbb{C}}^{l} \oplus$ $T^{0,1} \mathbb{H}_{\mathbb{C}}^{l}$ be its decomposition in $(1,0)$ and $(0,1)$ part. We extend the differential of $f$ by $\mathbb{C}$ linearity and still write $\mathrm{d} f: T^{\mathbb{C}} \mathbb{H}_{\mathbb{C}}^{m} \longrightarrow T^{\mathbb{C}} \mathbb{H}_{\mathbb{C}}^{n}$ (if the distinction is necessary we will use $\mathrm{d}^{\mathbb{C}} f$ ). Its components are

$$
\begin{aligned}
& \partial^{1,0} f: T^{\mathbb{C}_{\mathbb{H}}^{\mathbb{C}}} m \longrightarrow T^{1,0} \mathbb{H}_{\mathbb{C}}^{n} \\
& \partial^{0,1} f: T^{\mathbb{C}_{\mathbb{C}}^{\mathbb{C}}} \longrightarrow T^{0,1} \mathbb{H}_{\mathbb{C}}^{n}
\end{aligned}
$$

We extend $g_{n}$ by $\mathbb{C}$-linearity to $T^{\mathbb{C}} \mathbb{H}_{\mathbb{C}}^{n}$. We will sometimes write $(X, Y)=g_{n}(X, Y)$ and $|X|^{2}=(X, \bar{X})$ for $X, Y \in T^{\mathbb{C}} \mathbb{H}_{\mathbb{C}}^{n}$.

If $\left(e_{i}\right)_{1 \leq i \leq 2 m}=\left(e_{\alpha}, J e_{\alpha}\right)_{1 \leq \alpha \leq m}$ is an orthonormal $\mathbb{R}$-basis of $T \mathbb{H}_{\mathbb{C}}^{m}$, we set $z_{\alpha}=\frac{1}{2}\left(e_{\alpha}-i J e_{\alpha}\right)$. $\left(z_{\alpha}\right)_{1 \leq \alpha \leq m}$ is an orthogonal $\mathbb{C}$-basis of $T^{1,0} \mathbb{H}_{\mathbb{C}}^{m}$.

To lighten the notations, we will sometimes use $f_{\alpha}$ instead of $\partial^{1,0} f\left(z_{\alpha}\right)$ and $f_{\bar{\beta}}$ instead of $\partial^{1,0} f\left(\bar{z}_{\beta}\right)$, so that $\partial^{0,1} f\left(\bar{z}_{\alpha}\right)=\bar{f}_{\alpha}$ and $\partial^{0,1} f\left(z_{\beta}\right)=\bar{f}_{\bar{\beta}}$.

In the sequel, we will often restrict $\partial^{1,0} f$, resp. $\partial^{0,1} f$, to $T^{1,0} \mathbb{H}_{\mathbb{C}}^{m}$ and consider them as sections of $\operatorname{Hom}\left(T^{1,0} \mathbb{H}_{\mathbb{C}}^{m}, f^{\star} T^{1,0} \mathbb{H}_{\mathbb{C}}^{n}\right)$, resp. $\operatorname{Hom}\left(T^{1,0} \mathbb{H}_{\mathbb{C}}^{m}, f^{\star} T^{0,1} \mathbb{H}_{\mathbb{C}}^{n}\right)$. We will call $e^{\prime}(f)$, resp. $e^{\prime \prime}(f)$, the square of the norm of $\partial^{1,0} f$, resp. $\partial^{0,1} f$, namely:

$$
\begin{aligned}
e^{\prime}(f) & :=\left\|\partial^{1,0} f\right\|^{2}=2 \sum_{\alpha=1}^{m} g_{n}\left(\partial^{1,0} f\left(z_{\alpha}\right), \overline{\partial^{1,0} f\left(z_{\alpha}\right)}\right)=2 \sum_{\alpha=1}^{m}\left|f_{\alpha}\right|^{2}, \\
e^{\prime \prime}(f) & :=\left\|\partial^{0,1} f\right\|^{2}=2 \sum_{\alpha=1}^{m} g_{n}\left(\partial^{0,1} f\left(z_{\alpha}\right), \overline{\partial^{0,1} f\left(z_{\alpha}\right)}\right)=2 \sum_{\alpha=1}^{m}\left|f_{\bar{\alpha}}\right|^{2} .
\end{aligned}
$$

Note that with these definitions, the energy density of $f$ is given by $e(f):=\frac{1}{2}\|\mathrm{~d} f\|^{2}=$ $e^{\prime}(f)+e^{\prime \prime}(f)$. Again, we will often abreviate $e^{\prime}(f)$ and $e^{\prime \prime}(f)$ to $e^{\prime}$ and $e^{\prime \prime}$ when no confusion is possible.

Proof of Proposition 2.7. Theorem 2.1 shows that $f^{\star} R^{n}\left(z_{\alpha}, z_{\beta}, \bar{z}_{\alpha}, \bar{z}_{\beta}\right)=0$ for all $\alpha, \beta \in$ $\{1, \ldots, m\}$. Let us see what this implies in more details.

$$
\begin{aligned}
f^{\star} R^{n}\left(z_{\alpha}, z_{\beta}, \bar{z}_{\alpha}, \bar{z}_{\beta}\right)= & R^{n}\left(f_{\alpha}+\bar{f}_{\bar{\alpha}}, f_{\beta}+\bar{f}_{\bar{\beta}}, \bar{f}_{\alpha}+f_{\bar{\alpha}}, \bar{f}_{\beta}+f_{\bar{\beta}}\right) \\
= & -\frac{1}{2}\left[\left|f_{\alpha}\right|^{2}\left|f_{\bar{\beta}}\right|^{2}+\left|f_{\bar{\alpha}}\right|^{2}\left|f_{\beta}\right|^{2}-\left(f_{\bar{\alpha}}, \bar{f}_{\bar{\beta}}\right)\left(f_{\alpha}, \bar{f}_{\beta}\right)\right. \\
& -\left(f_{\beta}, \bar{f}_{\alpha}\right)\left(f_{\bar{\beta}}, \bar{f}_{\bar{\alpha}}\right)-\left(f_{\alpha}, \bar{f}_{\bar{\beta}}\right)\left(f_{\bar{\alpha}}, \bar{f}_{\beta}\right)-\left(f_{\beta}, \bar{f}_{\bar{\alpha}}\right)\left(f_{\bar{\beta}}, \bar{f}_{\alpha}\right) \\
& \left.+\left(f_{\beta}, \bar{f}_{\bar{\alpha}}\right)\left(f_{\bar{\alpha}}, \bar{f}_{\beta}\right)+\left(f_{\alpha}, \bar{f}_{\bar{\beta}}\right)\left(f_{\bar{\beta}}, \bar{f}_{\alpha}\right)\right] \\
= & -\frac{1}{2}\left(\left\|f_{\alpha} \wedge \bar{f}_{\bar{\beta}}-f_{\beta} \wedge \bar{f}_{\bar{\alpha}}\right\|^{2}+\left|\left(f_{\alpha}, \bar{f}_{\bar{\beta}}\right)-\left(f_{\beta}, \bar{f}_{\bar{\alpha}}\right)\right|^{2}\right)
\end{aligned}
$$

Therefore, for all $\alpha$ and $\beta$, we have $f_{\alpha} \wedge \bar{f}_{\bar{\beta}}=f_{\beta} \wedge \bar{f}_{\bar{\alpha}}$.
Suppose that the families $\left(f_{\alpha}\right)_{1 \leq \alpha \leq m}$ and $\left(f_{\bar{\beta}}\right)_{1 \leq \beta \leq m}$ are both of rank less than or equal to 1. We may assume that for all $\alpha$ there exits $\lambda_{\alpha}$ such that $f_{\alpha}=\lambda_{\alpha} f_{1}$ and that for some $k$ and for all $\beta$ there exists $\mu_{\beta}$ such that $f_{\bar{\beta}}=\mu_{\beta} f_{\bar{k}}$.

If $f_{1}=0$, then for all $\alpha, \mathrm{d}^{\mathbb{C}} f\left(z_{\alpha}\right)=\bar{f}_{\bar{\alpha}}$ and $\mathrm{d}^{\mathbb{C}} f\left(\bar{z}_{\alpha}\right)=f_{\bar{\alpha}}$. Therefore the complex rank of $\mathrm{d}^{\mathbb{C}} f$ is at most 2 , namely the real rank of $\mathrm{d} f$ is at most 2 . The same is true if $f_{\bar{k}}=0$.

If both $f_{1}$ and $f_{\bar{k}}$ are non zero, then from the fact that $f_{1} \wedge \bar{f}_{\bar{\alpha}}=f_{\alpha} \wedge \bar{f}_{\overline{1}}$ we deduce that $\overline{\mu_{\alpha}} f_{1} \wedge \bar{f}_{\bar{k}}=\lambda_{\alpha} \overline{\mu_{1}} f_{1} \wedge \bar{f}_{\bar{k}}$, that is, $\overline{\mu_{\alpha}}=\lambda_{\alpha} \overline{\mu_{1}}$. Then,

$$
\mathrm{d}^{\mathbb{C}} f\left(z_{\alpha}\right)=f_{\alpha}+\bar{f}_{\bar{\alpha}}=\lambda_{\alpha} f_{1}+\overline{\mu_{\alpha}} \bar{f}_{\bar{k}}=\lambda_{\alpha} f_{1}+\lambda_{\alpha} \overline{\mu_{1}} \bar{f}_{\bar{k}}=\lambda_{\alpha}\left(f_{1}+\bar{f}_{\overline{1}}\right)=\lambda_{\alpha} \mathrm{d}^{\mathbb{C}} f\left(z_{1}\right)
$$

Hence the family $\left(\mathrm{d}^{\mathbb{C}} f\left(z_{\alpha}\right)\right)_{1 \leq \alpha \leq m}$ has rank $\leq 1$. This also holds for the family $\left(\mathrm{d}^{\mathbb{C}} f\left(\bar{z}_{\beta}\right)\right)_{1 \leq \beta \leq m}$ and we conclude that the real rank of $f$ is again less than or equal to 2 .

In any case, we see that if the real rank of $f$ is at least 3 at some point, then the rank of one of the families $\left(f_{\alpha}\right)_{1 \leq \alpha \leq m}$ and $\left(f_{\bar{\beta}}\right)_{1 \leq \beta \leq m}$ is at least 2 at this point.

Suppose now that at some point of $\mathbb{H}_{\mathbb{C}}^{m}$, the rank of the family $\left(f_{\alpha}\right)_{1 \leq \alpha \leq m}$ is at least 2 , for example $f_{1} \wedge f_{2} \neq 0$. Then $f_{1} \wedge \bar{f}_{\overline{2}}=f_{2} \wedge \bar{f}_{\overline{1}}$ implies $f_{\overline{1}}=f_{\overline{2}}=0$. From $f_{1} \wedge \bar{f}_{\bar{\gamma}}=f_{\gamma} \wedge \bar{f}_{\overline{1}}$, we conclude that $f_{\bar{\gamma}}=0$ for all $1 \leq \gamma \leq m$, i.e. $e^{\prime \prime}=0$ at this point.

In the same way, if the rank of the family $\left(f_{\bar{\beta}}\right)_{1 \leq \beta \leq m}$ is at least 2 , then $e^{\prime}=0$.
Finally, since $f$ is pluriharmonic and the complexified sectional curvature of $\mathbb{H}_{\mathbb{C}}^{n}$ is zero on $\mathrm{d} f\left(T^{1,0} \mathbb{H}_{\mathbb{C}}^{m}\right)$ (see Theorem 2.1 above), and because $\mathbb{H}_{\mathbb{C}}^{n}$ is a Kähler symmetric space, it is known that $\partial^{1,0} f$, resp. $\partial^{0,1} f$, are holomorphic sections of the holomorphic bundles $\operatorname{Hom}\left(T^{1,0} \mathbb{H}_{\mathbb{C}}^{m}, f^{\star} T^{1,0} \mathbb{H}_{\mathbb{C}}^{n}\right)$, resp. $\operatorname{Hom}\left(T^{1,0} \mathbb{H}_{\mathbb{C}}^{m}, f^{\star} T^{0,1} \mathbb{H}_{\mathbb{C}}^{n}\right)(\mathbb{C} 89]$, Theorem 2.3). So they have a generic rank on $\mathbb{H}_{\mathbb{C}}^{m}$. Therefore, if for example the family $\left(f_{\alpha}\right)_{1 \leq \alpha \leq m}$ has rank at least 2 at some point, it has rank at least 2 on a dense open subset of $\mathbb{H}_{\mathbb{C}}^{m}$ and so $e^{\prime \prime}=0$ on a dense open subset of $\mathbb{H}_{\mathbb{C}}^{m}$, hence everywhere, and $f$ is holomorphic. Similarly, if $\operatorname{rk}\left(f_{\bar{\beta}}, 1 \leq \beta \leq m\right) \geq 2$ at some point, $f$ is antiholomorphic.

### 2.3. Some technical lemmas.

If $f: \mathbb{H}_{\mathbb{C}}^{m} \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$ is a $\rho$-equivariant pluriharmonic map whose rank is at most 2 everywhere, $f$ needs not be holomorphic nor antiholomorphic. Nevertheless, pluriharmonicity has other consequences that will be useful later. Namely, if $e^{\prime}(f)$ and $e^{\prime \prime}(f)$ are the previously defined squared norms of $\partial^{1,0} f$ and $\partial^{0,1} f$, we have

$$
\left\langle f^{\star} \omega_{n}, \omega_{m}\right\rangle:=\sum_{i, j=1}^{2 m} f^{\star} \omega_{n}\left(e_{i}, e_{j}\right) \omega_{m}\left(e_{i}, e_{j}\right)=2\left(e^{\prime}(f)-e^{\prime \prime}(f)\right)
$$

as it is easy to check. Because of this, some results on the energies $e^{\prime}(f)$ and $e^{\prime \prime}(f)$ will be needed in the proof of Theorem 3.1 and we shall prove them in this section.

The results stated here were obtained in complex dimension 1 in To79 and Wo79. The proofs in the general case (see also [Lic70]) go almost exactly as in the case $m=1$ and they are given only for completeness and to fix the notations.

We will work on the complexifications of the tangent spaces of $\mathbb{H}_{\mathbb{C}}^{m}$ and $\mathbb{H}_{\mathbb{C}}^{n}$ and therefore we extend all needed sections, tensors and operators defined on the real tangent spaces by $\mathbb{C}$-linearity to these complexifications.

Since $\partial^{1,0} f$ can be considered as a section of $\operatorname{Hom}\left(T^{\mathbb{C}} \mathbb{H}_{\mathbb{C}}^{m}, f^{\star} T^{\mathbb{C}} \mathbb{H}_{\mathbb{C}}^{n}\right)$, we can define its covariant derivative $\nabla \partial^{1,0} f \in \operatorname{Hom}\left(T^{\mathbb{C}} \mathbb{H}_{\mathbb{C}}^{m} \otimes T^{\mathbb{C}} \mathbb{H}_{\mathbb{C}}^{m}, f^{\star} T^{\mathbb{C}} \mathbb{H}_{\mathbb{C}}^{n}\right)$. It follows easily from the fact that $\left(\mathbb{H}_{\mathbb{C}}^{n}, g_{n}\right)$ is Kähler that $\nabla \partial^{1,0} f$ belongs in fact to $\operatorname{Hom}\left(T^{\mathbb{C}} \mathbb{H}_{\mathbb{C}}^{m} \otimes T^{\mathbb{C}} \mathbb{H}_{\mathbb{C}}^{m}, f^{\star} T^{1,0} \mathbb{H}_{\mathbb{C}}^{n}\right)$. We will call $\nabla^{\prime} \partial^{1,0} f \in \operatorname{Hom}\left(T^{1,0} \mathbb{H}_{\mathbb{C}}^{m} \otimes T^{1,0} \mathbb{H}_{\mathbb{C}}^{m}, f^{\star} T^{1,0} \mathbb{H}_{\mathbb{C}}^{n}\right)$ its restriction to $T^{1,0} \mathbb{H}_{\mathbb{C}}^{m} \otimes T^{1,0} \mathbb{H}_{\mathbb{C}}^{m}$. We define $\nabla^{\prime} \partial^{0,1} f \in \operatorname{Hom}\left(T^{1,0} \mathbb{H}_{\mathbb{C}}^{m} \otimes T^{1,0} \mathbb{H}_{\mathbb{C}}^{m}, f^{\star} T^{0,1} \mathbb{H}_{\mathbb{C}}^{n}\right)$ and $\nabla^{\prime} \mathrm{d} f=\nabla^{\prime} \partial^{1,0} f+\nabla^{\prime} \partial^{0,1} f \in$ $\operatorname{Hom}\left(T^{1,0} \mathbb{H}_{\mathbb{C}}^{m} \otimes T^{1,0} \mathbb{H}_{\mathbb{C}}^{m}, f^{\star} T^{\mathbb{C}} \mathbb{H}_{\mathbb{C}}^{n}\right)$ similarly. Note that $\left\|\nabla^{\prime} \mathrm{d} f\right\|,\left\|\nabla^{\prime} \partial^{1,0} f\right\|$ and $\left\|\nabla^{\prime} \partial^{0,1} f\right\|$ belong to $L^{2}(M)$ because $\|\nabla \mathrm{d} f\|$ does (Lemma 2.3).

In the entire section, $f$ denotes a $\rho$-equivariant pluriharmonic map $\mathbb{H}_{\mathbb{C}}^{m} \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$.
Lemma 2.8. We have

$$
\frac{1}{4} \Delta e^{\prime}=-\frac{1}{2}\left\|\nabla^{\prime} \partial^{1,0} f\right\|^{2}-2 R^{\prime}+\frac{m+1}{4} e^{\prime} \quad \text { and } \quad \frac{1}{4} \Delta e^{\prime \prime}=-\frac{1}{2}\left\|\nabla^{\prime} \partial^{0,1} f\right\|^{2}-2 R^{\prime \prime}+\frac{m+1}{4} e^{\prime \prime}
$$

where

$$
\begin{aligned}
R^{\prime} & =\sum_{\alpha, \beta} R^{n}\left(\mathrm{~d} f\left(z_{\alpha}\right), \mathrm{d} f\left(\bar{z}_{\alpha}\right), \partial^{1,0} f\left(z_{\beta}\right), \overline{\partial^{1,0} f\left(z_{\beta}\right)}\right), \\
R^{\prime \prime} & =\sum_{\alpha, \beta} R^{n}\left(\mathrm{~d} f\left(z_{\alpha}\right), \mathrm{d} f\left(\bar{z}_{\alpha}\right), \partial^{0,1} f\left(z_{\beta}\right), \overline{\partial^{0,1} f\left(z_{\beta}\right)}\right) .
\end{aligned}
$$

Proof. We make the computation for $\Delta e^{\prime}$, and we use normal coordinates:

$$
\frac{1}{4} \Delta e^{\prime}=-\sum_{\alpha} \nabla \mathrm{d} e^{\prime}\left(\bar{z}_{\alpha}, z_{\alpha}\right)=-\sum_{\alpha} \bar{z}_{\alpha} \cdot z_{\alpha} \cdot e^{\prime} .
$$

Now, $z_{\alpha} \cdot\left\langle\partial^{1,0} f, \overline{\partial^{1,0} f}\right\rangle=\left\langle\nabla_{z_{\alpha}} \partial^{1,0} f, \overline{\partial^{1,0} f}\right\rangle+\left\langle\partial^{1,0} f, \nabla_{z_{\alpha}} \overline{\partial^{1,0} f}\right\rangle$. The map $f$ is pluriharmonic and therefore $\nabla \mathrm{d} f(Z, \bar{W})=\nabla \mathrm{d} f(\bar{Z}, W)=0$ for all $Z, W$ in $T^{1,0} \mathbb{H}_{\mathbb{C}}^{m}$. Since $\mathbb{H}_{\mathbb{C}}^{m}$ is Kähler, $\nabla_{\bar{z}_{\alpha}} \partial^{1,0} f=\left(\nabla_{\bar{z}_{\alpha}} \mathrm{d} f\right)^{1,0}$ and hence vanishes identically on $T^{1,0} \mathbb{H}_{\mathbb{C}}^{m}$. It follows that $\nabla_{z_{\alpha}} \overline{\partial^{1,0} f}=$ $\overline{\nabla_{\bar{z}_{\alpha}} \partial^{1,0} f}=0$. Thus,

$$
\begin{aligned}
\frac{1}{4} \Delta e^{\prime} & =-\sum_{\alpha} \bar{z}_{\alpha} \cdot\left\langle\nabla_{z_{\alpha}} \partial^{1,0} f, \overline{\partial^{1,0} f}\right\rangle \\
& =-\sum_{\alpha}^{\alpha}\left[\left\langle\nabla_{\bar{z}_{\alpha}} \nabla_{z_{\alpha}} \partial^{1,0} f, \overline{\partial^{1,0} f}\right\rangle+\left\langle\nabla_{z_{\alpha}} \partial^{1,0} f, \overline{\nabla_{z_{\alpha}} \partial^{1,0} f}\right\rangle\right] \\
& =-\frac{1}{2}\left\|\nabla^{\prime} \partial^{1,0} f\right\|^{2}-\sum_{\alpha}\left\langle\nabla_{\bar{z}_{\alpha}} \nabla_{z_{\alpha}} \partial^{1,0} f, \overline{\partial^{1,0} f}\right\rangle .
\end{aligned}
$$

Therefore, since $\left(\nabla_{\bar{z}_{\alpha}} \nabla_{z_{\alpha}} \partial^{1,0} f\right)\left(z_{\beta}\right)=\left(\nabla_{\bar{z}_{\alpha}} \nabla_{z_{\alpha}} \partial^{1,0} f\right)\left(z_{\beta}\right)-\left(\nabla_{z_{\alpha}} \nabla_{\bar{z}_{\alpha}} \partial^{1,0} f\right)\left(z_{\beta}\right)$, we have

$$
\begin{aligned}
\sum_{\alpha}\left(\nabla_{\bar{z}_{\alpha}} \nabla_{z_{\alpha}} \partial^{1,0} f\right)\left(z_{\beta}\right) & =\sum_{\alpha} R^{n}\left(\mathrm{~d} f\left(z_{\alpha}\right), \mathrm{d} f\left(\bar{z}_{\alpha}\right)\right) \partial^{1,0} f\left(z_{\beta}\right)-\sum_{\alpha} \partial^{1,0} f\left(R^{m}\left(z_{\alpha}, \bar{z}_{\alpha}\right) z_{\beta}\right) \\
& =\sum_{\alpha} R^{n}\left(\mathrm{~d} f\left(z_{\alpha}\right), \mathrm{d} f\left(\bar{z}_{\alpha}\right)\right) \partial^{1,0} f\left(z_{\beta}\right)+\frac{1}{2} \partial^{1,0} f\left(\operatorname{Ric}^{m}\left(z_{\beta}\right)\right) .
\end{aligned}
$$

The result follows since the Ricci curvature tensor of $g_{m}$ is $-\frac{m+1}{2} g_{m}$.
We also have
Lemma 2.9. At each point of $\mathbb{H}_{\mathbb{C}}^{m}$ where $e^{\prime} \neq 0$, resp. $e^{\prime \prime} \neq 0$,

$$
\frac{1}{4} \Delta \log e^{\prime}=-\alpha^{\prime}-2 \frac{R^{\prime}}{e^{\prime}}+\frac{m+1}{4}, \text { resp. } \frac{1}{4} \Delta \log e^{\prime \prime}=-\alpha^{\prime \prime}-2 \frac{R^{\prime \prime}}{e^{\prime \prime}}+\frac{m+1}{4},
$$

where

$$
\alpha^{\prime}=\frac{1}{2 e^{\prime 2}}\left(\left\|\nabla^{\prime} \partial^{1,0} f\right\|^{2} e^{\prime}-\|\left\langle\nabla^{\prime} \partial^{1,0} f, \overline{\left.\partial^{1,0} f\right\rangle} \|^{2}\right)\right.
$$

and

$$
\alpha^{\prime \prime}=\frac{1}{2 e^{\prime \prime 2}}\left(\left\|\nabla^{\prime} \partial^{0,1} f\right\|^{2} e^{\prime \prime}-\|\left\langle\nabla^{\prime} \partial^{0,1} f, \overline{\left.\left.\partial^{0,1} f\right\rangle \|^{2}\right)}\right.\right.
$$

are both nonnegative by Cauchy-Schwarz's inequality.
Proof. Again, we make this (easy) computation only for $\frac{1}{4} \Delta \log e^{\prime}=\frac{1}{4 e^{\prime}} \Delta e^{\prime}+\frac{1}{4 e^{\prime 2}}\left\|\mathrm{~d} e^{\prime}\right\|^{2}$. Now, $\frac{1}{4}\left\|\mathrm{~d} e^{\prime}\right\|^{2}=\sum_{\alpha} \mathrm{d} e^{\prime}\left(z_{\alpha}\right) \mathrm{d} e^{\prime}\left(\bar{z}_{\alpha}\right)=\sum_{\alpha}\left\langle\nabla_{z_{\alpha}} \partial^{1,0} f, \overline{\partial^{1,0} f}\right\rangle\left\langle\partial^{1,0} f, \nabla_{\bar{z}_{\alpha}} \overline{\partial^{1,0} f}\right\rangle=\frac{1}{2}\left\|\left\langle\nabla^{\prime} \partial^{1,0} f, \overline{\partial^{1,0} f}\right\rangle\right\|^{2}$ where $\left\langle\nabla^{\prime} \partial^{1,0} f, \overline{\partial^{1,0} f}\right\rangle$ denotes the 1-form on $T^{1,0} \mathbb{H}_{\mathbb{C}}^{m}$ given by $\left\langle\nabla^{\prime} \partial^{1,0} f, \overline{\partial^{1,0} f}\right\rangle(z)=$ $\left\langle\nabla_{z} \partial^{1,0} f, \overline{\partial^{1,0} f}\right\rangle=2 \sum_{\beta} g_{n}\left(\left(\nabla_{z} \partial^{1,0} f\right)\left(z_{\beta}\right), \overline{\partial^{1,0} f}\left(z_{\beta}\right)\right)$.

Finally, easy computations show that:

## Lemma 2.10.

$$
\begin{aligned}
R^{\prime} & =\frac{1}{2} \sum_{\alpha, \beta}\left[\left|g_{n}\left(\partial^{1,0} f\left(z_{\alpha}\right), \overline{\partial^{1,0} f\left(z_{\beta}\right)}\right)\right|^{2}-\left|g_{n}\left(\partial^{1,0} f\left(z_{\beta}\right), \overline{\partial^{1,0} f\left(\bar{z}_{\alpha}\right)}\right)\right|^{2}\right]+\frac{1}{8} e^{\prime}\left(e^{\prime}-e^{\prime \prime}\right) \\
R^{\prime \prime} & =\frac{1}{2} \sum_{\alpha, \beta}\left[\left|g_{n}\left(\partial^{1,0} f\left(\bar{z}_{\alpha}\right), \overline{\partial^{1,0} f\left(\bar{z}_{\beta}\right)}\right)\right|^{2}-\left|g_{n}\left(\partial^{1,0} f\left(z_{\alpha}\right), \overline{\partial^{1,0} f\left(\bar{z}_{\beta}\right)}\right)\right|^{2}\right]+\frac{1}{8} e^{\prime \prime}\left(e^{\prime \prime}-e^{\prime}\right)
\end{aligned}
$$

Remark. In the sequel, we shall use the fact that all the functions involved in those three lemmas are well defined on $M=\Gamma \backslash \mathbb{H}_{\mathbb{C}}^{m}$.

## 3. Rigidity of Representations of lattices of $\mathrm{PU}(m, 1)$ into $\mathrm{PU}(n, 1)$

### 3.1. Burger-Iozzi invariant.

We again assume that $m \geq 2$.
Let $\Gamma$ be a torsion-free lattice in $\mathrm{PU}(m, 1)$, and let $\rho: \Gamma \longrightarrow \mathrm{PU}(n, 1)$ be a homomorphism. M. Burger and A. Iozzi assign to $\rho$ an invariant which can be defined as follows (see [BI01]).

Take any $\rho$-equivariant map $f: \mathbb{H}_{\mathbb{C}}^{m} \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$ and consider the pull-back $f^{\star} \omega_{n}$ of the Kähler form $\omega_{n}$ of $\mathbb{H}_{\mathbb{C}}^{n}$. Note that we can consider $f^{\star} \omega_{n}$ as a 2 -form on $M=\Gamma \backslash \mathbb{H}_{\mathbb{C}}^{m}$. The de Rham cohomology class $\left[f^{\star} \omega_{n}\right] \in H_{D R}^{2}(M)$ defined by $f^{\star} \omega_{n}$ is independent of the choice of the equivariant map $f$ since all such maps are homotopic, and therefore we call it $\left[\rho^{\star} \omega_{n}\right]$.

Now, Burger and Iozzi remark that the class $\left[\rho^{\star} \omega_{n}\right]$ is in the image of the natural comparison map from the $L^{2}$-cohomology group $H_{(2)}^{2}(M)$ of $M$ to the de Rham cohomology group $H_{D R}^{2}(M)$. Since $m \geq 2$, the comparison map is injective (see [Zu82]; the arithmeticity of the lattice $\Gamma$ is not necessary for the result in the present case), this yields a well-defined $L^{2}$-cohomology class, denoted by $\left[\rho^{\star} \omega_{n}\right]_{(2)}$, and they define (in a slightly different form)

$$
\tau(\rho):=\frac{1}{2 m} \int_{M}\left\langle\rho^{\star} \omega_{n}, \omega_{m}\right\rangle d V_{m}
$$

where $\rho^{\star} \omega_{n}$ is any $L^{2}$-form representing $\left[\rho^{\star} \omega_{n}\right]_{(2)}$ (observe that, because $\omega_{m}$ is parallel, $\tau(\rho)$ depends only on $\left[\rho^{\star} \omega_{n}\right]_{(2)}$, hence on $\rho$ ).
Remark. In complex dimension 1 and for a uniform lattice $\Gamma, \tau(\rho)$ is also well-defined and coincides with the classical Toledo invariant (cf. [To89]).

The main result of [BI01] then reads:

Theorem 3.1. Under the above assumptions,

$$
|\tau(\rho)| \leq \operatorname{Vol}(M) .
$$

Moreover, equality holds if and only if there exists a $\rho$-equivariant totally geodesic isometric embedding $\mathbb{H}_{\mathbb{C}}^{m} \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$.

Let $\Gamma$ be a torsion-free lattice in $\operatorname{SU}(m, 1)$. Via the natural inclusion of $\mathrm{SU}(m, 1)$ into $\operatorname{PU}(n, 1)$ for $n>m$, we obtain the so-called standard representation of $\Gamma$ into $\operatorname{PU}(n, 1)$ (which is of course $\mathbb{C}$-Fuchsian). Theorem 3.1 then implies:

Corollary 3.2. Let $\Gamma$ be a torsion-free lattice in $\operatorname{SU}(m, 1), m \geq 2$. Then any deformation of the standard representation of $\Gamma$ into $\mathrm{PU}(n, 1)(n>m)$ is also $\mathbb{C}$-Fuchsian.

Proof. Since $\Gamma$ is torsion-free, it projects isomorphically into $\operatorname{PU}(m, 1)$. We can therefore consider the standard representation of $\Gamma$ as a representation of a lattice of $\mathrm{PU}(m, 1)$ and apply Theorem 3.1. Now, when seen as a cohomology class in $H_{D R}^{2}(M),\left[\rho^{\star} \omega_{n}\right]$ is a characteristic class of the principal $\mathrm{PU}(n, 1)$-bundle over $M$ associated to $\rho$ and so, it is constant on connected components of $\operatorname{Hom}(\Gamma, \operatorname{PU}(n, 1))$. On the other hand, $|\tau(\rho)|=\operatorname{Vol}(M)$ holds when $\rho$ is the standard representation of $\Gamma$ into $\operatorname{PU}(n, 1)$, hence the result.

The main tool in To89 and [BI01 is bounded cohomology. Corlette in Co88 was the first to obtain Corollary 3.2 for $m \geq 2$ and $\Gamma$ cocompact. He worked with an invariant similar to $\tau(\rho)$, the volume of $\rho$ (see the introduction and the remark at the end of Section 3.2), and he used harmonic maps techniques to obtain a result equivalent to Theorem 3.1.

We will now show that Theorem 3.1 is a consequence of our results on harmonic maps.

### 3.2. Proof of Theorem 3.1.

The next two lemmas have important consequences. They show that the invariant is welldefined, that is, that the de Rham cohomology class $\left[\rho^{\star} \omega_{n}\right]$ can be represented by an $L^{2}$-form. Moreover, they imply that we don't have to worry about non reductive representations and that we can use the finite energy harmonic map we constructed to compute the invariant in the case of reductive ones.

Lemma 3.3. Suppose the representation $\rho$ is not reductive. Then $\left[\rho^{\star} \omega_{n}\right]=0$ and hence $\tau(\rho)=0$.

Proof. Let $f$ be any $\rho$-equivariant map from $\mathbb{H}_{\mathbb{C}}^{m}$ to $\mathbb{H}_{\mathbb{C}}^{n}$. The image $\rho(\Gamma)$ fixes a point at infinity in $\mathbb{H}_{\mathbb{C}}^{n}$ which we can take to be $\infty$ in the Siegel model. We then have $\omega_{n}=-d^{c} t^{\prime}$ where $t^{\prime}=\log \left(2 \operatorname{Re}\left(w^{\prime}\right)-\left\langle\left\langle z^{\prime}, z^{\prime}\right\rangle\right\rangle\right)$ (see section 1$)$. As can be easily checked, the 1 -form $\varsigma:=-\mathrm{d}^{c} t^{\prime}$ is invariant by the stabilizer in $\mathrm{PU}(n, 1)$ of the fixed point. Therefore, the form $f^{\star} \varsigma$ on $\mathbb{H}_{\mathbb{C}}^{m}$ goes down to a form defined on the whole quotient $M$ such that $f^{\star} \omega_{n}=\mathrm{d} f^{\star} \varsigma$. Hence the de Rham cohomology class of $f^{\star} \omega_{n}$ is zero and the invariant $\tau(\rho)$ vanishes.

We may now assume that $\rho$ is reductive. Theorems 1.1 and 1.2 then guarantee the existence of a finite energy $\rho$-equivariant harmonic map $f: \mathbb{H}_{\mathbb{C}}^{m} \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$.

We first prove that we can use $f$ to compute $\tau(\rho)$. This follows from the
Lemma 3.4. The 2 -form $f^{\star} \omega_{n}$ is $L^{2}$.
Proof. If $f$ is (anti)holomorphic, it follows from the generalization of the Schwarz-Pick lemma (see Ko70) that in an obvious sense, $f^{\star} g_{n} \leq g_{m}$. An easy computation then shows that $\left\|f^{\star} \omega_{n}\right\| \leq 2 m$ everywhere on $M: f^{\star} \omega_{n}$ is a $L^{2}$-form.

Assume now that $f$ is not (anti)holomorphic. Then Proposition 2.7 implies that $\mathrm{rk}_{\mathbb{R}} f \leq 2$. When $\mathrm{rk}_{\mathbb{R}} f \leq 1, f^{\star} \omega_{n}=0$, whereas $\mathrm{rk}_{\mathbb{R}} \mathrm{d}_{x} f=2\left(\right.$ or $^{\left.\mathrm{rk}_{\mathbb{C}} \mathrm{d}_{x}^{\mathbb{C}} f=2\right) \text { is equivalent to }}$

$$
\operatorname{dim}_{\mathbb{C}} \mathrm{d}_{x}^{\mathbb{C}} f\left(T^{1,0} \mathbb{H}_{\mathbb{C}}^{m}\right)=1 \text { and } \mathrm{d}_{x}^{\mathbb{C}} f\left(T^{1,0} \mathbb{H}_{\mathbb{C}}^{m}\right) \cap \overline{\mathrm{d}_{x}^{\mathbb{C}} f\left(T^{1,0} \mathbb{H}_{\mathbb{C}}^{m}\right)}=\{0\}
$$

But, if $\mathrm{d}_{x}^{\mathbb{C}} f\left(T^{1,0} \mathbb{H}_{\mathbb{C}}^{m}\right)$ contains no real vectors, it contains no purely imaginary vectors, thus $\mathrm{d}_{x} f(X)=0$ if and only if $\mathrm{d}_{x}^{\mathbb{C}} f(X-i J X)=0$. This means that $\operatorname{Ker}_{x} f$ is $J$-invariant, hence, we may choose an orthonormal basis $\left(e_{i}\right)_{1 \leq i \leq 2 m}=\left(e_{i}, J e_{i}\right)_{1 \leq i \leq m}$ of $T_{x} \mathbb{H}_{\mathbb{C}}^{m}$ in which computations give

$$
\left\langle f^{\star} \omega_{n}, f^{\star} \omega_{n}\right\rangle=\sum_{i, j=1}^{2 m} f^{\star} \omega_{n}\left(e_{i}, e_{j}\right)^{2}=2 f^{\star} \omega_{n}\left(e_{1}, J e_{1}\right)^{2}=2\left(e^{\prime}-e^{\prime \prime}\right)^{2} .
$$

Thus, if $\mathrm{rk}_{\mathbb{R}} f=2$, we have $\left\|f^{\star} \omega_{n}\right\|^{2}=2\left(e^{\prime}-e^{\prime \prime}\right)^{2}$ on $\mathbb{H}_{\mathbb{C}}^{m}$. Using the fact that $\mathrm{rk}_{\mathbb{R}} f=2$ in the formulae of Lemma 2.10, we find that $R^{\prime} \geq \frac{1}{4} e^{\prime}\left(e^{\prime}-e^{\prime \prime}\right)$ and $R^{\prime \prime} \geq \frac{1}{4} e^{\prime \prime}\left(e^{\prime \prime}-e^{\prime}\right)$ so that $\left(e^{\prime}-e^{\prime \prime}\right)^{2} \leq 4\left(R^{\prime}+R^{\prime \prime}\right)$. Now, adding the two equalities in Lemma 2.8, we get

$$
2\left(e^{\prime}-e^{\prime \prime}\right)^{2} \leq 8\left(R^{\prime}+R^{\prime \prime}\right)=-\Delta e-2\left\|\nabla^{\prime} \mathrm{d}^{\mathbb{C}} f\right\|^{2}+(m+1) e .
$$

Since $f$ has finite energy, we conclude that $f^{\star} \omega_{n}$ is $L^{2}$ (see the proof of Lemma 2.3 and the beginning of section 2.3).

We now prove the inequality $|\tau(\rho)| \leq \operatorname{Vol}(M)$.
The above lemma implies that the Burger-Iozzi invariant of $\rho$ is given by

$$
\tau(\rho)=\frac{1}{2 m} \int_{M}\left\langle f^{\star} \omega_{n}, \omega_{m}\right\rangle d V_{m}=\frac{1}{m} \int_{M}\left(e^{\prime}-e^{\prime \prime}\right) d V_{m}
$$

since $\left\langle f^{\star} \omega_{n}, \omega_{m}\right\rangle=2\left(e^{\prime}-e^{\prime \prime}\right)$.
Therefore, if $f$ is a complex rank $r$ holomorphic map, the Schwarz-Pick lemma implies that $0 \leq\left\langle f^{\star} \omega_{n}, \omega_{m}\right\rangle \leq 2 r$ at each point of $M$, whereas if $f$ is a rank $r$ antiholomorphic map, then $-2 r \leq\left\langle f^{\star} \omega_{n}, \omega_{m}\right\rangle \leq 0$. Integrating these inequalities yields the result.

If $f$ is neither holomophic nor antiholomorphic, then we know from Proposition 2.7 that $\mathrm{rk}_{\mathbb{R}} f \leq 2$. We will prove that in this case,

$$
\left|\int_{M}\left\langle f^{\star} \omega_{n}, \omega_{m}\right\rangle d V_{m}\right| \leq(m+1) \operatorname{Vol}(M)
$$

As in the proof of Lemma 3.4. we have $R^{\prime} \geq \frac{1}{4} e^{\prime}\left(e^{\prime}-e^{\prime \prime}\right)$ and $R^{\prime \prime} \geq \frac{1}{4} e^{\prime \prime}\left(e^{\prime \prime}-e^{\prime}\right)$. Lemma 2.9 then implies that

$$
\Delta \log e^{\prime} \leq 2\left(e^{\prime \prime}-e^{\prime}\right)+m+1,
$$

resp.

$$
\Delta \log e^{\prime \prime} \leq 2\left(e^{\prime}-e^{\prime \prime}\right)+m+1,
$$

at each point where $e^{\prime} \neq 0$, resp. $e^{\prime \prime} \neq 0$.
Let $\varepsilon>0$ be a regular value of $e^{\prime}$. We set $M_{\varepsilon}=\left\{x \in M, e^{\prime}(x)>\varepsilon\right\}$, and we introduce the cut-off function $\eta_{R}$ defined in the proof of Proposition 2.2.

By Green's formula we have

$$
\int_{M_{\varepsilon}}\left[\eta_{R} \Delta \log e^{\prime}-\left\langle\operatorname{grad} \eta_{R}, \operatorname{grad} \log e^{\prime}\right\rangle\right] d V_{m}=\int_{\partial M_{\varepsilon}} \eta_{R}\left\langle\nu, \operatorname{grad} \log e^{\prime}\right\rangle d A \geq 0
$$

The latter is nonnegative because $\nu$ is the inward unit vector field along $\partial M_{\varepsilon}$ which is pointwise orthogonal to $\partial M_{\varepsilon}$. From

$$
\left|\left\langle\operatorname{grad} \eta_{R}, \operatorname{grad} \log e^{\prime}\right\rangle\right| \leq \sqrt{2}\left\|\mathrm{~d} \eta_{R}\right\| \frac{\left\|\nabla^{\prime} \partial^{1,0} f\right\|}{\sqrt{e^{\prime}}} \leq \sqrt{2}\left\|\mathrm{~d} \eta_{R}\right\| \frac{\left\|\nabla^{\prime} \partial^{1,0} f\right\|}{\sqrt{\varepsilon}} \text { on } M_{\varepsilon},
$$

we get

$$
\int_{M_{\varepsilon}} \eta_{R} \Delta \log e^{\prime} d V_{m} \geq-\sqrt{\frac{2}{\varepsilon}} \frac{C}{R} \int_{M}\left\|\nabla^{\prime} \partial^{1,0} f\right\| d V_{m} \geq-\sqrt{\frac{2}{\varepsilon}} \frac{C}{R}\left[\operatorname{Vol}(M) \int_{M}\left\|\nabla^{\prime} \partial^{1,0} f\right\|^{2} d V_{m}\right]^{1 / 2}
$$

Using the fact that $\left\|\nabla^{\prime} \partial^{1,0} f\right\| \in L^{2}(M)$, we obtain that, for some constant $C^{\prime}$ and for all $R>1$,

$$
\int_{M_{\varepsilon}} \eta_{R} \Delta \log e^{\prime} d V_{m} \geq-\sqrt{\frac{2}{\varepsilon}} \frac{C^{\prime}}{R}
$$

and so

$$
\int_{M_{\varepsilon}}\left(2\left(e^{\prime \prime}-e^{\prime}\right)+m+1\right) d V_{m}=\lim _{R \rightarrow+\infty} \int_{M_{\varepsilon}} \eta_{R}\left(2\left(e^{\prime \prime}-e^{\prime}\right)+m+1\right) d V_{m} \geq 0
$$

The subset $\left\{x \in M, e^{\prime}(x)=0\right\}$ has zero measure since $\partial^{1,0} f$ is a holomorphic section of $\operatorname{Hom}\left(T^{1,0} \mathbb{H}_{\mathbb{C}}^{m}, f^{\star} T^{1,0} \mathbb{H}_{\mathbb{C}}^{n}\right)$. Hence, letting $\varepsilon \rightarrow 0$, we conclude that

$$
\int_{M}\left\langle f^{\star} \omega_{n}, \omega_{m}\right\rangle d V_{m}=\int_{M} 2\left(e^{\prime}-e^{\prime \prime}\right) d V_{m} \leq(m+1) \operatorname{Vol}(M)
$$

Integrating $\Delta \log e^{\prime \prime}$ in the same way, we get the required inequality.
Assume now that equality holds in the theorem: $\left|\int_{M}\left\langle f^{\star} \omega_{n}, \omega_{m}\right\rangle d V_{m}\right|=2 m \operatorname{Vol}(M)$.
Since $m \geq 2$, it follows from the proof above that $f$ is a complex rank $m$ holomorphic or antiholomorphic map. Since the inequality $\left|\left\langle f^{\star} \omega_{n}, \omega_{m}\right\rangle\right| \leq 2 m$ is therefore true pointwise, the global equality implies that $\left|\left\langle f^{\star} \omega_{n}, \omega_{m}\right\rangle\right|=2 m$ everywhere on $M: f$ is an isometry. On the other hand, the Bochner-type formula $(\diamond)$ with $Q=I$ reads:

$$
\int_{M}\|\nabla \mathrm{~d} f\|^{2} d V_{m}=\frac{1}{2} \int_{M}\left(\operatorname{Scal}\left(f^{\star} R^{n}\right)-\frac{1}{2 m}\|\mathrm{~d} f\|^{2} \operatorname{Scal}\left(R^{m}\right)\right) d V_{m}
$$

Since $f$ is an (anti)holomorphic isometry, $\|\mathrm{d} f\|^{2}=2 m$ and $f^{\star} R^{n}=R^{m}$. Therefore $\nabla \mathrm{d} f=0$, namely, $f$ is totally geodesic and we are done.
Remark. It should be noted that if one is interested only in proving Corollary 3.2, it is actually possible to define another invariant, that one could call the $L^{2}$-volume of the representation, in the following way: just take the $m$-th exterior power of any $L^{2}$-form representing the $L^{2}$ cohomology class $\left[\rho^{\star} \omega_{n}\right]_{(2)}$ and integrate it over $M$. The so-obtained number is independent of the choice of the $L^{2}$-representative and therefore depends only on $\rho$. One can then prove Theorem 3.1 with $\tau(\rho)$ replaced by this (suitably normalized) $L^{2}$-volume. The proof is in fact easier since this invariant will be zero if the real rank of the harmonic map is less than 2 m . Therefore one does not need to deal with non (anti)holomorphic maps.

Nevertheless, there exist representations $\rho$ with zero $L^{2}$-volume but $\tau(\rho) \neq 0$ and one can hope to be able to get informations on these representations from the Burger-Iozzi invariant that the volume would not give. An example of such a representation was given by Livné in his thesis (Liv81], see also [Ka98]). He constructed a (closed) complex hyperbolic manifold $M$ of complex dimension 2 and a surjective holomorphic map $f$ from $M$ to a (closed) Riemann surface $\Sigma$ such that the induced map on the fundamental groups is surjective. This gives a representation of the lattice $\pi_{1}(M) \subset \mathrm{PU}(2,1)$ into $\mathrm{PU}(1,1)$. Of course the volume of this representation is zero. But its Burger-Iozzi invariant does not vanish since the pull-back $f^{\star} \omega_{1}$ is a semi-positive ( 1,1 )-form on $M$ which is not identically zero.

## 4. The case of non-uniform lattices of $\operatorname{PU}(1,1)$

In this section we want to extend the previous results to the 1-dimensional case, namely the case of non-uniform lattices of $\mathrm{PU}(1,1)$. We remark that if $\rho$ is a representation of such a lattice into $\mathrm{PU}(n, 1)$, the Burger-Iozzi invariant of $\rho$ is not defined since $H_{D R}^{2}(M)=0$ and the comparison map $H_{(2)}^{2}(M) \longrightarrow H_{D R}^{2}(M)$ is of course not injective anymore. As we mentionned in the introduction, Corollary 3.2 fails in this case. Indeed, Gusevskii and Parker prove in GP00 that there exist lattices in $\mathrm{PU}(1,1)$ admitting quasi-Fuchsian deformations into $\mathrm{PU}(2,1)$.

Geometrically, a torsion-free lattice $\Gamma<\mathrm{PU}(1,1)$ is the fundamental group of the complete hyperbolic surface of finite volume $M=\Gamma \backslash \mathbb{H}_{\mathbb{C}}^{1}$. It turns out that we don't need a Riemannian structure on $M$ to define an invariant associated to representations of its fundamental group into $\mathrm{PU}(1,1)$. We will therefore work in the more general setting of fundamental groups of orientable surfaces of finite topological type.

Let $M$ be the open surface obtained by removing a finite number of points $m_{1}, \ldots, m_{p}$, called punctures, from a closed orientable surface $\bar{M}$ of genus $g$. We will call such an $M$ a $p$ times punctured closed orientable surface of genus $g$. We assume that the Euler characteristic $\chi(M)=2-2 g-p$ of $M$ is negative. Let $\Gamma=\pi_{1}(M)$ be the fundamental group of $M$ and $\pi: \widetilde{M} \longrightarrow M$ be the universal cover of $M$.

Loops going once around the puncture $m_{i}$ in the direction prescribed by the orientation of $\bar{M}$ correspond to a conjugacy class $c_{i}$ of elements of $\Gamma$. The elements of the conjugacy classes $c_{i}$ are called peripheral. For each $i$, choose $\gamma_{i} \in c_{i}$ and denote by $\left\langle\gamma_{i}\right\rangle$ the cyclic subgroup generated by $\gamma_{i}$. There exist small disjoint open topological discs $D_{i} \subset \bar{M}$ around each $m_{i}$ and disjoint open simply-connected sets $U_{i}$ in $\widetilde{M}$, precisely invariant under $\left\langle\gamma_{i}\right\rangle$ (meaning that $\gamma_{i} U_{i}=U_{i}$ and $\gamma U_{i} \cap U_{i}=\emptyset$ if $\left.\gamma \notin\left\langle\gamma_{i}\right\rangle\right)$, such that the punctured discs $D_{i}^{\star}:=D_{i} \backslash\left\{m_{i}\right\} \subset M$ are given by $D_{i}^{\star}=\left\langle\gamma_{i}\right\rangle \backslash U_{i}$ (This can for example be seen by uniformizing $M$ as a finite volume hyperbolic surface, and then choosing precisely invariant horospherical neighbourhoods of the parabolic fixed points of $\Gamma$ ).

Let $\rho$ be a homomorphism from $\Gamma$ to $\mathrm{PU}(n, 1)$. At the beginning of section 3.2 (there, $\Gamma$ was a lattice in $\mathrm{PU}(m, 1)$ for $m \geq 2)$, we saw how to find a compactly supported 2 -form in the de Rham cohomology class $\left[\rho^{\star} \omega_{n}\right]$. Here, as we said, this class is zero but we shall in the same way associate to $\rho$ a class in the cohomology with compact support.

For each $i$, choose a fixed point $\xi_{i}$ of $\rho\left(\gamma_{i}\right)$, in $\mathbb{H}_{\mathbb{C}}^{n}$ if $\rho\left(\gamma_{i}\right)$ is elliptic, else in $\partial_{\infty} \mathbb{H}_{\mathbb{C}}^{n}$, and then a Kähler potential $\psi_{i}$ of $\omega_{n}$, invariant by the stabilizer in $\mathrm{PU}(n, 1)$ of $\xi_{i}$ if $\xi_{i} \in \mathbb{H}_{\mathbb{C}}^{n}$ or by the stabilizer of the horospheres centered at $\xi_{i}$ if $\xi_{i} \in \partial_{\infty} \mathbb{H}_{\mathbb{C}}^{n}$. If $\xi_{i} \in \mathbb{H}_{\mathbb{C}}^{n}$, we can assume that $\xi_{i}=0$ in the ball model of $\mathbb{H}_{\mathbb{C}}^{n}$ and take $\psi_{i}=\log (1-\langle\langle z, z\rangle\rangle)$ (here $\langle\langle\rangle$,$\rangle denotes the$ standard Hermitian form on $\left.\mathbb{C}^{n}\right)$. If $\xi_{i} \in \partial_{\infty} \mathbb{H}_{\mathbb{C}}^{n}$, we can take $\psi_{i}=\log \left(2 \operatorname{Re}\left(w^{\prime}\right)-\left\langle\left\langle z^{\prime}, z^{\prime}\right\rangle\right\rangle\right)$, where $\left(w^{\prime}, z^{\prime}\right) \in \mathbb{C} \times \mathbb{C}^{n-1}$ are horospherical coordinates centered at $\xi_{i}$. Note that, up to an additive constant, these potentials are unique. The potential $\psi_{i}$ is invariant by $\rho\left(\gamma_{i}\right)$ only if $\rho\left(\gamma_{i}\right)$ is elliptic or parabolic but the 1-form $\varsigma_{i}:=-\mathrm{d}^{c} \psi_{i}$, which satisfies $\mathrm{d} \varsigma_{i}=\omega_{n}$, is always invariant by $\rho\left(\gamma_{i}\right)$.

Given a $\rho$-equivariant map $f: \widetilde{M} \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$, we can pull-back the Kähler form $\omega_{n}$ and the forms $\varsigma_{i}$ to get a 2 -form $f^{\star} \omega_{n}$ invariant by $\Gamma$ and 1-forms $f^{\star} \varsigma_{i}$ invariant by $\left\langle\gamma_{i}\right\rangle$. We can therefore consider $f^{\star} \omega_{n}$ as a 2 -form on $M$ and each $f^{\star} \varsigma_{i}$ as a 1-form on the punctured disc $D_{i}^{\star}$ (by restricting it first to $U_{i}$ ). If now $\chi$ is a function identically equal to 0 outside the $D_{i}^{\star}$ 's and to 1 in small neighbourhoods of the punctures, we get a compactly supported 2-form $f^{\star} \omega_{n}-\mathrm{d}\left(\sum_{i} \chi f^{\star} \varsigma_{i}\right)$ on $M$. This yields a class $\left[f^{\star} \omega_{n}-\mathrm{d}\left(\sum_{i} \chi f^{\star} \varsigma_{i}\right)\right]_{c}$ in the second cohomology group with compact support $H_{c}^{2}(M)$.

Proposition-definition 4.1. The class $\left[f^{\star} \omega_{n}-\mathrm{d}\left(\sum_{i} \chi f^{\star} s_{i}\right)\right]_{c}$ depends only on the representation $\rho$, and will therefore be denoted by $\left[\rho^{\star} \omega_{n}\right]_{c}$. Moreover, we set

$$
\tau(\rho)=\int_{M}\left[\rho^{\star} \omega_{n}\right]_{c}
$$

Proof. This class is clearly independent of the cut-off function $\chi$, and therefore also of the choice of the punctured discs $D_{i}^{\star}$ where the $f^{\star} s_{i}$ 's are defined.

Now, let $f_{1}$ and $f_{2}$ be two $\rho$-equivariant maps $\widetilde{M} \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$. The map $f_{2}$ is homotopic to a $\rho$-equivariant map $f_{3}$ such that, when seen as sections of the $\mathbb{H}_{\mathbb{C}}^{n}$-bundle on $M$ associated to $\rho, f_{3}=f_{1}$ on the set $\{\chi<1\}$ and $f_{3}=f_{2}$ close enough to the punctures. Then there exists a compactly supported 1 -form $\alpha$ such that $f_{2}^{\star} \omega_{n}=f_{3}^{\star} \omega_{n}+\mathrm{d} \alpha$. Hence we have

$$
f_{2}^{\star} \omega_{n}-\mathrm{d}\left(\sum_{i} \chi f_{2}^{\star} \varsigma_{i}\right)=f_{3}^{\star} \omega_{n}-\mathrm{d}\left(\sum_{i} \chi f_{3}^{\star} \varsigma_{i}\right)+\mathrm{d}\left(\sum_{i} \chi\left(f_{3}^{\star} s_{i}-f_{2}^{\star} \varsigma_{i}\right)\right)+\mathrm{d} \alpha
$$

But

$$
f_{3}^{\star} \omega_{n}-\mathrm{d}\left(\sum_{i} \chi f_{3}^{\star} \varsigma_{i}\right)=\left\{\begin{array}{l}
f_{1}^{\star} \omega_{n}-\mathrm{d}\left(\sum_{i} \chi f_{1}^{\star} s_{i}\right) \text { on }\{\chi<1\} \\
0 \text { on }\{\chi=1\}
\end{array}=f_{1}^{\star} \omega_{n}-\mathrm{d}\left(\sum_{i} \chi f_{1}^{\star} \varsigma_{i}\right) \text { on } M .\right.
$$

Therefore

$$
f_{2}^{\star} \omega_{n}-\mathrm{d}\left(\sum_{i} \chi f_{2}^{\star} \varsigma_{i}\right)=f_{1}^{\star} \omega_{n}-\mathrm{d}\left(\sum_{i} \chi f_{1}^{\star} s_{i}\right)+\mathrm{d}\left[\sum_{i} \chi\left(f_{3}^{\star} s_{i}-f_{2}^{\star} s_{i}\right)+\alpha\right]
$$

From the definition of $f_{3}$, the 1 -form inside the brackets is compactly supported. Hence $\left[f^{\star} \omega_{n}-\mathrm{d}\left(\sum_{i} \chi f^{\star} s_{i}\right)\right]_{c}$ does not depend on the $\rho$-equivariant map $f$.

If $\rho\left(\gamma_{j}\right)$ is elliptic for some $j$, it might fix more than one point in $\mathbb{H}_{\mathbb{C}}^{n}$. We therefore have to check that choosing another fixed point, say $\xi_{j}^{\prime}$, to define the Kähler potential does not change the class. Let $\psi_{j}^{\prime}$ be the Kähler potential associated to $\xi_{j}^{\prime}$ and $\varsigma_{j}^{\prime}=-\mathrm{d}^{c} \psi_{j}^{\prime}$ the corresponding 1 -form. To compute our class, we can choose the $\rho$-equivariant map $f: \widetilde{M} \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$ to be constant equal to $\xi_{j}$ in $U_{j}$ so that $f^{\star} \varsigma_{j}=f^{\star} \varsigma_{j}^{\prime}=0$ on $D_{j}^{\star}$. Therefore

$$
f^{\star} \omega_{n}-\mathrm{d}\left(\sum_{i} \chi f^{\star} \varsigma_{i}\right)=f^{\star} \omega_{n}-\mathrm{d}\left(\sum_{i \neq j} \chi f^{\star} \varsigma_{i}\right),
$$

and the cohomology class is not affected.
In the same way, if $\rho\left(\gamma_{j}\right)$ is hyperbolic for some $j$, we must show that we can choose any of the two fixed points $\xi_{j}$ and $\xi_{j}^{\prime}$ of $\rho\left(\gamma_{j}\right)$ indifferently. In this case, we can arrange that the equivariant map $f$ maps $U_{j}$ to the axis of $\rho\left(\gamma_{j}\right)$. If we take the potential $\psi_{j}$ associated to $\xi_{j}$ (resp. $\xi_{j}^{\prime}$ ) to define $\varsigma_{j}$ then $\varsigma_{j}=-\mathrm{d}^{c} t=\mathrm{d} t \circ J$ in horospherical coordinates $(z, v, t)$ chosen so that $\xi_{j}=\infty$ and $\xi_{j}^{\prime}=0$ (resp. $\xi_{j}^{\prime}=\infty$ and $\xi_{j}=0$ ). Since in these coordinates the axis of $\rho\left(\gamma_{j}\right)$ is the set $\{z=0, v=0\}, f^{\star} \varsigma_{j}=0$ on $D_{j}^{\star}$. Again,

$$
f^{\star} \omega_{n}-\mathrm{d}\left(\sum_{i} \chi f^{\star} \varsigma_{i}\right)=f^{\star} \omega_{n}-\mathrm{d}\left(\sum_{i \neq j} \chi f^{\star} \varsigma_{i}\right)
$$

Finally, it is easy to check that a different choice of the peripheral elements $\gamma_{i}$ (and hence of the $U_{i}$ 's) gives the same cohomology class. Indeed, let $\gamma_{j}$ and $\gamma_{j}^{\prime}=\gamma \gamma_{j} \gamma^{-1}$ be two elements of the conjugacy class $c_{i}$. We denote with primes the objects associated to $\gamma_{j}^{\prime}$ (for example, $\psi_{j}^{\prime}$ is the Kähler potential associated to a fixed point $\xi_{j}^{\prime}$ of $\left.\rho\left(\gamma_{j}^{\prime}\right)\right)$. If $\rho\left(\gamma_{j}\right)$, and hence $\rho\left(\gamma_{j}^{\prime}\right)$, is elliptic or hyperbolic, we can choose as above the equivariant map $f$ so that $f^{\star} s_{i}=0$ on $U_{i}$ and $f^{\star} \varsigma_{i}^{\prime}=0$ on $U_{i}^{\prime}=\gamma U_{i}$. Thus we can assume that $\rho\left(\gamma_{j}\right)$ and $\rho\left(\gamma_{j}^{\prime}\right)$ are both parabolic. But
then $\xi_{j}^{\prime}=\rho(\gamma) \xi_{j}, \psi_{j}^{\prime}=\psi_{j} \circ \rho\left(\gamma^{-1}\right)$ and $f^{\star} \varsigma_{j}^{\prime}=\left(\gamma^{-1}\right)^{\star} f^{\star} \varsigma_{j}$ on $\mathbb{H}_{\mathbb{C}}^{1}$. Therefore the restrictions of $f^{\star} \varsigma_{j}^{\prime}$ to $U_{j}^{\prime}=\gamma U_{j}$ and of $f^{\star} \varsigma_{j}$ to $U_{j}$ induce the same form on $D_{j}^{\star}=\left\langle\gamma_{j}\right\rangle \backslash U_{j}=\left\langle\gamma_{j}^{\prime}\right\rangle \backslash U_{j}^{\prime}$.

Before stating the main theorem of this section, we need the following definitions.
Definition 4.2. Let $\Gamma$ be the fundamental group of a p-times punctured closed orientable surface $M$.

A homomorphism $\rho$ from $\Gamma$ in $\mathrm{PU}(n, 1)$ is called tame if it maps no peripheral element of $\Gamma$ to a hyperbolic isometry of $\mathbb{H}_{\mathbb{C}}^{n}$.

A homomorphism $\rho$ from $\Gamma$ in $\mathrm{PU}(1,1)$ is called a uniformization representation if it is an isomorphism onto a torsion-free discrete subgroup $\rho(\Gamma)<\mathrm{PU}(1,1)$ such that $M$ and $\rho(\Gamma) \backslash \mathbb{H}_{\mathbb{C}}^{1}$ are diffeomorphic.

Theorem 4.3. Let $M$ be a p-times punctured closed orientable surface of genus $g$. Assume that $\chi(M)=2-2 g-p<0$. Let $\Gamma$ be the fundamental group of $M$ and $\rho: \Gamma \longrightarrow \mathrm{PU}(n, 1)$ be a homomorphism. Then $|\tau(\rho)| \leq-2 \pi \chi(M)$ and equality holds if and only if the image $\rho(\Gamma)$ of $\Gamma$ stabilizes a totally geodesic copy of $\mathbb{H}_{\mathbb{C}}^{1}$ in $\mathbb{H}_{\mathbb{C}}^{n}$, hence $\rho$ can be seen as a homomorphism from $\Gamma$ into $\mathrm{PU}(1,1)$ and moreover, as such, $\rho$ is a uniformization representation. If in addition $\rho$ is tame (and equality holds) then $\rho(\Gamma)<\mathrm{PU}(1,1)$ is a lattice.

Remark. (i) The number $\tau(\rho)$, up to sign, depends only on the diffeomorphism type of the surface $M$ in the following sense. If $\phi: M^{\prime} \longrightarrow M$ is a diffeomorphism and if we consider the homomorphism $\rho^{\prime}=\rho \circ \phi_{\star}$ of the fundamental group $\Gamma^{\prime}$ of $M^{\prime}$ into $\mathrm{PU}(n, 1)$, then it is easily checked that $\tau\left(\rho^{\prime}\right)= \pm \tau(\rho)$ depending on whether $\phi$ is orientation preserving or reversing.
(ii) In particular, if $\rho: \Gamma \longrightarrow \Gamma^{\prime}<\mathrm{PU}(1,1)$ is a uniformization representation, so that $M$ is diffeomorphic to $M^{\prime}:=\Gamma^{\prime} \backslash \mathbb{H}_{\mathbb{C}}^{1}$, then $\tau(\rho)= \pm \int_{M^{\prime}}\left[\omega_{1}-\mathrm{d}\left(\sum_{i} \chi \varsigma_{i}\right)\right]$. Let $M_{0}^{\prime}$ be the convex core of $M^{\prime} . M_{0}^{\prime}$ is a finite volume complete hyperbolic surface whose boundary consists of finitely many disjoint closed simple geodesics $c_{k}$ (corresponding to the conjugacy classes of peripheral elements of $\Gamma$ sent by $\rho$ to hyperbolic isometries of $\mathbb{H}_{\mathbb{C}}^{1}$ ). We can assume that $\chi=0$ on the boundary of $M_{0}^{\prime}$. It is easy to see (cf. the proof of Proposition 4.5 below) that the 1 -forms $\varsigma_{j}$ corresponding to punctures of $M_{0}^{\prime}$ are $L^{1}$ forms and therefore that

$$
\int_{M_{0}^{\prime}}\left[\omega_{1}-\mathrm{d}\left(\sum_{j} \chi \varsigma_{j}\right)\right]=\int_{M_{0}^{\prime}} \omega_{1}=\operatorname{Vol}\left(M_{0}^{\prime}\right)=-2 \pi \chi\left(M_{0}^{\prime}\right)=-2 \pi \chi(M)
$$

Moreover,

$$
\int_{M^{\prime} \backslash M_{0}^{\prime}}\left[\omega_{1}-\mathrm{d}\left(\sum_{k} \chi \varsigma_{k}\right)\right]=\sum_{k} \int_{c_{k}} \mathrm{~d} \varsigma_{k}=0
$$

as can be seen from the proof of Proposition 4.1. Hence $\tau(\rho)= \pm 2 \pi \chi(M)$ as Theorem 4.3 says.
(iii) Our definition of the invariant $\tau$ makes no reference to a Riemannian structure on the surface $M$. This means that we can equip $M$ with any Riemannian metric we want and that our results will be independent of this particular choice. Now, the uniformization theorem implies that there exist complete hyperbolic metrics of finite volume on $M$. If we choose such a metric $g_{1}$ on $M$, we obtain a uniformization representation $u: \Gamma \longrightarrow \mathrm{PU}(1,1)$ and $M$ is diffeomorphic to $u(\Gamma) \backslash \mathbb{H}_{\mathbb{C}}^{1}$. The composition $\rho \circ u^{-1}$ is a homomorphism from the torsionfree lattice $u(\Gamma)$ of $\mathrm{PU}(1,1)$ to $\mathrm{PU}(n, 1)$. These identifications are completely transparent for the invariant $\tau$ and we can therefore assume that $\Gamma$ is a torsion-free lattice in $\operatorname{PU}(1,1)$ and that $M=\Gamma \backslash \mathbb{H}_{\mathbb{C}}^{1}$ with its complete hyperbolic metric $g_{1}$ of finite volume. The punctured neighbourhoods $D_{i}^{\star}$ of the punctures $m_{i}$ will then be seen as cusps of $M$ and will often be denoted by $C_{i}$.

The metric on $M$ allows to talk about $L^{2}$-cohomology groups and if we call $\left[\rho^{\star} \omega_{n}\right]_{(2)}$ the image in $H_{(2)}^{2}(M)$ of $\left[\rho^{\star} \omega_{n}\right]_{c}$ under the comparison map $H_{c}^{2}(M) \longrightarrow H_{(2)}^{2}(M)$, we have

$$
\tau(\rho)=\frac{1}{2} \int_{M}\left\langle\left[\rho^{\star} \omega_{n}\right]_{(2)}, \omega_{1}\right\rangle d V_{1}
$$

where $\omega_{1}$ is the Kähler form of $g_{1}$.
(iv) If $\rho$ is not reductive, that is, if $\rho(\Gamma)$ fixes a point in $\partial_{\infty} \mathbb{H}_{\mathbb{C}}^{n}$, then $\tau(\rho)=0$. Indeed, we can use this fixed point to define as before a 1-form $\varsigma$ invariant by $\rho(\Gamma)$. Then $f^{\star} \omega_{n}=\mathrm{d} f^{\star} \varsigma$ on $M$. Hence $f^{\star} \omega_{n}-\mathrm{d}\left(\chi f^{\star} \varsigma\right)=\mathrm{d}\left((1-\chi) f^{\star} \varsigma\right)$, that is $\left[f^{\star} \omega_{n}-\mathrm{d}\left(\chi f^{\star} \varsigma\right)\right]_{c}=0$ and $\tau(\rho)=0$.

Hence we assume from now on that all considered representations are reductive.
The proof of Theorem 4.3 is easier if one deals only with tame representations. Since almost all the ideas are needed in this case, we present it separately in Section 4.1 and we explain in Section 4.2 how to adapt the arguments to handle the general case.

### 4.1. Tame representations.

Note that a representation $\rho$ of a torsion-free lattice $\Gamma$ of $\mathrm{PU}(1,1)$ into $\mathrm{PU}(n, 1)$ is tame if and only if $\rho$ maps no parabolic elements of $\Gamma$ to hyperbolic elements of $\operatorname{PU}(n, 1)$. The name "tame" is motivated by the following proposition:

Proposition 4.4. Let $\Gamma<\mathrm{PU}(1,1)$ be a torsion-free lattice, and let $\rho: \Gamma \longrightarrow \mathrm{PU}(n, 1)$ be a homomorphism. There exists a $\rho$-equivariant map $f: \mathbb{H}_{\mathbb{C}}^{1} \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$ of finite energy if and only if $\rho$ is tame.

Proof. When $\rho$ is tame, we may easily construct a $\rho$-equivariant map with finite energy. In fact, we use the same method as in Theorem 1.2; since the fundamental group of each cusp of $M$ is generated by a single parabolic element, the construction is much simpler. Namely, it is sufficient to define the $\rho$-equivariant map on each cusp of $M$. Let $C$ be a cusp of $M$ and let $\gamma$ be a parabolic element of $\Gamma$ generating $\pi_{1}(C)$ (via the usual identification). We can choose (horospherical) coordinates $(v, t)$ on $\mathbb{H}_{\mathbb{C}}^{1}$ such that $\gamma(v, t)=(v+a, t)$ and such that $C$ is isometric to the quotient by $\langle\gamma\rangle$ of the subset $D:=[0, a] \times[0,+\infty)$ of $\mathbb{H}_{\mathbb{C}}^{1}$ endowed with the metric $g_{1}=e^{-2 t} d v^{2}+d t^{2}$.

If $f: \mathbb{H}_{\mathbb{C}}^{1} \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$ is any $\rho$-equivariant map, the energy of $f$ in the cusp is given by

$$
E_{C}(f)=\frac{1}{2} \int_{0}^{+\infty} \int_{0}^{a}\|\mathrm{~d} f\|_{(v, t)}^{2} e^{-t} d v d t
$$

When $\rho(\gamma)$ is elliptic, we map $D$ to a fixed point of $\rho(\gamma)$ and the energy of $f$ in the cusp is zero. If $\rho(\gamma)$ is parabolic, let $\left(z^{\prime}, v^{\prime}, t^{\prime}\right)$ be adapted horospherical coordinates on $\mathbb{H}_{\mathbb{C}}^{n}$. We define

$$
\begin{aligned}
f:[0, a] \times[0,+\infty) & \longrightarrow \mathbb{H}_{\mathbb{C}}^{n} \\
(v, t) & \longmapsto(\varphi(v), 2 t)
\end{aligned}
$$

where $\varphi$ is a map from $[0, a]$ into a horosphere $H S^{\prime} \subset \mathbb{H}_{\mathbb{C}}^{n}$ such that $\varphi(a)=\rho(\gamma) \varphi(0)$. Computing as in the proof of Theorem 1.2 (case 1: $\rho\left(\gamma_{0}\right)$ is parabolic), we get the following estimation of the energy of $f$ in the cusp:

$$
E_{C}(f) \leq \frac{1}{2} \int_{0}^{+\infty} \int_{0}^{a}\left(4+\|\mathrm{d} \varphi\|^{2}\right) e^{-t} d v d t \leq 2 a+E(\varphi)<+\infty .
$$

Suppose now that $\rho(\gamma)$ is hyperbolic and let $f: \mathbb{H}_{\mathbb{C}}^{1} \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$ be any $\rho$-equivariant map. For every $t$, we denote by $c_{t}:[0, a] \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$ the curve $v \longmapsto f(v, t)$. Since $\rho(\gamma)$ is hyperbolic, there
exists $\delta>0$ such that the distance between $f(0, t)$ and $f(a, t)=\rho(\gamma) f(0, t)$ is at least $\delta$, hence the length $l\left(c_{t}\right)$ of $c_{t}$ is at least $\delta$. This implies that

$$
\frac{1}{2} \int_{0}^{a}\|\mathrm{~d} f\|_{(v, t)}^{2} d v \geq e^{2 t} E\left(c_{t}\right)=\frac{e^{2 t}}{2} \int_{0}^{a}\left\|\mathrm{~d} c_{t}\right\|^{2} d v \geq \frac{e^{2 t}}{2 a}\left(l\left(c_{t}\right)\right)^{2} \geq \frac{e^{2 t}}{2 a} \delta^{2}
$$

and

$$
E_{C}(f) \geq \frac{\delta^{2}}{2 a} \int_{0}^{+\infty} e^{t} d t=+\infty
$$

The next proposition shows that for a tame representation $\rho$, the invariant $\tau(\rho)$ can be computed with any finite energy $\rho$-equivariant map $\mathbb{H}_{\mathbb{C}}^{1} \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$.

Proposition 4.5. Let $\Gamma<\operatorname{PU}(1,1)$ be a torsion-free lattice and let $\rho: \Gamma \longrightarrow \mathrm{PU}(n, 1)$ be a tame homomorphism. Then,

$$
\tau(\rho)=\frac{1}{2} \int_{M}\left\langle f^{\star} \omega_{n}, \omega_{1}\right\rangle d V_{1}
$$

for any finite energy $\rho$-equivariant map $f: \mathbb{H}_{\mathbb{C}}^{1} \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$.
Remark. One could therefore take this as a definition of the invariant for tame representations. This gives a formulation very similar to the classical one for the Toledo invariant of a uniform lattice.

Note that the energy finiteness assumption is necessary as the following simple example shows. Let $\Gamma$ be a torsion-free lattice in $\operatorname{SU}(1,1)$ generated by two hyperbolic elements $\gamma_{1}$ and $\gamma_{2}$ such that $\gamma_{0}=\left[\gamma_{1}, \gamma_{2}\right]$ is parabolic. Then $M=\Gamma \backslash \mathbb{H}_{\mathbb{C}}^{1}$ is diffeomorphic to a oncepunctured torus. Let $\rho$ be the inclusion $\Gamma \longrightarrow \mathrm{PU}(1,1)$. Keeping the same notations as in the proof of Proposition 4.4, we can define $C^{\infty}$ maps $f_{\mu}: M \longrightarrow M$ which equal identity outside the cusp, and given by $f_{\mu}:[0, a] \times[0,+\infty) \longrightarrow[0, a] \times \mathbb{R},(v, t) \longmapsto(v, \mu(t))$ in $D$, where $\mu \in C^{\infty}([0,+\infty), \mathbb{R})$. The energy of $f_{\mu}$ in the cusp is $\frac{a}{2} \int_{0}^{+\infty}\left(e^{t-2 \mu(t)}+\mu^{\prime}(t)^{2} e^{-t}\right) d t$. Moreover, a formal computation gives: $\frac{1}{2} \int_{C}\left\langle f_{\mu}^{\star} \omega_{1}, \omega_{1}\right\rangle d V_{1}=\int_{C} f_{\mu}^{\star} \omega_{1}=a \int_{0}^{+\infty} \mu^{\prime}(t) e^{-\mu(t)} d t=$ $a\left(1-\lim _{t \rightarrow+\infty} e^{-\mu(t)}\right)$. Note that this limit needs not exist and that for every $c \in[-\infty, \tau(\rho)]$, we may choose $\mu$ such that $\int_{M} f_{\mu}^{\star} \omega_{1}=c$. But if the energy of $f_{\mu}$ is finite, then $\mu(t) \rightarrow+\infty$ as $t \rightarrow+\infty$, and then, $\int_{M} f_{\mu}^{\star} \omega_{1}=\int_{M} \omega_{1}=\tau(\rho)$.
Proof. Let $f: \mathbb{H}_{\mathbb{C}}^{1} \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$ be a finite energy $\rho$-equivariant map. Since $f$ has finite energy, $\left\langle f^{\star} \omega_{n}, \omega_{1}\right\rangle=2\left(e^{\prime}(f)-e^{\prime \prime}(f)\right)$ is integrable on $M$ and we can write

$$
\tau(\rho)=\frac{1}{2} \int_{M}\left\langle f^{\star} \omega_{n}, \omega_{1}\right\rangle d V_{1}-\frac{1}{2} \int_{M}\left\langle\mathrm{~d}\left(\sum_{i} \chi f^{\star} \varsigma_{i}\right), \omega_{1}\right\rangle d V_{1} .
$$

The function $\left\langle\mathrm{d}\left(\sum_{i} \chi f^{\star} \varsigma_{i}\right), \omega_{1}\right\rangle$ is integrable on $M$ (because so is $\left\langle f^{\star} \omega_{n}, \omega_{1}\right\rangle$ in the cusps of $M$ ). Therefore, if we prove that the 1-form $\sum_{i} \chi f^{\star} \varsigma_{i}$ is an $L^{1}$-form, the Stokes formula of Ga54] will apply and $\int_{M}\left\langle\mathrm{~d}\left(\sum_{i} \chi f^{\star} \varsigma_{i}\right), \omega_{1}\right\rangle d V_{1}$ will vanish as wanted.

Now, if the generator $\gamma_{i}$ of the fundamental group of the cusp $C_{i}$ is sent by $\rho$ to a parabolic element of $\operatorname{PU}(n, 1)$, we know that in horocyclic coordinates relative to the fixed point of $\rho\left(\gamma_{i}\right)$ we have $g_{n}=d t^{\prime 2}+\varsigma_{i}^{2}+4 e^{-t^{\prime}}\left\langle\left\langle d z^{\prime}, d z^{\prime}\right\rangle\right\rangle$. Therefore,

$$
\|\mathrm{d} f\|^{2}=\operatorname{tr}_{g_{1}}\left(f^{\star} g_{n}\right) \geq\left\|f^{\star} \varsigma_{i}\right\|^{2} .
$$

If $\rho\left(\gamma_{i}\right)$ is elliptic, we can assume that one of its fixed points is 0 in the ball model of $\mathbb{H}_{\mathbb{C}}^{n}$. In this case, we have $\varsigma_{i}=-\mathrm{d}^{c} \psi_{i}$ with $\psi_{i}=\log (1-\langle\langle z, z\rangle\rangle)$. The metric $g_{n}$ is then given by

$$
g_{n}=4 \frac{\langle\langle d z, d z\rangle\rangle}{1-\langle\langle z, z\rangle\rangle}+\left(\mathrm{d} \psi_{i}\right)^{2}+\left(\mathrm{d}^{c} \psi_{i}\right)^{2} .
$$

Again, we have $\left\|f^{\star} \varsigma_{i}\right\| \leq\|\mathrm{d} f\|$ on $C_{i}$.
The form $f^{\star} \varsigma_{i}$ is therefore $L^{1}$ on $C_{i}$ for each $i$. The proposition follows.
We are now in position to prove Theorem 4.3 in the case of reductive and tame representations.
Proof of Theorem 4.3. Since $\rho$ is assumed to be tame and reductive, Proposition 4.4 and Theorem 1.1 imply the existence of a $\rho$-equivariant harmonic map $f: \mathbb{H}_{\mathbb{C}}^{1} \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$ of finite energy.

Proceeding exactly as in section 3.2 yields the inequality $|\tau(\rho)| \leq \operatorname{Vol}(M)$.
Before treating the equality case, we remark that there are particular representations for which it is immediate that equality can not hold:

Claim 4.6. Suppose that $\rho$ maps the conjugacy classes $c_{q+1}, \ldots, c_{p}$ of peripheral elements of $\Gamma$ to $\mathrm{id} \in \operatorname{PU}(n, 1)$ and denote by $M^{\prime}$ the surface obtained by removing only the points $m_{1}, \ldots$, $m_{q}$ from the closed orientable surface $\bar{M}$ of genus $g$. Then, $|\tau(\rho)| \leq \max \left(0,-2 \pi \chi\left(M^{\prime}\right)\right)$.
Proof. The representation $\rho$ admits a factorisation $\rho=\rho^{\prime} \circ i_{\star}$ where $i_{\star}: \Gamma \longrightarrow \Gamma^{\prime}:=\pi_{1}\left(M^{\prime}\right)$ is the representation induced by the inclusion $i: M \longrightarrow M^{\prime}$ and $\rho^{\prime}: \Gamma^{\prime} \longrightarrow \mathrm{PU}(n, 1)$ is a homomorphism. We may equip $M^{\prime}$ with a complex structure such that the inclusion $i$ is holomorphic.
If $\chi\left(M^{\prime}\right)<0$, we endow $M^{\prime}$ with a complete hyperbolic metric of finite volume. Then, $\left|\tau\left(\rho^{\prime}\right)\right| \leq-2 \pi \chi\left(M^{\prime}\right)$ by the above argument. If $f$ is any section of the bundle $M^{\prime} \times{ }_{\rho^{\prime}} \mathbb{H}_{\mathbb{C}}^{n}$, the restriction of $f$ to $M$ can be seen as a section of the bundle $M \times{ }_{\rho} \mathbb{H}_{\mathbb{C}}^{n}$ whose energy is finite on neighbourhoods of the punctures $m_{q+1}, \ldots, m_{p}$ (since energy finiteness only depends on the conformal structure of $M$ ). Then, by Proposition 4.5, $\tau(\rho)=\int_{M}\left[f^{\star} \omega_{n}-\mathrm{d}\left(\sum_{i=1}^{q} \chi f^{\star} \varsigma_{i}\right)\right]=$ $\int_{M^{\prime}}\left[f^{\star} \omega_{n}-\mathrm{d}\left(\sum_{i=1}^{q} \chi f^{\star} \varsigma_{i}\right)\right]=\tau\left(\rho^{\prime}\right)$.

If $\chi\left(M^{\prime}\right)=0$, then $M^{\prime}$ is holomorphically equivalent to either $\mathbb{C}^{\star}$ or an elliptic curve. In the first case, $\Gamma^{\prime} \simeq \mathbb{Z}$ and so $\rho(\Gamma)$ is generated by a single element. If this element is parabolic or hyperbolic, then $\rho$ is not reductive and if it is elliptic, there exists a constant $\rho$-equivariant map. Thus, $\tau(\rho)=0$. In the second case, $\Gamma^{\prime}$ is abelian, generated by two elements $\gamma_{1}$ and $\gamma_{2}$. If $\rho^{\prime}\left(\gamma_{1}\right)$ is parabolic or hyperbolic, by commutation, $\rho^{\prime}\left(\gamma_{2}\right)$ must fix the fixed point(s) of $\rho^{\prime}\left(\gamma_{1}\right)$ and $\rho$ is not reductive. If $\rho^{\prime}\left(\gamma_{1}\right)$ and $\rho^{\prime}\left(\gamma_{2}\right)$ are elliptic, they must have a common fixed point in $\mathbb{H}_{\mathbb{C}}^{n}$, as they commute. In either case, $\tau(\rho)=0$.

Finally, if $\chi\left(M^{\prime}\right)>0, M^{\prime}$ is simply connected and $\tau(\rho)$ is of course zero.
We suppose now that the equality $|\tau(\rho)|=-2 \pi \chi(M)$ holds.
The harmonic map $f: \mathbb{H}_{\mathbb{C}}^{1} \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$ needs not be (anti)holomorphic as in the higher dimensional case, but we know that its real rank is 2 at some point, hence on a dense open subset $\mathcal{U}$ of $\mathbb{H}_{\mathbb{C}}^{1}$ by a result of [Sa78], and that in the proof of Theorem 3.1, one of the inequalities concerning $R^{\prime}$ and $R^{\prime \prime}$ is necessarily an equality on $M$.

Suppose for example that $R^{\prime}=\frac{1}{4} e^{\prime}\left(e^{\prime}-e^{\prime \prime}\right)$ (the case where the inequality becoming an equality is $R^{\prime \prime} \geq \frac{1}{4} e^{\prime \prime}\left(e^{\prime \prime}-e^{\prime}\right)$ is handled similarly), that is

$$
\left|g_{n}\left(\partial_{x}^{1,0} f\left(z_{1}\right), \overline{\partial_{x}^{1,0} f\left(\bar{z}_{1}\right)}\right)\right|^{2}=g_{n}\left(\partial_{x}^{1,0} f\left(z_{1}\right), \overline{\partial_{x}^{1,0} f\left(z_{1}\right)}\right) g_{n}\left(\partial_{x}^{1,0} f\left(\bar{z}_{1}\right), \overline{\partial_{x}^{1,0} f\left(\bar{z}_{1}\right)}\right)
$$

for every $x \in \mathbb{H}_{\mathbb{C}}^{1}$.

This has the following simple but very important consequence: for all $x \in \mathbb{H}_{\mathbb{C}}^{1}, \mathrm{~d}_{x} f\left(T_{x} \mathbb{H}_{\mathbb{C}}^{1}\right)$ is contained in a complex one-dimensional subspace of $T_{f(x)} \mathbb{H}_{\mathbb{C}}^{n}$. To prove this fact, we only need to consider points where $\mathrm{rk}_{\mathbb{R}} \mathrm{d}_{x} f=2$. We remark that a vector subspace $V \subset T_{f(x)} \mathbb{H}_{\mathbb{C}}^{n}$ is $J$-invariant if and only if

$$
V^{\mathbb{C}}=V^{1,0} \oplus V^{0,1}=V^{1,0} \oplus \overline{V^{1,0}}
$$

where $V^{\mathbb{C}} \subset T_{f(x)}^{\mathbb{C}} \mathbb{H}_{\mathbb{C}}^{n}$ is the complexification of $V$ and $V^{1,0}$ resp. $V^{0,1}$ is the projection of $V^{\mathbb{C}}$ on $T_{f(x)}^{1,0} \mathbb{H}_{\mathbb{C}}^{n}$ resp. $T_{f(x)}^{0,1} \mathbb{H}_{\mathbb{C}}^{n}$. So, if $V=\mathrm{d}_{x} f\left(T_{x} \mathbb{H}_{\mathbb{C}}^{1}\right)$ is a 2-dimensional real subspace of $T_{f(x)} \mathbb{H}_{\mathbb{C}}^{n}$, $V$ is $J$-invariant if and only if $\operatorname{dim}_{\mathbb{C}} V^{1,0}=1$. Since $V^{1,0}$ is spanned by $\partial_{x}^{1,0} f\left(z_{1}\right)$ and $\partial_{x}^{1,0} f\left(\bar{z}_{1}\right)$, the above equality holds if and only if $\mathrm{d}_{x} f\left(T_{x} \mathbb{H}_{\mathbb{C}}^{1}\right)$ is contained in a complex one-dimensional subspace of $T_{f(x)} \mathbb{H}_{\mathbb{C}}^{n}$, according to the equality part of Cauchy-Schwarz's inequality.

Now, we will use the same trick as A. G. Reznikov in his paper Re93]. Namely, we equip the product $\mathbb{H}_{\mathbb{C}}^{1} \times \mathbb{H}_{\mathbb{C}}^{n}$ with the metric $h_{\varepsilon}:=\varepsilon g_{1}+g_{n}($ for a given $\varepsilon>0)$ and we consider the map $\phi: \mathbb{H}_{\mathbb{C}}^{1} \longrightarrow \mathbb{H}_{\mathbb{C}}^{1} \times \mathbb{H}_{\mathbb{C}}^{n}$ given by $\phi(x)=(x, f(x))$. $\phi$ is then a harmonic embedding of $\mathbb{H}_{\mathbb{C}}^{1}$ into $\mathbb{H}_{\mathbb{C}}^{1} \times \mathbb{H}_{\mathbb{C}}^{n}$. Let $\kappa_{\phi^{\star} h_{\varepsilon}}$ be the Gaussian curvature of the induced metric $\phi^{\star} h_{\varepsilon}$ on $\mathbb{H}_{\mathbb{C}}^{1}$. Since $\phi$ is harmonic, Lemma C. 4 of [Re93] implies that for all $x \in \mathbb{H}_{\mathbb{C}}^{1}, \kappa_{\phi^{\star} h_{\varepsilon}}(x)$ is less than or equal to the value of the sectional curvature of $h_{\varepsilon}$ on the 2-plane $T_{\phi(x)} \phi\left(\mathbb{H}_{\mathbb{C}}^{1}\right) \subset T_{\phi(x)}\left(\mathbb{H}_{\mathbb{C}}^{1} \times \mathbb{H}_{\mathbb{C}}^{n}\right)$ (which we will call $\kappa_{h_{\varepsilon}}\left(T_{\phi(x)} \phi\left(\mathbb{H}_{\mathbb{C}}^{1}\right)\right)$ ).

Since $\phi$ is equivariant w.r.t. the isometric diagonal action of $\Gamma$ on $\mathbb{H}_{\mathbb{C}}^{1} \times \mathbb{H}_{\mathbb{C}}^{n}$, these quantities are well defined over the Riemann surface $M=\Gamma \backslash \mathbb{H}_{\mathbb{C}}^{1}$. Moreover, we have
Claim 4.7. For each $\varepsilon>0$,

$$
\operatorname{Area}_{g_{1}}(M)=-2 \pi \chi(M)=-\int_{M} \kappa_{\phi^{\star} h_{\varepsilon}} d V_{\phi^{\star} h_{\varepsilon}}
$$

where $d V_{\phi^{\star} h_{\varepsilon}}$ denotes the volume element of the pull-back metric $\phi^{\star} h_{\varepsilon}$ on $M$.
Proof. We shall apply the main theorem of [Li97] to $\phi^{\star} h_{\varepsilon}$. This metric is clearly complete (because $g_{1}$ is), and has everywhere non-positive curvature. We have to prove that

$$
\lim _{r \rightarrow+\infty} \frac{\operatorname{Area}_{\phi^{\star}} h_{\varepsilon}\left(B_{\phi^{\star}} h_{\varepsilon}(r)\right)}{r^{2}}=0
$$

where $B_{\phi^{\star} h_{\varepsilon}}(r)$ is the geodesic ball of radius $r$ (at a fixed point that belongs, for example, to the compact part of $M$ ) w.r.t. the metric $\phi^{\star} h_{\varepsilon}$. Now, simple computations show that

$$
d V_{\phi^{\star} h_{\varepsilon}}=\sqrt{\varepsilon^{2}+2 \varepsilon e(f)+\frac{1}{2}\left\|f f^{\star} \omega_{n}\right\|^{2}} d V_{1} .
$$

Since $f$ has finite energy and $f^{\star} \omega_{n}$ is $L^{2}$, the metric $\phi^{\star} h_{\varepsilon}$ has in fact finite volume and we obtain the expected limit. Then, it follows from [Li97] that $\phi^{\star} h_{\varepsilon}$ has finite total curvature and this implies the result (as it is explained in the same paper).

On $\mathcal{U}$, the pull-back $f^{\star} g_{n}$ is a Riemannian metric. Let $\kappa_{f^{\star} g_{n}}$ denote its Gaussian curvature. Again, since $f$ is harmonic, for all $x \in \mathcal{U}, \kappa_{f^{\star} g_{n}}(x)$ is less than or equal to $\kappa_{g_{n}}\left(T_{f(x)} f\left(\mathbb{H}_{\mathbb{C}}^{1}\right)\right)$, the sectional curvature of the 2-plane $T_{f(x)} f\left(\mathbb{H}_{\mathbb{C}}^{1}\right) \subset T_{f(x)} \mathbb{H}_{\mathbb{C}}^{n}$ w.r.t. the metric $g_{n}$. Since $T_{f(x)} f\left(\mathbb{H}_{\mathbb{C}}^{1}\right)$ is a complex line in $T_{f(x)} \mathbb{H}_{\mathbb{C}}^{n}$, its sectional curvature is -1 . Hence, for all $x \in \mathcal{U}$, $\kappa_{f^{\star} g_{n}}(x) \leq-1$.

Assume that at a certain point $p$ of $\mathcal{U}, \kappa_{f \star g_{n}}(p)$ is stricly less than -1 . We are going to prove that this can not happen. For, if it does, there exist $\alpha>0$ and a small disc $D$ around $p$ whose closure $\bar{D}$ is contained in $\mathcal{U}$ such that for all $x \in D, \kappa_{f \star g_{n}}(x) \leq-1-\alpha$. We can
assume that the image of $D$ in $M$ is diffeomorphic to $D$, and we still call it $D$. Now, if $K$ is any compact subset of $\mathcal{V}:=\Gamma \backslash \mathcal{U}$ containing $D$, we have

$$
\begin{aligned}
\operatorname{Area}_{g_{1}}(M) & \geq-\int_{K} \kappa_{\phi^{\star} h_{\varepsilon}}(x) d V_{\phi^{\star} h_{\varepsilon}} \\
& \geq-\int_{D} \kappa_{\phi^{\star} h_{\varepsilon}}(x) d V_{\phi^{\star} h_{\varepsilon}}-\int_{K \backslash D} \kappa_{h_{\varepsilon}}\left(T_{\phi(x)} \phi\left(\mathbb{H}_{\mathbb{C}}^{1}\right)\right) d V_{\phi^{\star} h_{\varepsilon}}
\end{aligned}
$$

since on $K, \kappa_{\phi^{\star} h_{\varepsilon}} \leq \kappa_{h_{\varepsilon}} \leq 0$.
When $\varepsilon$ goes to 0 , the induced metric $\phi^{\star} h_{\varepsilon}$ clearly goes to $f^{\star} g_{n}$ on $D$ and therefore, on $D$, the curvature and the volume form of $\phi^{\star} h_{\varepsilon}$ respectively go to the curvature and volume form of $f^{\star} g_{n}$. In the same manner, on the compact set $K, \kappa_{h_{\varepsilon}}\left(T_{\phi(x)} \phi\left(\mathbb{H}_{\mathbb{C}}^{1}\right)\right)$ goes to $\kappa_{g_{n}}\left(T_{f(x)} f\left(\mathbb{H}_{\mathbb{C}}^{1}\right)\right)$. Again, all involved quantities are well defined on $\mathcal{V}$. Hence

$$
\operatorname{Area}_{g_{1}}(M) \geq-\int_{D} \kappa_{f^{\star} g_{n}}(x) d V_{f^{\star} g_{n}}-\int_{K \backslash D} \kappa_{g_{n}}\left(T_{f(x)} f\left(\mathbb{H}_{\mathbb{C}}^{1}\right)\right) d V_{f^{\star} g_{n}}
$$

Using $\kappa_{g_{n}}\left(T_{f(x)} f\left(\mathbb{H}_{\mathbb{C}}^{1}\right)\right)=-1$ and remembering our assumption on $D$, we obtain

$$
\operatorname{Area}_{g_{1}}(M) \geq(1+\alpha) \operatorname{Area}_{f^{\star} g_{n}}(D)+\operatorname{Area}_{f^{\star} g_{n}}(K \backslash D)=\alpha \operatorname{Area}_{f^{\star} g_{n}}(D)+\operatorname{Area}_{f^{\star} g_{n}}(K)
$$

This is true for any compact subset $K$ of $\mathcal{V}$, hence

$$
\operatorname{Area}_{g_{1}}(M) \geq \alpha \operatorname{Area}_{f^{\star} g_{n}}(D)+\operatorname{Area}_{f^{\star} g_{n}}(\mathcal{V})
$$

Since $f(\mathcal{U})$ is (locally) a 1-dimensional complex submanifold of $\mathbb{H}_{\mathbb{C}}^{n}$, the volume form of the metric induced on it by $g_{n}$ is simply the restriction of $\omega_{n}$. Therefore the volume form of the metric $f^{\star} g_{n}$ on $\mathcal{V}$ is $\left|f^{\star} \omega_{n}\right|=\left|e^{\prime}-e^{\prime \prime}\right| \omega_{1}$. Hence

$$
\begin{aligned}
\operatorname{Area}_{g_{1}}(M) & \geq \alpha \operatorname{Area}_{f^{\star} g_{n}}(D)+\int_{\mathcal{V}}\left|e^{\prime}-e^{\prime \prime}\right| \omega_{1} \\
& =\alpha \operatorname{Area}_{f^{\star} g_{n}}(D)+\int_{M}\left|e^{\prime}-e^{\prime \prime}\right| \omega_{1} \\
& \geq \alpha \operatorname{Area}_{f^{\star} g_{n}}(D)+\left|\int_{M}\left(e^{\prime}-e^{\prime \prime}\right) \omega_{1}\right|
\end{aligned}
$$

But we assumed that $\int_{M}\left(e^{\prime}-e^{\prime \prime}\right) \omega_{1}=\operatorname{Area}_{g_{1}}(M)$, so this gives $e^{\prime}-e^{\prime \prime} \geq 0$ on $M$ and the desired contradiction: we conclude that the curvature of the metric induced by $g_{n}$ on $f(\mathcal{U})$ is everywhere -1 .

If we now take a small open set $U$ in $\mathcal{U}$ on which $f$ is an embedding, $f(U)$ is a complex, hence minimal, submanifold of $\mathbb{H}_{\mathbb{C}}^{n}$ whose sectional curvature equals the restriction of the sectional curvature of the ambient manifold $\mathbb{H}_{\mathbb{C}}^{n}: f(U)$ must be totally geodesic and therefore contained in a complex geodesic. By a result of Sampson (Sa78), this implies that $f$ maps $\mathbb{H}_{\mathbb{C}}^{1}$ entirely into this complex geodesic. In particular, $\rho(\Gamma)$ stabilizes this totally geodesic copy of $\mathbb{H}_{\mathbb{C}}^{1}$ and $\rho$ thus induces a tame representation of $\Gamma$ into $\mathrm{PU}(1,1)$ that we shall still denote by $\rho$ in the sequel. In the same way, $f$ will be seen as a $\rho$-equivariant map from $\mathbb{H}_{\mathbb{C}}^{1}$ into $\mathbb{H}_{\mathbb{C}}^{1}$.

Claim 4.8. The map $f: \mathbb{H}_{\mathbb{C}}^{1} \longrightarrow \mathbb{H}_{\mathbb{C}}^{1}$ is a local diffeomorphism.
Proof. We apply the method of Toledo in To79] Theorem 4.2, to prove that $e^{\prime}-e^{\prime \prime}>0$ everywhere on $M$. We first go back to the proof of Theorem 3.1, keeping the same notations. We know that $\Delta \log e^{\prime}=2\left(e^{\prime \prime}-e^{\prime}\right)+2$, wherever $e^{\prime}>0$. The zeros of $e^{\prime}$ (if any) are known to be isolated and of finite order. Suppose that there exists $x \in M$ such that $e^{\prime}(x)=0$, and let $r_{0}, \varepsilon_{0}>0$ such that $e^{\prime} \geq \varepsilon_{0}$ on $\partial \mathrm{D}\left(x, r_{0}\right)$ (where $\mathrm{D}\left(x, r_{0}\right)$ is the disc of radius $r_{0}$ about $x$ ). For any $0<r<r_{0}$ and $0<\varepsilon<\varepsilon_{0}$, we set $M_{r, \varepsilon}=\left(M_{\varepsilon} \backslash \mathrm{D}\left(x, r_{0}\right)\right) \cup\left(\mathrm{D}\left(x, r_{0}\right) \backslash \mathrm{D}(x, r)\right)$ and $M_{\varepsilon}^{\prime}=M_{\varepsilon} \cup \mathrm{D}\left(x, r_{0}\right)$.

By Green's formula, we have

$$
\int_{M_{r, \varepsilon}}\left[\eta_{R} \Delta \log e^{\prime}-\left\langle\operatorname{grad} \eta_{R}, \operatorname{grad} \log e^{\prime}\right\rangle\right] d V_{1}=\int_{\partial M_{r, \varepsilon}} \eta_{R}\left\langle\nu, \operatorname{grad} \log e^{\prime}\right\rangle d s
$$

and letting $r \rightarrow 0$, then $R \rightarrow+\infty$, we get

$$
\int_{M_{\varepsilon}^{\prime}}\left(2\left(e^{\prime \prime}-e^{\prime}\right)+2\right) d V_{1} \geq 4 \pi q
$$

where $2 q>0$ is the order of the zero of $e^{\prime}$ at $x$ (see To79 and the proof of Theorem 3.1). But, letting $\varepsilon \rightarrow 0$, we get a contradiction with $\tau(\rho)=-2 \pi \chi(M)$. Finally, we thus have $e^{\prime}>0$ and $e^{\prime}-e^{\prime \prime} \geq 0$ everywhere on $M$.

Now, if $e^{\prime}(x)=e^{\prime \prime}(x)$ at some point $x \in M$, there exists a positive constant $A$ such that

$$
-\Delta \log \frac{e^{\prime}}{e^{\prime \prime}} \leq 4\left(e^{\prime}-e^{\prime \prime}\right) \leq A \log \frac{e^{\prime}}{e^{\prime \prime}}
$$

in a neighbourhood of $x$. This implies (see SY78 and references therein) that $\log \left(e^{\prime} / e^{\prime \prime}\right)$ is identically zero on a non empty open subset of $M$, which is impossible, as the rank of $\mathrm{d} f$ is generically 2 .

We may therefore pull-back the complex structure of the target to obtain a complex structure on $\mathbb{H}_{\mathbb{C}}^{1}$ with respect to which $f$ is holomorphic. In the following discussion, $M$ will be endowed with the new induced complex structure.

As a consequence of the uniformization theorem, there exists a properly discontinuous subgroup $\Gamma^{\prime}<\mathrm{PU}(1,1)$ such that $M$ is holomorphically equivalent to $\Gamma^{\prime} \backslash \mathbb{H}_{\mathbb{C}}^{1}$.
Claim 4.9. The discrete subgroup $\Gamma^{\prime}<\mathrm{PU}(1,1)$ is a lattice.
Proof. Each puncture of $M$ has a neighbourhood which is holomorphically equivalent to either a punctured disc or an annulus. We have to show that the latter cannot happen.

Let $m$ be a puncture of $M$ and suppose that $m$ has a neighbourhood $U$ which is holomorphically equivalent to an annulus

$$
A_{b}=\left\{z \in \mathbb{C}, e^{-\pi / b}<|z|<1\right\}, b>0
$$

In the former complex structure, we have horocyclic coordinates $(v, t) \in[0, a] \times[0,+\infty)$ on a fundamental domain of the cusp $C$ corresponding to $m$. Let us denote as before by $c_{t}:[0, a] \longrightarrow \mathbb{H}_{\mathbb{C}}^{1}$ the curve $v \longmapsto f(v, t)$. Since the harmonic $\rho$-equivariant map $f$ has finite energy, there exists a sequence $\left(t_{k}\right)_{k \in \mathbb{N}}$ going to infinity such that $c_{k}:=c_{t_{k}}$ verifies

$$
\lim _{k \rightarrow+\infty}\left(l\left(c_{k}\right)\right)^{2} \leq \lim _{k \rightarrow+\infty} a \int_{0}^{a}\left\|\mathrm{~d} c_{k}\right\|^{2} d v \leq \lim _{k \rightarrow+\infty} a e^{-t_{k}} \int_{0}^{a}\|\mathrm{~d} f\|_{\left(v, t_{k}\right)}^{2} e^{-t_{k}} d v=0
$$

where $l\left(c_{k}\right)$ is the length of $c_{k}$.
If $\gamma$ is a peripheral element of $\Gamma$ generating $\pi_{1}(C), \rho(\gamma)$ is either parabolic or elliptic, because $\rho$ is tame. In the first case, $f$ induces a holomorphic function $f_{U}$ from $A_{b}$ into $\langle\rho(\gamma)\rangle \backslash \mathbb{H}_{\mathbb{C}}^{1}$ (which is holomorphically equivalent to a punctured disc). The loops $\alpha_{k}:[0, a] \longrightarrow C, v \longmapsto\left(v, t_{k}\right)$ define a sequence $\left(\beta_{k}\right)$ of simple loops in $A_{b}$ (with index 1 with respect to the origin), and whose supports converge to the circle $C_{b}=\left\{z \in \mathbb{C},|z|=e^{-\pi / b}\right\}$. As $\lim _{k \rightarrow+\infty} l\left(c_{k}\right)=0$, for every $\varepsilon>0$, there exists $k_{\varepsilon} \in \mathbb{N}$ such that $k \geq k_{\varepsilon}$ implies $\left|f_{U} \circ \beta_{k}(t)\right| \leq \varepsilon$ for any $t \in[0, a]$. By the maximum principle, $\left|f_{U}(z)\right| \leq \varepsilon$ for any $z$ in the connected component of $A_{b}$ lying between the support of $\beta_{k_{\varepsilon}}$ and $C_{b}$. But the holomorphic function $f_{U}$ admits a Laurent series
expansion $f_{U}(z)=\sum_{n \in \mathbb{Z}} a_{n} z^{n}$ on $A_{b}$ and since, for every $e^{-\pi / b}<r<1$,

$$
\left|a_{n}\right|=\left|\frac{1}{2 i \pi} \int_{|z|=r} \frac{f(\zeta)}{\zeta^{n+1}} d \zeta\right| \leq e^{n \pi / b} \sup _{|z|=r}|f(z)|,
$$

$f_{U}$ should be identically zero, which is a contradiction.
If $\rho(\gamma)$ is elliptic, we know from Claim 4.6 that $\rho(\gamma) \neq \mathrm{id}$ and thus we may assume that it is a rotation with center $0 \in \mathbb{H}_{\mathbb{C}}^{1}$ and angle $\theta \in(0,2 \pi)$. We now use the cover map $S \longrightarrow A_{b}$, $w \longmapsto e^{i z / b}$ where $S$ is the strip $\{w \in \mathbb{C}, 0<\operatorname{Im} w<\pi\}$. The $\rho$-equivariant map $f$ induces a holomorphic map $f_{S}: S \longrightarrow \mathbb{H}_{\mathbb{C}}^{1}$ such that $f(w+2 \pi b)=e^{i \theta} f(w)$ for every $w \in S$. But the holomorphic map $g_{S}: S \longrightarrow \mathbb{C}, w \longmapsto e^{-i \frac{\theta}{2 \pi b} w} f_{S}(w)$ descends to $A_{b}$ and, since $\left|g_{S}\right| \leq e^{\frac{\theta}{2 b}}\left|f_{S}\right|$, the previous arguments applied to $g_{S}$ instead of $f_{U}$ do imply the same contradiction.

The conclusion of Theorem 4.3 now easily follows. Indeed, denoting by $\omega_{1}^{\prime}$ the Kähler form of the induced complete hyperbolic metric of finite volume on $M \simeq \Gamma^{\prime} \backslash \mathbb{H}_{\mathbb{C}}^{1}$, we may apply the Schwarz-Pick lemma to obtain $0 \leq\left\langle f^{\star} \omega_{n}, \omega_{1}^{\prime}\right\rangle \leq 2$ pointwise. Moreover,

$$
\tau(\rho)=\int_{M} f^{\star} \omega_{n}=\frac{1}{2} \int_{M}\left\langle f^{\star} \omega_{n}, \omega_{1}^{\prime}\right\rangle d V_{1}^{\prime} \leq \int_{M} d V_{1}^{\prime}=-2 \pi \chi(M) .
$$

Therefore $\left\langle f^{\star} \omega_{n}, \omega_{1}^{\prime}\right\rangle=2$ everywhere and hence, if $\rho^{\prime}: \Gamma^{\prime} \longrightarrow \mathrm{PU}(1,1)$ is the representation induced by $\rho, f$ defines a $\rho^{\prime}$-equivariant isometry from $\mathbb{H}_{\mathbb{C}}^{1}$ onto $\mathbb{H}_{\mathbb{C}}^{1}$. In particular, $\rho^{\prime}$ (and so $\rho$ ) is injective, and $\rho^{\prime}\left(\Gamma^{\prime}\right)=\rho(\Gamma)$ must be discrete. Finally, since $f$ descends to an isometry between $\Gamma^{\prime} \backslash \mathbb{H}_{\mathbb{C}}^{1}$ and $\rho(\Gamma) \backslash \mathbb{H}_{\mathbb{C}}^{1}, \rho$ is a uniformization representation and $\rho(\Gamma)$ is a lattice in $\mathrm{PU}(1,1)$.

### 4.2. The general case.

Let now $\Gamma$ be the fundamental group of a $p$-times punctured closed orientable surface of negative Euler characteristic $M$ and $\rho: \Gamma \longrightarrow \mathrm{PU}(n, 1)$ be a reductive representation. In the following, we shall prove Theorem 4.3 in this general setting.

As usual, we identify $\Gamma$ with a non-uniform torsion-free lattice in $\operatorname{PU}(1,1)$ and $M$ with $\Gamma \backslash \mathbb{H}_{\mathbb{C}}^{1}$, and we call $C_{1}, \ldots, C_{p}$ the cusps of $M$. We can assume that $\rho$ is not tame, namely that for some $1 \leq k \leq p, \rho$ maps the peripheral elements corresponding to the punctures $m_{1}, \ldots, m_{k}$ to hyperbolic isometries of $\mathrm{PU}(n, 1)$.

We begin by showing that the proof given in the previous section, until Claim 4.8 included, still holds here.

If one looks carefully at the preceding arguments, one sees that what is really needed is the existence of a $\rho$-equivariant harmonic map $f: \mathbb{H}_{\mathbb{C}}^{1} \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$ such that:
(a) the norm $\|\nabla \mathrm{d} f\|$ is square integrable and hence $f^{\star} \omega_{n}$ is an $L^{2}$ form;
(b) we have $\tau(\rho)=\frac{1}{2} \int_{M}\left\langle f^{\star} \omega_{n}, \omega_{1}\right\rangle d V_{1}$, which together with (a) implies as before the inequality $|\tau(\rho)| \leq \operatorname{Vol}(M)$;
(c) Claim 4.7 holds and hence the equality $|\tau(\rho)|=\operatorname{Vol}(M)$ implies as for tame representations that $f$ is a $\rho$-equivariant immersion from $\mathbb{H}_{\mathbb{C}}^{1}$ into a totally geodesic copy of $\mathbb{H}_{\mathbb{C}}^{1}$ in $\mathbb{H}_{\mathbb{C}}^{n}$.
According to Proposition 4.4 there is no finite energy $\rho$-equivariant map $\mathbb{H}_{\mathbb{C}}^{1} \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$, so that we can not apply Theorem 1.1. Nevertheless, we shall prove that:
(i) there exists a (infinite energy) $\rho$-equivariant harmonic map $f: \mathbb{H}_{\mathbb{C}}^{1} \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$;
(ii) we have a control on the energy density of $f$ at infinity;
(iii) this control implies (a), (b) and (c).

We shall use the notations of the proofs of the tame case.
(i) Our method is the same as in JZ97. We first fix a particular $\rho$-equivariant map $\Phi$ : $\mathbb{H}_{\mathbb{C}}^{1} \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$. As usual, we only define it in each cusp $C_{i}$ of $M$ and extend it on the compact part of $M$. We call $\gamma_{i}$ a generator of the fundamental group of $C_{i}$ (when working on a single cusp, we shall drop the subscript $i$ ). The notations are as in the proof of Proposition 4.4. Let $C$ be a cusp of $M$. If $\rho(\gamma)$ is parabolic or elliptic, then we define $\Phi$ as we defined $f$ in Proposition 4.4 . If $\rho(\gamma)$ is hyperbolic, let $\varphi:[0, a] \longrightarrow \mathbb{H}_{\mathbb{C}}^{n}$ be a map sending $[0, a]$ proportionally to arclength into the axis of $\rho(\gamma)$ (which is a geodesic in $\mathbb{H}_{\mathbb{C}}^{n}$ ), and such that $\varphi(a)=\rho(\gamma) \varphi(0)$. Then, define $\Phi$ by $\Phi(v, t)=\varphi(v)$ (we have $\|\mathrm{d} \Phi\|_{(v, t)}=e^{t} \frac{\delta}{a}$ where $\delta$ is the translation length of $\rho(\gamma)$ ).

For each $s \in \mathbb{R}_{+}$, let $M_{s}$ be the compact Riemann surface with boundary, obtained from $M$ by deleting the end $\{t>s\}$ in each cusp of $M$, and let $F_{s}$ be the restriction of the fiber bundle $F:=\mathbb{H}_{\mathbb{C}}^{1} \times{ }_{\rho} \mathbb{H}_{\mathbb{C}}^{n}$ to $M_{s}$. From Co92, we get the existence of a harmonic section $f_{s}$ of $F_{s}$ which agrees with $\Phi$ on $\partial M_{s}$ and whose energy satisfies $E\left(f_{s}\right) \leq E\left(\Phi_{\mid M_{s}}\right)$.

We want to prove that there exists a strictly increasing sequence $\left(s_{n}\right)_{n \in \mathbb{N}}$ going to infinity, such that $f_{s_{n}}$ converges uniformly on every compact subset of $M$. The limiting function $f$ will be the (infinite energy) harmonic section of $F$ we are looking for. Using the same argument as in Co92, it is sufficient to prove that for each $0 \leq t \leq s$, the energy $E\left(f_{s \mid M_{t}}\right)$ is bounded independently of $s$.

Let $C$ be a cusp of $M$ such that $\rho(\gamma)$ is hyperbolic. If we set $C_{s}=C \cap M_{s}$, from the proof of Proposition 4.4 it is immediate that

$$
E\left(f_{s \mid C_{s} \backslash C_{t}}\right) \geq E\left(\Phi_{\mid C_{s} \backslash C_{t}}\right)=\frac{\delta^{2}}{2 a}\left(e^{s}-e^{t}\right) .
$$

Moreover, denoting by $M^{\prime}$ the union of the compact part of $M$ with the cusps $C_{k+1}, \ldots, C_{p}$, and letting $M^{\prime \prime}=M \backslash M^{\prime}\left(M_{s}^{\prime}=M^{\prime} \cap M_{s}, M_{s}^{\prime \prime}=M^{\prime \prime} \cap M_{s}\right)$, we have

$$
E\left(\Phi_{\mid M_{s}^{\prime}}\right) \leq E\left(\Phi_{\mid M^{\prime}}\right)<+\infty
$$

and so, because of the energy minimizing property of $f_{s}$,

$$
E\left(f_{s \mid M_{t}}\right) \leq E\left(f_{s}\right)-E\left(\Phi_{\mid M_{s}^{\prime \prime} \backslash M_{t}^{\prime \prime}}\right) \leq E\left(\Phi_{\mid M^{\prime} \cup M_{t}^{\prime \prime}}\right)
$$

which is independent of $s$.
(ii) For each $i \in\{1, \ldots, k\}$, we now define

$$
\begin{aligned}
\alpha_{i}:[0,+\infty) & \longrightarrow \mathbb{R} \\
t & \longmapsto \frac{1}{2} \int_{0}^{a_{i}}\left(\|\mathrm{~d} f\|_{(v, t)}^{2}-\|\mathrm{d} \Phi\|_{(v, t)}^{2}\right) e^{-t} d v,
\end{aligned}
$$

where $(v, t) \in\left[0, a_{i}\right] \times[0,+\infty)$ are the usual horospherical coordinates in the cusp $C_{i}$. The $\alpha_{i}$ 's are non-negative functions (see the proof of Proposition 4.4).

We claim that the energy density $e(f)$ is controlled by $e(\Phi)$, namely:

$$
\int_{0}^{+\infty} \alpha_{i}(t) d t<+\infty \text { for any } i \in\{1, \ldots, k\}, \text { and } E\left(f_{\mid M^{\prime}}\right)<+\infty
$$

Otherwise, there exists $r>0$ such that

$$
E\left(f_{\mid M_{r}}\right) \geq 2+E\left(\Phi_{\mid M^{\prime} \cup M_{r}^{\prime \prime}}\right)<+\infty .
$$

We may choose $s>r$ such that

$$
\left|e\left(f_{s}\right)-e(f)\right| \leq \frac{1}{\operatorname{Vol}\left(M_{r}\right)}
$$

on $M_{r}$ and so,

$$
E\left(f_{s \mid M_{r}}\right) \geq 1+E\left(\Phi_{\mid M^{\prime} \cup M_{r}^{\prime \prime}}\right) .
$$

Now, we have

$$
\begin{aligned}
E\left(f_{s}\right)=\int_{M_{s}} e\left(f_{s}\right) d V_{1} & \geq 1+E\left(\Phi_{\mid M^{\prime} \cup M_{r}^{\prime \prime}}\right)+E\left(\Phi_{\mid M_{s}^{\prime \prime} \backslash M_{r}^{\prime \prime}}\right) \\
& \geq 1+E\left(\Phi_{\mid M_{s}}\right)
\end{aligned}
$$

But $f_{s}$ and $\Phi$ coincide on $\partial M_{s}$ and this contradicts the energy minimizing property of the harmonic map $f_{s}$.
(iii) Proof of (a). First, we remark that the map $\Phi$ used to construct $f$ is a rank one harmonic map on $M^{\prime \prime}$, and in fact, it is totally geodesic. From the formula of Eells-Sampson (see Lemma 2.3), we thus get $\Delta e(\Phi)=2 e(\Phi)$ on $M^{\prime \prime}$. Let now be $e_{\Phi}: M \longrightarrow \mathbb{R}$ be a $\left(C^{2}\right)$ function such that $e_{\Phi \mid M^{\prime}}=0$ and $e_{\Phi \mid M^{\prime \prime} \backslash M_{1}^{\prime \prime}}=e(\Phi)_{\mid M^{\prime \prime} \backslash M_{1}^{\prime \prime}}$. If $K$ is the compact subset $\overline{M_{1}^{\prime \prime} \backslash M_{0}^{\prime \prime}}$, we have

$$
\Delta\left(e(f)-e_{\Phi}\right)=-\|\nabla \mathrm{d} f\|^{2}+\operatorname{Scal}\left(f^{\star} R^{n}\right)+2\left(e(f)-e_{\Phi}\right)
$$

on $M \backslash K$. Denoting by $\eta_{R}(R>1)$ the usual cut-off functions, and using Green's formula, we obtain

$$
\begin{aligned}
\int_{M} \eta_{R} \Delta\left(e(f)-e_{\Phi}\right) d V_{1} & =\int_{M}\left(e(f)-e_{\Phi}\right)\left(\Delta \eta_{R}\right) d V_{1} \\
& =\int_{M^{\prime}} e(f)\left(\Delta \eta_{R}\right) d V_{1}+\sum_{i=1}^{k} \int_{0}^{+\infty}\left(\Delta \eta_{R}\right) \alpha_{i}(t) d t+A
\end{aligned}
$$

because $\eta_{R}$ only depends on $t$ in the cusps (here $A$ is a constant independent of $R$ ). Like in Lemma 2.3. the inequality $\|\nabla \mathrm{d} f\|^{2} \leq-\Delta\left(e(f)-e_{\Phi}\right)+2\left(e(f)-e_{\Phi}\right)$, which is valid on $M \backslash K$, now implies that $\|\nabla \mathrm{d} f\| \in L^{2}(M)$ and, like in Lemma 3.4, we also obtain that $f^{\star} \omega_{n}$ is $L^{2}$.

Proof of (b). Since $f^{\star} \omega_{n}$ is $L^{2}$, we see from the proof of Proposition 4.5 that we only need to show that each $f^{\star} \varsigma_{i}$ is an $L^{2}$ form to get $\tau(\rho)=\frac{1}{2} \int_{M}\left\langle f^{\star} \omega_{n}, \omega_{1}\right\rangle d V_{1}$.

Let $i \in\{1, \ldots, k\}$. We take horospherical coordinates $\left(z^{\prime}, v^{\prime}, t^{\prime}\right) \in \mathbb{H}_{\mathbb{C}}^{n}$ such that the axis of $\rho\left(\gamma_{i}\right)$ is given by $z^{\prime}=v^{\prime}=0$. Then $\Phi^{\star} \varsigma_{i}=\Phi^{\star} d z^{\prime}=0$, that is $\|\mathrm{d} \Phi\|=\left\|\Phi^{\star} d t^{\prime}\right\|$. Since $\int_{0}^{a_{i}}\left(\left\|f^{\star} d t^{\prime}\right\|_{(v, t)}^{2}-\left\|\Phi^{\star} d t^{\prime}\right\|_{(v, t)}^{2}\right) d v \geq 0$ for each $t$, the fact that $\|\mathrm{d} f\|^{2} \geq\left\|f^{\star} d t^{\prime}\right\|^{2}+\left\|f^{\star} \varsigma_{i}\right\|^{2}$ together with (ii) implies that $f^{\star} \varsigma_{i}$ is $L^{2}$.

The proof of the inequality $|\tau(\rho)| \leq \operatorname{Vol}(M)$ is like in section 3.2, the key point being that $\|\nabla \mathrm{d} f\| \in L^{2}(M)$.

Proof of (c). We only have to show that for each $\varepsilon>0$,

$$
\lim _{r \rightarrow+\infty} \frac{\operatorname{Area}_{\phi^{\star} h_{\varepsilon}}\left(B_{\phi^{\star} h_{\varepsilon}}(r)\right)}{r^{2}}=0
$$

In a cusp $C_{i}$ for which $\rho\left(\gamma_{i}\right)$ is hyperbolic, we first compare the volume element $d V_{\phi^{\star} h_{\varepsilon}}$ to $d V_{\varepsilon g_{1}+\Phi^{\star} g_{n}}=\sqrt{\varepsilon^{2}+2 \varepsilon e(\Phi)} d V_{1}$ induced by the model map $\Phi$. For any $s>0$, we immediately get

$$
\operatorname{Area}_{\phi^{\star} h_{\varepsilon}}\left(C_{i, s}\right) \leq \operatorname{Area}_{\varepsilon g_{1}+\Phi^{\star} g_{n}}\left(C_{i, s}\right)+2 \int_{0}^{s} \alpha_{i}(t) d t+\frac{1}{2 \varepsilon} \int_{C_{i, s}}\left\|f^{\star} \omega_{n}\right\|^{2} d V_{1} .
$$

Recall that $f^{\star} \omega_{n}$ is $L^{2}$ and, since $E\left(f_{\mid M^{\prime}}\right)<+\infty$, the other cusps have finite area w.r.t. $\phi^{\star} h_{\varepsilon}$. So

$$
\begin{aligned}
\lim _{r \rightarrow+\infty} \frac{\operatorname{Area}_{\phi^{\star} h_{\varepsilon}}\left(B_{\phi^{\star} h_{\varepsilon}}(r)\right)}{r^{2}} & \leq \lim _{r \rightarrow+\infty} \frac{\operatorname{Area}_{\varepsilon g_{1}+\Phi^{\star} g_{n}}\left(B_{\phi^{\star} h_{\varepsilon}}(r) \cap M^{\prime \prime}\right)}{r^{2}} \\
& \leq \lim _{r \rightarrow+\infty} \frac{\operatorname{Area}_{\varepsilon g_{1}+\Phi^{\star} g_{n}}\left(B_{g_{1}}\left(\frac{r}{\sqrt{\varepsilon}}\right) \cap M^{\prime \prime}\right)}{r^{2}}
\end{aligned}
$$

where $B_{g_{1}}(r)$ is the geodesic ball of radius $r$ with respect to the metric $g_{1}$.
But $B_{g_{1}}\left(\frac{r}{\sqrt{\varepsilon}}\right) \cap C_{i} \subset C_{i, \frac{r}{\sqrt{\varepsilon}}}$, and $d V_{\varepsilon g_{1}+\Phi^{\star} g_{n}}=\sqrt{\varepsilon^{2}+\varepsilon \frac{\delta_{i}^{2}}{a_{i}^{2}} e^{2 t}} e^{-t} d v d t$ on $C_{i}$. From this, we conclude that there exists a positive constant $A_{\varepsilon}$ such that

$$
\operatorname{Area}_{\varepsilon g_{1}+\Phi^{\star} g_{n}}\left(B_{g_{1}}\left(\frac{r}{\sqrt{\varepsilon}}\right) \cap M^{\prime \prime}\right) \leq A_{\varepsilon} r
$$

Thus the area of balls for the metric $\phi^{\star} h_{\varepsilon}$ grows at most linearly and the result follows.
As in the tame case, we therefore know that $\rho(\Gamma)$ stabilizes a totally geodesic copy of $\mathbb{H}_{\mathbb{C}}^{1}$ in $\mathbb{H}_{\mathbb{C}}^{n}$, and hence we can consider $\rho$ as a homomorphism from $\Gamma$ to $\operatorname{PU}(1,1)$. Moreover, the harmonic map $f: \mathbb{H}_{\mathbb{C}}^{1} \longrightarrow \mathbb{H}_{\mathbb{C}}^{1}$ is a local diffeomorphism and therefore we can pull-back the complex structure of the target to $M$. The uniformization theorem implies that there exists a complete hyperbolic metric $g_{1}^{\prime}$ on $M$ compatible with this new complex structure and a discrete torsion-free subgroup $\Gamma^{\prime}=u(\Gamma)$ of $\operatorname{PU}(1,1)$, isomorphic to $\Gamma$, such that $\left(M, g_{1}^{\prime}\right)$ is isometric to $\Gamma^{\prime} \backslash \mathbb{H}_{\mathbb{C}}^{1}$.

However, contrary to the case of tame representations, Claim 4.9 does not a priori hold and we will assume that $\Gamma^{\prime}$ is not a lattice, namely that the volume of $M$ with respect to $g_{1}^{\prime}$ is infinite (in fact, Claim 4.9 and the proof of Claim 4.10 below imply that $\Gamma^{\prime}$ is a lattice if and only if $\rho$ is tame). This means that some punctures, say $m_{1}, \ldots, m_{q}$, have neighbourhoods of infinite $g_{1}^{\prime}$-volume. More precisely, peripheral elements of $\Gamma$ corresponding to these punctures are sent by the new uniformization representation $u$ to hyperbolic isometries of $\operatorname{PU}(1,1)$ and for each $i \in\{1, \ldots, q\}, m_{i}$ has a neighbourhood $g_{1}^{\prime}$-isometric to the annulus

$$
A_{i}:=\left\{z \in \mathbb{C}, e^{-\pi / b_{i}}<|z|<a_{i}\right\}, b_{i}>0,1>a_{i}>e^{-\pi / 2 b_{i}},
$$

endowed with the metric

$$
\left(\frac{2 b_{i}}{|z| \sin \left(b_{i} \log |z|\right)}\right)^{2} d z d \bar{z}
$$

In particular, there exist disjoint simple closed $g_{1}^{\prime}$-geodesics $c_{1}, \ldots, c_{q}$ in $M$, corresponding to the circles $\left\{|z|=e^{-\pi / 2 b_{i}}\right\}$ in $A_{i}$, such that $c_{i}$ is null-homotopic in $M \cup\left\{m_{i}\right\}$ for each $i \in\{1, \ldots, q\}$.

If we cut $M$ along these geodesics and double the remaining finite volume part $M_{0}$ (whose interior is diffeomorphic to $M$ ) along its (possibly disconnected) $g_{1}^{\prime}$-geodesic boundary, we obtain a surface $2 M$, on which $g_{1}^{\prime}$ extends by symmetry to a complete hyperbolic metric of finite volume. We call $\sigma$ the $g_{1}^{\prime}$-isometric (antiholomorphic) involution of $2 M$. Note that the Euler characteristic $\chi(2 M)$ of $2 M$ equals $2 \chi(M)$.

We want to extend the representation $\rho: \Gamma \longrightarrow \mathrm{PU}(1,1)$ to a representation $2 \rho$ of the fundamental group $2 \Gamma$ of $2 M$. For this we need an adapted presentation of $2 \Gamma$.

The hyperbolic metric $g_{1}^{\prime}$ on $2 M$ allows us to identify $2 \Gamma$ with a lattice in $\operatorname{PU}(1,1)$ and $2 M$ with $2 \Gamma \backslash \mathbb{H}_{\mathbb{C}}^{1}$. Call again $\pi: \mathbb{H}_{\mathbb{C}}^{1} \longrightarrow 2 M$ the covering projection and choose a connected component $X^{\prime}$ of $\mathbb{H}_{\mathbb{C}}^{1} \backslash \bigcup_{i=1}^{q} \pi^{-1}\left(c_{i}\right)$. Since $\bigcup_{i=1}^{q} \pi^{-1}\left(c_{i}\right)$ consists of disjoint geodesics of $\mathbb{H}_{\mathbb{C}}^{1}$, $X^{\prime}$ is a convex set whose stabilizer in $2 \Gamma$ can be identified with $\Gamma^{\prime}$. Moreover, $\pi\left(X^{\prime}\right)=\Gamma^{\prime} \backslash X^{\prime}$
is just the interior of $M_{0}$, seen as a subset of $2 M$. For each $i$ choose a lift $\tilde{c}_{i}$ of $c_{i}$ in the boundary of $X^{\prime}$, a generator $\gamma_{i}^{\prime}$ of the cyclic stabilizer of $\tilde{c}_{i}$ in $\Gamma^{\prime}$, and call $\sigma_{i}$ the symmetry w.r.t. $\tilde{c}_{i}$. It is easy to see that the symmetries $\sigma_{i}$ are lifts of $\sigma$ and therefore for all $i$ there exists $h_{i} \in 2 \Gamma$ such that $\sigma_{i}=h_{i} \sigma_{1}$.

Let $X^{\prime \prime}=\sigma_{1} X^{\prime}$ be the other component of $\mathbb{H}_{\mathbb{C}}^{1} \backslash \bigcup_{i=1}^{q} \pi^{-1}\left(c_{i}\right)$ adjacent to $\tilde{c}_{1}$, and call $\Gamma^{\prime \prime}$ $\left(=\sigma_{1} \Gamma^{\prime} \sigma_{1}\right)$ its stabilizer in $2 \Gamma$. Note that, for all $i, h_{i}^{-1} \gamma_{i}^{\prime} h_{i}=\sigma_{1} \sigma_{i} \gamma_{i}^{\prime} \sigma_{i} \sigma_{1}=\sigma_{1} \gamma_{i}^{\prime} \sigma_{1}$ belongs to $\Gamma^{\prime \prime}$.

The fundamental group $2 \Gamma$ of $2 M$ has the following abstract presentation (see for example [Se77]). It is generated by the sets $\Gamma^{\prime}$ and $\Gamma^{\prime \prime}$, together with the $h_{i}$ 's, subject to the relations:

- the relations of the groups $\Gamma^{\prime}$ and $\Gamma^{\prime \prime}$;
- for all $i \in\{1, \ldots, q\}$, the element $\gamma_{i}^{\prime}$ of $\Gamma^{\prime}$ is identified with the element $\gamma_{i}^{\prime \prime}:=h_{i}^{-1} \gamma_{i}^{\prime} h_{i}$ of $\Gamma^{\prime \prime}$.

Now, it follows from the discussion in the proof of Claim 4.9 that for all $i \in\{1, \ldots, q\}$, $g_{i}:=\rho \circ u^{-1}\left(\gamma_{i}^{\prime}\right)$ is an hyperbolic isometry of $\mathbb{H}_{\mathbb{C}}^{1}$. Call $s_{i}$ the symmetry w.r.t. the axis of $g_{i}$ and set

$$
\left\{\begin{array}{l}
2 \rho\left(\gamma^{\prime}\right)=\rho \circ u^{-1}\left(\gamma^{\prime}\right) \text { for } \gamma^{\prime} \in \Gamma^{\prime} ; \\
2 \rho\left(\gamma^{\prime \prime}\right)=s_{1}\left(\rho \circ u^{-1}\left(\sigma_{1} \gamma^{\prime \prime} \sigma_{1}\right)\right) s_{1} \text { for } \gamma^{\prime \prime} \in \Gamma^{\prime \prime} ; \\
2 \rho\left(h_{i}\right)=s_{i} s_{1} \text { for } i \in\{1, \ldots, q\} .
\end{array}\right.
$$

This is clearly compatible with the relations of $2 \Gamma$ and hence $2 \rho$ is a well-defined representation of $2 \Gamma$ into $\mathrm{PU}(1,1)$.

Claim 4.10. The representation $2 \rho: 2 \Gamma \longrightarrow \mathrm{PU}(1,1)$ is tame.
Proof. It is enough to show that, for any $i \in\{q+1, \ldots, p\}$, the peripheral elements $\gamma_{i}$ corresponding to the punctures $m_{i}$ are not mapped by $\rho$ to hyperbolic isometries of $\mathbb{H}_{\mathbb{C}}^{1}$.

Suppose that one of them is. We drop the subscript $i$. The quotient $\langle u(\gamma)\rangle \backslash \mathbb{H}_{\mathbb{C}}^{1}$ is holomorphically equivalent to a punctured disc $D^{\star}=\{z \in \mathbb{C}, 0<|z|<1\}$, whereas the quotient $\langle\rho(\gamma)\rangle \backslash \mathbb{H}_{\mathbb{C}}^{1}$ is holomorphically equivalent to an annulus $A=\{z \in \mathbb{C}, a<|z|<1\}$. The $\rho$-equivariant holomorphic map $f: \mathbb{H}_{\mathbb{C}}^{1} \longrightarrow \mathbb{H}_{\mathbb{C}}^{1}$ induces a holomorphic map $f$ from $D^{\star}$ to $A$. This map is bounded and hence extends to a holomorphic map from the whole disc $D$ to $\bar{A}$. This is impossible since the $\rho$-equivariance of $f$ implies that for all $0<r<1$, the loops $f\left(\left\{z \in D^{\star},|z|=r\right\}\right)$ are homotopically non trivial.

Claim 4.11. The invariant $\tau(2 \rho)$ is maximal: $\tau(2 \rho)=-2 \pi \chi(2 M)$.
Proof. Since $2 \rho$ is tame, there exists a finite energy $2 \rho$-equivariant map $f: \mathbb{H}_{\mathbb{C}}^{1} \longrightarrow \mathbb{H}_{\mathbb{C}}^{1}$ and $\tau(2 \rho)=\int_{2 M} f^{\star} \omega_{1}$.

For each $i \in\{1, \ldots, q\}$, we may choose a symmetric tubular neighbourhood $N_{i}$ of the geodesic $c_{i}$ and a lift $\widetilde{N}_{i}$ in $\mathbb{H}_{\mathbb{C}}^{1}$ such that $N_{i}=\left\langle\gamma_{i}^{\prime}\right\rangle \backslash \widetilde{N}_{i}$. Let $\psi_{i}$ be a Kähler potential associated as before to a fixed point $\xi_{i}$ of $2 \rho\left(\gamma_{i}^{\prime}\right)$ and let $\varsigma_{i}$ be the corresponding 1 -form. Then $f^{\star} \varsigma_{i}$ is a 1 -form on $N_{i}$ such that $f^{\star} \omega_{1}=\mathrm{d} f^{\star} \varsigma_{i}$. If now $\eta$ is a symmetric cut-off function on $2 M$ identically equal to 0 outside the $N_{i}$ 's and to 1 close to the geodesics $c_{i}$, we get:

$$
\begin{aligned}
\tau(2 \rho) & =\int_{2 M}\left[f^{\star} \omega_{1}-\mathrm{d}\left(\sum_{i=1}^{q} \eta f^{\star} \varsigma_{i}\right)\right] \\
& =\int_{M_{0}}\left[f^{\star} \omega_{1}-\mathrm{d}\left(\sum_{i=1}^{q} \eta f^{\star} \varsigma_{i}\right)\right]+\int_{\sigma M_{0}}\left[f^{\star} \omega_{1}-\mathrm{d}\left(\sum_{i=1}^{q} \eta f^{\star} \varsigma_{i}\right)\right] \\
& =\int_{M_{0}}\left[f^{\star} \omega_{1}-\mathrm{d}\left(\sum_{i=1}^{q} \eta f^{\star} \varsigma_{i}\right)\right]-\int_{M_{0}}\left[\sigma^{\star} f^{\star} \omega_{1}-\mathrm{d}\left(\sum_{i=1}^{q} \eta \sigma^{\star} f^{\star} \varsigma_{i}\right)\right] .
\end{aligned}
$$

Extending the form $f^{\star} \omega_{1}-\mathrm{d}\left(\sum_{i=1}^{q} \eta f^{\star} \varsigma_{i}\right)$ on $M_{0}$ by zero to a form on $M$, one sees that $\int_{M}\left[f^{\star} \omega_{1}-\mathrm{d}\left(\sum_{i=1}^{q} \eta f^{\star} \varsigma_{i}\right)\right]=\tau(\rho)$.

The forms $\sigma^{\star} f^{\star} \omega_{1}$ on $M$ and $\sigma^{\star} f^{\star} \varsigma_{i}$ on $N_{i}$ are respectively induced by $\sigma_{1}^{\star} f^{\star} \omega_{1}$ on $\mathbb{H}_{\mathbb{C}}^{1}$ and $\sigma_{1}^{\star} f^{\star} s_{i}$ on $\sigma_{1} \widetilde{N}_{i}$. Since $s_{1}^{\star} \omega_{1}=-\omega_{1}$, one has $\sigma_{1}^{\star} f^{\star} \omega_{1}=-\left(s_{1} \circ f \circ \sigma_{1}\right)^{\star} \omega_{1}$. Moreover, $\varsigma_{i}=s_{1}^{\star} s_{1}^{\star}\left(-\mathrm{d}^{c} \psi_{i}\right)=s_{1}^{\star} \mathrm{d}^{c}\left(\psi_{i} \circ s_{1}\right)$, and $\psi_{i} \circ s_{1}=\psi_{i} \circ s_{i} \circ s_{1}=2 \rho\left(h_{i}\right)^{\star} \psi_{i}$. Hence $\sigma_{1}^{\star} f^{\star} \varsigma_{i}=$ $-\left(s_{1} \circ f \circ \sigma_{1}\right)^{\star} 2 \rho\left(h_{i}\right)^{\star} \varsigma_{i}$. Since $s_{1} \circ f \circ \sigma_{1}$ is also $\rho$-equivariant, the 1-form $\left(s_{1} \circ f \circ \sigma_{1}\right)^{\star} 2 \rho\left(h_{i}\right)^{\star} \varsigma_{i}=$ $h_{i}^{\star}\left(s_{1} \circ f \circ \sigma_{1}\right)^{\star} \varsigma_{i}$ on $h_{i}^{-1} \widetilde{N}_{i}=\sigma_{1} \widetilde{N}_{i}$ induces $\left(s_{1} \circ f \circ \sigma_{1}\right)^{\star} \varsigma_{i}$ on $N_{i}$. Finally,

$$
-\int_{M_{0}}\left[\sigma^{\star} f^{\star} \omega_{1}-\mathrm{d}\left(\sum_{i=1}^{q} \eta \sigma^{\star} f^{\star} \varsigma_{i}\right)\right]=\int_{M_{0}}\left[\left(s_{1} \circ f \circ \sigma_{1}\right)^{\star} \omega_{1}-\mathrm{d}\left(\sum_{i=1}^{q} \eta\left(s_{1} \circ f \circ \sigma_{1}\right)^{\star} \varsigma_{i}\right)\right]
$$

and the r.h.s. again equals $\tau(\rho)$ since $s_{1} \circ f \circ \sigma_{1}$ is a $\rho$-equivariant map. Hence $\tau(2 \rho)=2 \tau(\rho)=$ $-4 \pi \chi(M)=-2 \pi \chi(2 M)$ and the lemma is proved.

The representation $2 \rho: 2 \Gamma \longrightarrow \mathrm{PU}(1,1)$ is therefore a tame representation of maximal invariant. It follows from the results of section 4.1 that there exists a $2 \rho$-equivariant isometry $f$ from $\mathbb{H}_{\mathbb{C}}^{1}$ onto $\mathbb{H}_{\mathbb{C}}^{1}$. Since $f$ is a fortiori $\left(\rho \circ u^{-1}\right)$-equivariant, $\rho$ is a uniformization representation and we are done.

## References

[Au60] L. Auslander, Bieberbach's theorems on space groups and discrete uniform subgroups of Lie groups, Ann. of Math. 71, 1960, 579-5906
[Bi97] O. Biquard, Métriques d'Einstein à cusps et équations de Seiberg-Witten, J. Reine Angew. Math. 490, 1997, 129-1545
[BI01] M. Burger and A. Iozzi, Bounded cohomology and representation varieties of lattices in PSU $(1, n)$, Preprint, 2001 2, 17,18
[BI03] M. Burger and A. Iozzi, Letter, 20034
[BM02] M. Burger and N. Monod, Continuous bounded cohomology and applications to rigidity theory, Geom. Funct. Anal. 12, 2002, 219-280 2
[CT89] J. A. Carlson and D. Toledo, Harmonic mappings of Kähler manifolds to locally symmetric spaces, Publ. Math. I.H.E.S. 69, 1989, 173-201 14,15
[Co88] K. Corlette, Flat G-bundles with canonical metrics, J. Diff. Geom. 28, 1988, 361-382 2 , 18
[Co92] K. Corlette, Archimedean superrigidity and hyperbolic geometry, Ann. of Math. 135, 1992, 164-182 13 ( 4 6 1131
[ES64] J. Eells and J. H. Sampson, Harmonic mappings of Riemannian manifolds, Amer. J. Math. 86, 1964, 109-160 2,12
[Ga54] M. P. Gaffney, A special Stokes's theorem for complete Riemannian manifolds, Ann. of Math. 60, 1954, 140-145 25
[Go85] W. M. Goldman, Representations of fundamental groups of surfaces, Geometry and topology (College Park, Md., 1983/84), Lecture Notes in Math. 1167, 95-117, Springer, Berlin, 19851
[Go99] W. M. Goldman, Complex hyperbolic geometry, The Clarendon Press, Oxford University Press, New York, 19995
[GM87] W. M. Goldman and J. J. Millson, Local rigidity of discrete groups acting on complex hyperbolic space, Invent. Math. 88, 1987, 495-520 1
[GS92] M. Gromov and R. Schoen, Harmonic maps into singular spaces and p-adic superrigidity for lattices in groups of rank one, Publ. Math. I.H.E.S. 76, 1992, 165-246 1
[GP00] N. Gusevskii and J. R. Parker, Representations of free Fuchsian groups in complex hyperbolic space, Topology 39, 2000, 33-60 321
[GP03] N. Gusevskii and J. R. Parker, Complex hyperbolic quasi-Fuchsian groups and Toledo's invariant, Geom. Dedicata 97, 2003, 151-185 3
[Hel84] S. Helgason, Groups and geometric analysis, Pure and Applied Math. 113, Academic Press, NewYork, 198412
[Her91] L. Hernandez, Kähler manifolds and $\frac{1}{4}$-pinching, Duke Math. J. 62, 1991, 601-611 14
[HS96] C. Hummel and V. Schroeder, Cusp closing in rank one symmetric spaces, Invent. Math. 123, 1996, 283-307 5
[II002] A. Iozzi, Bounded cohomology, boundary maps and rigidity of representations into Homeo $+\left(\mathrm{S}^{1}\right)$ and $\mathrm{SU}(1, n)$, Rigidity in dynamics and geometry (Cambridge, 2000), Springer, Berlin, 2002, 237-260 2
[JM87] D. Johnson and J. J. Millson, Deformation spaces associated to compact hyperbolic manifolds, Discrete groups in geometry and analysis (New Haven, Conn., 1984), Progr. Math. 67, Birkhäuser, Boston, MA, 1987, 48-106 1
[JZ97] J. Jost and K. Zuo, Harmonic maps of infinite energy and rigidity results for representations of fundamental groups of quasiprojective varieties, J. Diff. Geom. 47, 1997, 469-503 331
[Ka98] M. Kapovich, On normal subgroups in the fundamental groups of complex surfaces, arXiv:math.GT/9808085 20
[Ko70] S. Kobayashi, Hyperbolic manifolds and holomorphic mappings, Marcel Dekker, New York, 197018
[Li97] P. Li, Complete surfaces of at most quadratic area growth, Comm. Math. Helv. 72, 1997, 67-71 27
[Lic70] A. Lichnerowicz, Applications harmoniques et variétés kählériennes, Symp. Math III Bologna, 1970, 341-402 16
[Liv81] R. Livné, On certain covers of the universal elliptic curve, Ph.D. Thesis, Harvard University, 198120
[Ma91] G. A. Margulis, Discrete subgroups of semisimple Lie groups, Ergeb. der Math. 17, Springer-Verlag, Berlin, 19911
[MSY93] N. Mok, Y.-T. Siu and S.-K. Yeung, Geometric superrigidity, Invent. Math. 113, 1993, 57-83 3 10
[Mo73] G. D. Mostow, Strong rigidity of locally symmetric spaces, Annals of Math. Studies 78, Princeton Univ. Press, Princeton N.J., 19731
[Pa95] P. Pansu, Sous-groupes discrets des groupes de Lie : rigidité, arithméticité, Séminaire Bourbaki, Vol. 1993/94, Astérisque 227, Société Mathématique de France, Paris, 1995, 69-105 10
[Re93] A. G. Reznikov, Harmonic maps, hyperbolic cohomology and higher Milnor inequalities, Topology 32, 1993, 899-907 27
[Sa78] J. H. Sampson, Some properties and applications of harmonic mappings, Ann. Sci. École Norm. Sup. 11, 1978, 211-228 26, 28
[Se77] J.-P. Serre, Arbres, amalgames, $\mathrm{SL}_{2}$, Astérisque 46, Société Mathématique de France, Paris, 197734
[Si80] Y.-T. Siu, The complex-analyticity of harmonic maps and the strong rigidity of compact Kähler manifolds, Ann. of Math. 112, 1980, 73-1112
[SY78] R. Schoen and S.-T. Yau, On univalent harmonic maps between surfaces, Invent. Math. 44, 1978, 265-278 29
[To79] D. Toledo, Harmonic mappings of surfaces to certain Kähler manifolds, Math. Scand. 45, 1979, 13-26 16. 28,29
[To89] D. Toledo, Representations of surface groups in complex hyperbolic space, J. Diff. Geom. 29, 1989, 125-1332, 3, 17,18
[Wo79] J. C. Wood, Holomorphicity of certain harmonic maps from a surface to complex projective $n$-space, J. London Math. Soc. 20, 1979, 137-142 16
[Zu82] S. Zucker, $L_{2}$-cohomology of warped products and arithmetic groups, Invent. Math. 70, 1982, 169-218 17

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